Motives	Introduction and that's why …	What is $f(R)$	Our f(R) type	Next and be

Further analysis with metric-affine f(R) gravity

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August 19, 2024





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Image: August 19, 2024

Motives	Introduction and that's why ···	What is $f(R)$	Our f(R) type	Next and beyond the horizons …	Biblography
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6 Biblography

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Motives	Introduction and that's why …	What is $f(R)$	Our f(R) type	Next and beyond the horizons ···	Biblography
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Backgrou	Ind and Motivation				

Why this topic:

- If we check our main goal, and set back for a while!
- Basically we need some presentations of this category.
- It does really root in even unimaginable levels.
- For a batter understanding of our universe.



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- Studying cosmology and pontificating through it has been our horizon, especially lately.
- As we gain more knowledge, many obstacles come across our trajectory.
- Some of these problems is: the late Universe acceleration.
- Many attmepts has been proposed. E.g., models that requires exotic sources AND; on the other hand, modified gravity via non-Einsteinian dynamics.
- The Hubble issue.



Motives	Introduction and that's why …	What is $f(R) = 0$	Our f(R) type	Next and beyond the horizons …	Biblography
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- GR had a wide range of acceptance then, but it took physicists only 4 year to start questioning its uniqueness!
- Early vs. Late Universe.
- Some modification on the Einstein-Hilbert action.
- It is not that easy task, the competition is hard.
- There are mainly two variational principles to derive Einstein's equations from the Einstein-Hilbert action; Metric variation & Palatini variation.



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Motives Introduction and that's why \cdots What is f(R) Our f(R) type Next and beyond the horizons \cdots Biblography ocoor metric-affine f(R):

• But there is a third one; when the assumption of the Independence of matter action from the connection \rightarrow **metric-affine** f(R) **gravity** which makes it the most general!

$$S_{MA} = \frac{1}{2\kappa} \int d^4x \sqrt{-g} f(\mathcal{R}) + S_M(g_{\mu\nu}, \Gamma^{\lambda}_{\mu\nu}, \psi)$$

- where $\kappa \equiv 8\pi G$, *G* is the gravitational constant, *g* is the determinant of the metric, ψ collectively denotes the matter fields, $\mathcal{R} = g^{\mu\nu} \mathcal{R}_{\mu\nu}$.
- We need **hypermomentum** to mimics the definition of the stress-energy tensor as following

$$\Delta_{\lambda}^{\mu\nu} \equiv -\frac{2}{\sqrt{-g}} \frac{\delta S_M}{\delta \Gamma_{\mu\nu}^{\lambda}}$$

• The way is long from here \cdots



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The Hor	izons:				

- It is the most general case of f(R) for its enriched phenomenology, e.g., matter-induced, non-metricity, and torsion, which comes quite naturally since it is introduced by particles with spin.
- It is not a metric theory, hence the name! $\rightarrow T^{\mu\nu}$ is not divergence-free with respect to the covariant derivative defined with the Levi-Civita connection (nor with $\bar{\nabla}_{\mu}$)
- The physics meaning of the last statement is questionable and **further analysis** *is needed* since in the metric-affine gravity $T_{\mu\nu}$ does not really carry the usual stress-energy tensor, and we already have the hypermomentum which describes matter characteristics.
- Viability, representations, and Post-Newtonian limits



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Thanks a bundle !

Don't hesitate reaching me out for any questions or collaboration

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