

SNOLAB | July 8-19 2024

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TRISEP

$$\frac{dR}{dE_R} = \frac{\rho_0}{m_\chi m_N} \int_{v_{min}}^{\infty} v f(\vec{v}) \frac{d\vec{v}}{dE_R} \cdot a v$$

STANDARD MODEL PHYSICS AND BEYOND
STATISTICS & COMPUTING FOR PARTICLE PHYSICS
DARK MATTER CANDIDATES AND DETECTION

NEUTRINOS PHYSICS

LONG BASELINE NEUTRINO MEASUREMENTS

NEUTRINOLESS DOUBLE BETA DECAY

FOCUSED LECTURES ON UNDERGROUND TOPICS

$$\frac{\Delta m_{ij}^2 L}{4E} \approx 1.267 \frac{\Delta m_{ij}^2 [eV^2] \times L [km]}{E [GeV]}$$



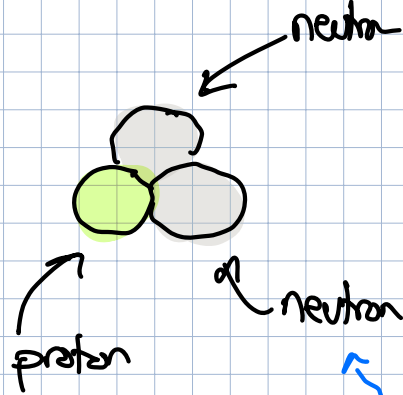
The Standard Model

Problem Solutions

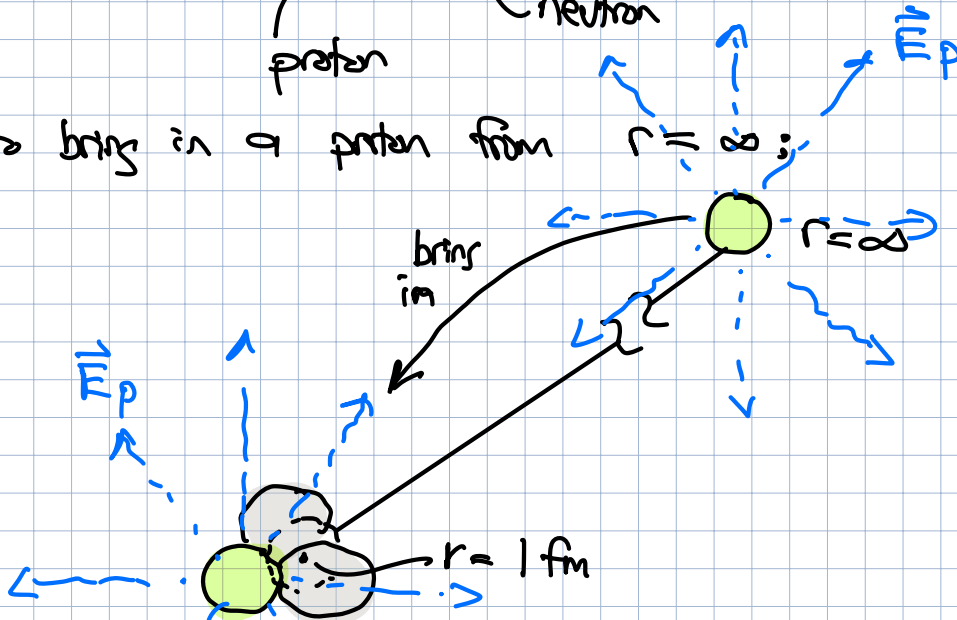
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Problem 1.1

I have a tritium nucleus:



I want to bring in a proton from $r = \infty$:



The work that must be done to achieve this is:

$r_0 = 1 \text{ fm}$ from the centre of the original proton

$$W = \int_{\infty}^{r_0} \vec{F} \cdot d\vec{r} = q \Delta V = q (V_{r_0} - V_{\infty})$$

$V_{\infty} = 0$, and $V_{r_0} = k \frac{q}{r_0}$ where $q = e = 1.602 \times 10^{-19} \text{ C}$

$$\frac{1}{4\pi\epsilon_0} = 8.99 \times 10^9 \frac{\text{Nm}^2}{\text{C}^2}$$

$$W = e^2 k / r_0 = (1.602 \times 10^{-19} \text{ C})^2 (8.99 \times 10^9 \frac{\text{Nm}^2}{\text{C}^2}) / (1 \times 10^{-15} \text{ m}) = 2.307 \times 10^{-13} \text{ J}$$

$$1.602 \times 10^{-19} \text{ J} = 1 \text{ eV} \Rightarrow W = 1.440 \times 10^6 \text{ eV} \approx 1 \text{ MeV}$$

Comment: this is absurdly consistent with the observed energy scale of nuclear phenomena!

BONUS: the wavelength of a photon emitted from a nuclear process would correspond to this ~ 1 MeV energy scale

$$E = hf = 2\pi\hbar f \quad c = \lambda f \Rightarrow E = 2\pi\hbar \frac{c}{\lambda}$$

$$\hbar \approx 1.504 \times 10^{-34} \text{ J}\cdot\text{s}$$

or (from pdg. bl. gw, "Physical Constants")

Planck constant
Planck constant, reduced

h
 $\hbar \equiv h/2\pi$

$6.626\,070\,15 \times 10^{-34} \text{ J}\cdot\text{s}$ (or J/Hz) §
 $1.054\,571\,817 \dots \times 10^{-34} \text{ J}\cdot\text{s}$
 $= 6.582\,119\,569 \dots \times 10^{-22} \text{ MeV}\cdot\text{s}$

exact
exact*
exact*

$$\hbar c \approx 197 \text{ MeV}\cdot\text{fm}$$

$c = 2.998 \times 10^{25} \text{ fm/s}$
 $\text{fm} = 10^{-15} \text{ m}$

$$\hookrightarrow \lambda = \frac{2\pi (197 \text{ MeV}\cdot\text{fm})}{1 \text{ MeV}}$$

$$= 2\pi \cdot 197 \text{ fm} = 1238 \text{ fm} \approx 1 \text{ pm}$$

A picometre wavelength is a gamma ray.

Problem 1.2

Convert the electron rest mass in MKS \rightarrow MeV/c²

using $E^2 = p^2 c^2 + m^2 c^4 \xrightarrow{\text{at rest}} E = mc^2$

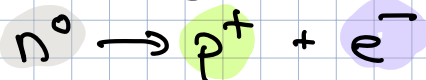
$E = mc^2 = 9.109 \times 10^{-31} \text{ kg} \times (2.998 \times 10^8 \text{ m/s})^2$
 $= 8.187 \times 10^{-14} \text{ J} \xrightarrow{\text{eV}} 511.1 \times 10^3 \text{ eV}$

Mass-energy

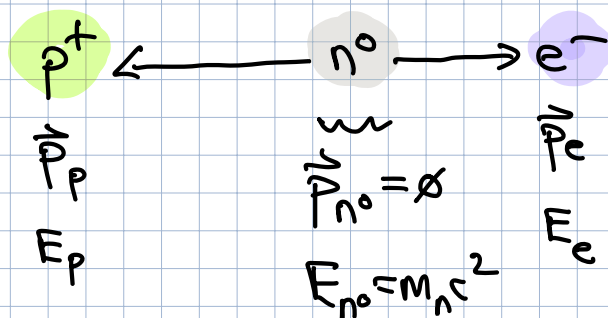
so $m_e = \frac{E}{c^2} = 0.511 \text{ MeV}/c^2$

Problem 1.3

Treat β -decay of free neutron or if the process proceeds by:



treat n^0 as being at rest. then:



Recall that energy + momentum are separately conserved:

$$E_n = E_p + E_e \rightarrow m_n c^2 = E_p + E_e$$

$$\vec{p}_n = \vec{p}_p + \vec{p}_e \rightarrow \emptyset = \vec{p}_p + \vec{p}_e \rightarrow \vec{p}_p = -\vec{p}_e$$

$|\vec{p}_p| = |\vec{p}_e|$

And in total:

$$P_n = P_p + P_e$$

$$\text{where } \begin{cases} P = (E, \vec{p}c) \\ P^2 = E^2 - \vec{p} \cdot \vec{p} c^2 \\ = m^2 c^4 \end{cases} \text{ 4-vector!}$$

$$P_e = P_n - P_p$$

$$P_e^2 = (P_n - P_p)^2$$

$$m_e^2 = P_n^2 - 2P_n P_p + P_p^2 = m_n^2 c^4 - 2E_n E_p + m_p^2 c^4$$

$$\begin{aligned} m_e^2 &= m_n^2 c^4 + m_p^2 c^4 - 2(E_n E_p - \underbrace{\vec{p}_n \cdot \vec{p}_p c^2}_{\text{vanishes, } \vec{p}_n = \vec{p}_p}) \\ &= m_n^2 c^4 + m_p^2 c^4 - 2m_n c^2 E_p \end{aligned}$$

so

$$E_p = \frac{m_n^2 + m_p^2 - m_e^2}{2m_n c^2} c^4$$

Then:

$$\begin{aligned} E_e = E_n - E_p &= m_n c^2 - \frac{m_n^2 + m_p^2 - m_e^2}{2m_n c^2} c^4 \\ &= \frac{2m_n^2 c^4}{2m_n c^2} - \frac{m_n^2 + m_p^2 - m_e^2}{2m_n c^2} c^4 \end{aligned}$$

$$E_e = \frac{m_n^2 - m_p^2 + m_e^2}{2m_n c^2} c^4$$

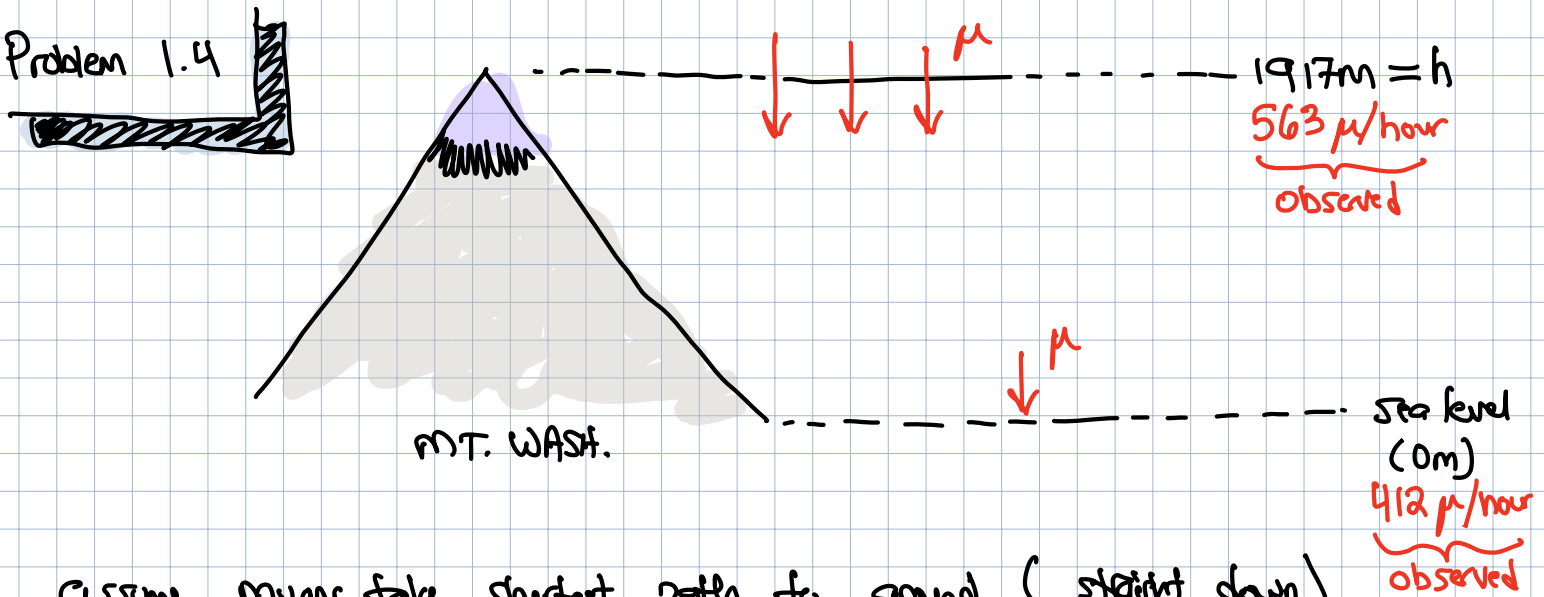
This yields a single value for E_e determined from;

$$\left. \begin{aligned} m_n &= 939.56542052 \text{ MeV}/c^2 \\ m_p &= 938.27208943 \text{ MeV}/c^2 \\ m_e &= 0.51099895069 \text{ MeV}/c^2 \end{aligned} \right\}$$

$$E_e = 1.2926... \text{ MeV}$$

which looks nothing like a real β -decay spectrum.

Problem 1.4



assume muons take shortest path to ground (straight down).
 at h , $N_\mu = 563$ muon (in 1 hour).

Classical calculation

For $\tau = 2.1969 \dots \mu\text{s}$ lifetime, each $\Delta t = \tau$ results
 in a reduction of the survival probability by $1/e = 0.367 \dots$

what is the probability of survival for muons if we ignore
 relativity?

↳ how long does it take muons to cross the 1917m
 gap?

$$\Delta y = h = v_\mu \Delta t \rightarrow \Delta t = \frac{h}{v_\mu}$$

$$\Delta t_{\min} \text{ when } v_\mu = 0.9954c \Rightarrow \frac{h}{v_{\mu \max}} = 6.4238 \times 10^{-6} \text{ s}$$

$$\Delta t_{\max} \text{ when } v_\mu = 0.9950c \Rightarrow \frac{h}{v_{\mu \min}} = 6.4264 \times 10^{-6} \text{ s}$$

$$P_{\text{survival}}^{\min} = e^{-\Delta t_{\max}/\tau} = 0.05366$$

$$P_{\text{survival}}^{\max} = e^{-\Delta t_{\min}/\tau} = 0.05372$$

so classically, $N_{\text{sealers}} = N_h \times \begin{cases} P_{\text{survival}}^{\text{max}} = 30.246 \\ P_{\text{survival}}^{\text{min}} = 30.210 \end{cases}$

$$\langle N_{\text{sealers}} \rangle = 30.228 \pm 0.018 \quad (\text{classical})$$

This is way below the observation.

Relativistic Calculations

We can use either of two approaches:

1) Muan's perspective: it measures proper length, and observes the height h to be contracted by $h' = h/\gamma$
 where $\gamma = (1 - \beta^2)^{-1/2}$

2) earth observer perspective: the muan's clock appears to run slowly, so its lifetime is dilated to $\tau' = \gamma \tau$

What is γ ?

$$\gamma_{\text{min}} = (1 - \beta_{\text{min}}^2)^{-1/2} = 10.013$$

$$\gamma_{\text{max}} = (1 - \beta_{\text{max}}^2)^{-1/2} = 10.438$$

I will use the second approach - it's a little faster. Try the first approach to confirm the results are the same.

$$P_{\text{survival}}^{\text{min}} = e^{-\Delta t_{\text{max}} / \gamma_{\text{min}} \tau} = 0.7467$$

$$P_{\text{survival}}^{\text{max}} = e^{-\Delta t_{\text{min}} / \gamma_{\text{max}} \tau} = 0.7557$$

$$\langle N_{\text{seal}} \rangle = N_{\mu} \frac{P_{\text{max}} + P_{\text{min}}}{2} = 422.9 \pm 2.5 \text{ relativity}$$

The real experiment observed 412 μ /hour. Applying Poisson errors to the count, $412 \pm \sqrt{412} = 412 \pm 20$, we find remarkable agreement:

$$\langle N_{\mu} \rangle_{\text{expected}} = 422.9 \pm 2.5$$

$$N_{\mu}^{\text{observed}} = 412 \pm 20$$

Wow!
Σ

Problem 1.5

See solution in Python in Github. I used `zfit` to model the unbinned data using

$$P(t) = A e^{-\lambda t} + B$$

$\tau = 1/\lambda$

Problem 2.1

We want a simple method that estimates the mass of a hypothetical omega (Ω) baryon.

We are told:

$$m_{\Delta} \sim 1232 \text{ MeV}/c^2 \quad (\Delta \in \Delta^-, \Delta^0, \Delta^+, \Delta^{++})$$

$$m_{\Sigma^*} \sim 1385 \text{ MeV}/c^2 \quad (\Sigma^* \in \Sigma^{*-}, \Sigma^{*0}, \Sigma^{*+})$$

$$m_{\Xi^*} \sim 1533 \text{ MeV}/c^2 \quad (\Xi^* \in \Xi^{*-}, \Xi^{*0})$$

are the "mass splittings" between neighboring states in the incomplete decuplet the same, or similar?

$$m_{\Sigma^*} - m_{\Delta} = 1385 - 1232 = 153 \text{ MeV}/c^2$$

$$m_{\Xi^*} - m_{\Sigma^*} = 1533 - 1385 = 148 \text{ MeV}/c^2$$

$$\langle \Delta m \rangle = \frac{153 + 148}{2} = 150.5 \text{ MeV}/c^2$$

huh. remarkably similar!

so estimate $m_{\Omega} = m_{\Xi^*} + \langle \Delta m \rangle$

$$= 1533 + 151 = 1684 \text{ MeV}/c^2$$

The real observed mass is $m_{\Omega}^{\text{obs}} = 1672 \text{ MeV}/c^2$.

This was a remarkable guess, and definitely hinted at an underlying rule.

The quark picture tells us that, for example, $\Xi^{*0} = \bar{s}\bar{s}u$ (or ssu)

we then expect $B_{\text{hadrons}} = 70\%$. Look at PDG to see:

W⁺ DECAY MODES

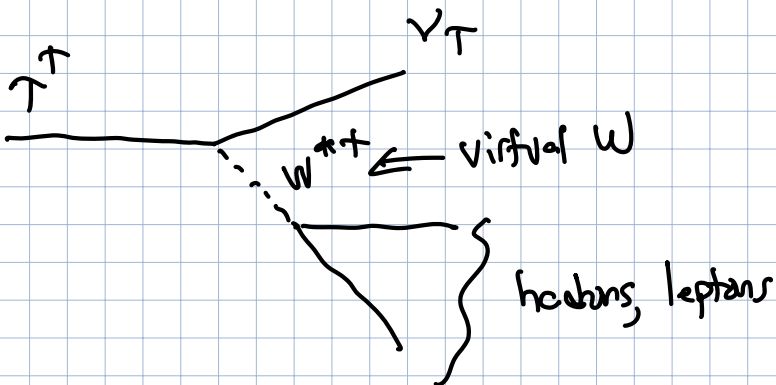
W⁻ modes are charge conjugates of the modes below.

| Mode | Fraction (Γ_i / Γ) | Scale Factor/ Conf. Level | P(MeV/c) |
|------------------------|----------------------------------|------------------------------|----------|
| Γ_1 $\ell^+\nu$ | ⁽¹⁾ (10.86 ± 0.09)% | | |
| Γ_2 $e^+\nu$ | (10.71 ± 0.16)% | 40185 | |
| Γ_3 $\mu^+\nu$ | (10.63 ± 0.15)% | 40185 | |
| Γ_4 $\tau^+\nu$ | (11.38 ± 0.21)% | 40165 | |
| Γ_5 hadrons | (67.41 ± 0.27)% | | |

$$\underbrace{\Gamma_2 + \Gamma_3 + \Gamma_4}_{\text{observed}} = (10.71 + 10.63 + 11.38)\% = 32.72\% \approx \underline{33\%} \Rightarrow \text{pretty close to expectation from } W \text{ universality!}$$

• Predict the branching fraction of $\tau^+ \rightarrow \text{leptons}$.

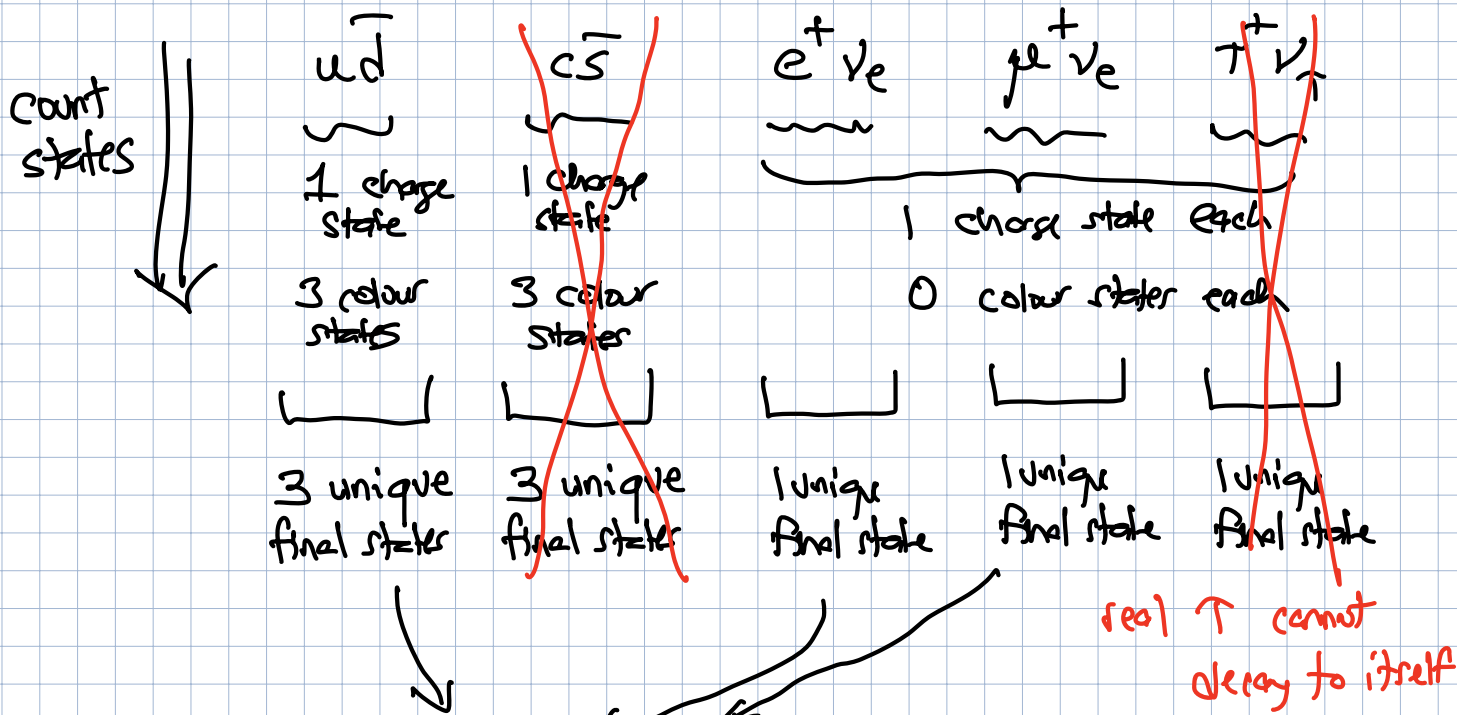
$$M_\tau = 1.78 \text{ GeV}/c^2 \longrightarrow W^+ \rightarrow c\bar{s} \text{ makes hadrons in final state.}$$



Lightest $c\bar{s}$ meson: D_s^+ , $1968 \text{ MeV}/c^2 > m_\tau!$
 Lighter charm mesons: D^+, D^0
 $\sim 1900 \text{ MeV}/c^2 > m_\tau!$
 $\tau^+ \rightarrow c\bar{s} \nu_\tau$ forbidden by energy conservation!

we state count again, adjusting for forbidden final states

⇓ see next page.



$$\tau \rightarrow \text{anything} : 3 + 1 + 1 = 5$$

$$B_{\tau \rightarrow \text{leptons}} = \frac{2}{5} = 0.4 \quad (40\%)$$

$$B_{\tau \rightarrow \text{hadrons}} = \frac{3}{5} = 0.6 \quad (60\%)$$

Compare to data:

| | | | | | |
|---|---|----------------|----------------------------------|-----|---|
| Γ_3 | $\mu^- \bar{\nu}_\mu \nu_\tau$ | ^[1] | $(17.39 \pm 0.04)\%$ | 885 | ▼ |
| Γ_4 | $\mu^- \bar{\nu}_\mu \nu_\tau \gamma$ | ^[2] | $(3.67 \pm 0.08) \times 10^{-3}$ | 885 | ▼ |
| Γ_5 | $e^- \bar{\nu}_e \nu_\tau$ | ^[1] | $(17.82 \pm 0.04)\%$ | 888 | ▼ |
| Γ_6 | $e^- \bar{\nu}_e \nu_\tau \gamma$ | ^[2] | $(1.83 \pm 0.05)\%$ | 888 | ▼ |
| Γ_7 | $h^- \geq 0 K_L^0 \nu_\tau$ | | $(12.03 \pm 0.05)\%$ | 883 | ▼ |
| Γ_8 | $h^- \nu_\tau$ | | $(11.51 \pm 0.05)\%$ | 883 | ▼ |
| Γ_9 | $\pi^- \nu_\tau$ | ^[1] | $(10.82 \pm 0.05)\%$ | 883 | ▼ |
| Γ_{10} | $K^- \nu_\tau$ | ^[1] | $(6.96 \pm 0.10) \times 10^{-3}$ | 820 | ▼ |
| Γ_{11} | $h^- \geq 1 \text{ neutrals } \nu_\tau$ | | $(37.00 \pm 0.09)\%$ | | ▼ |
| ▼ Modes with three charged particles | | | | | |
| Γ_{62} | $h^- h^- h^+ \geq 0 K_L^0 \nu_\tau$ | | $(15.20 \pm 0.06)\%$ | 861 | ▼ |

$$\Gamma_3 + \Gamma_5 = 17.39\% + 17.82\% = 35.21\% \approx 35\%$$

$$\Gamma_8 + \Gamma_{11} + \Gamma_{62} = 11.51\% + 37\% + 15.2\% = 63.71\% \approx 64\%$$

99% \approx

So compare 35% (observed) to 40% (estimated) \rightarrow
not bad!

The simple requirements imposed by the SM makes for
powerful constraints.