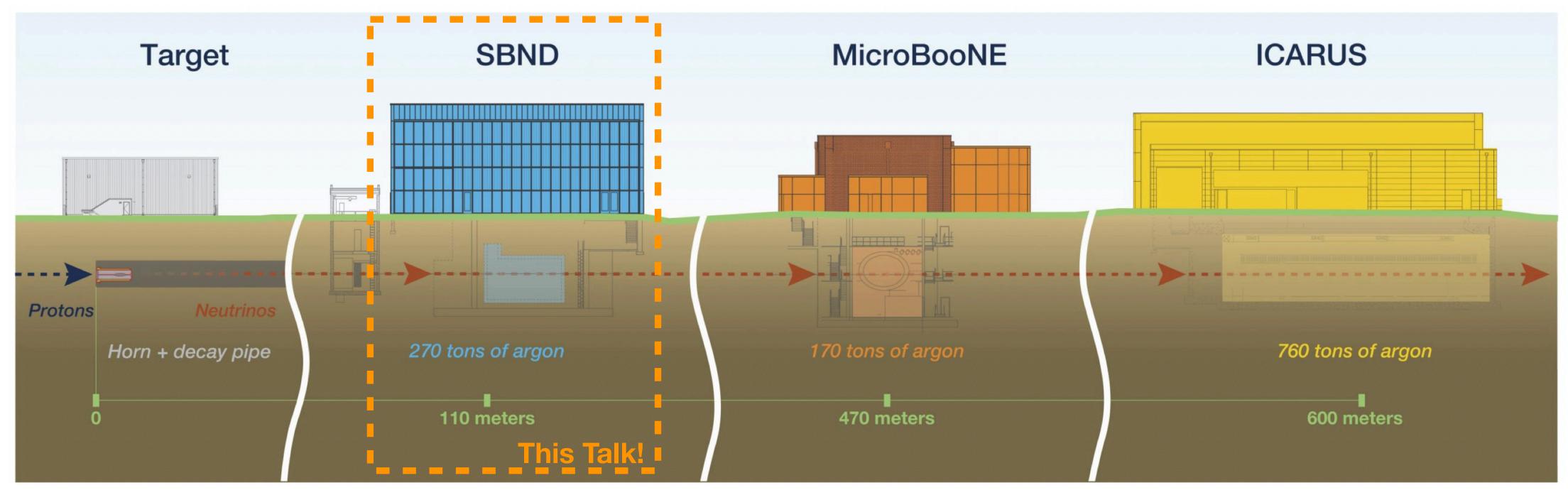




The SBN Program at Fermilab

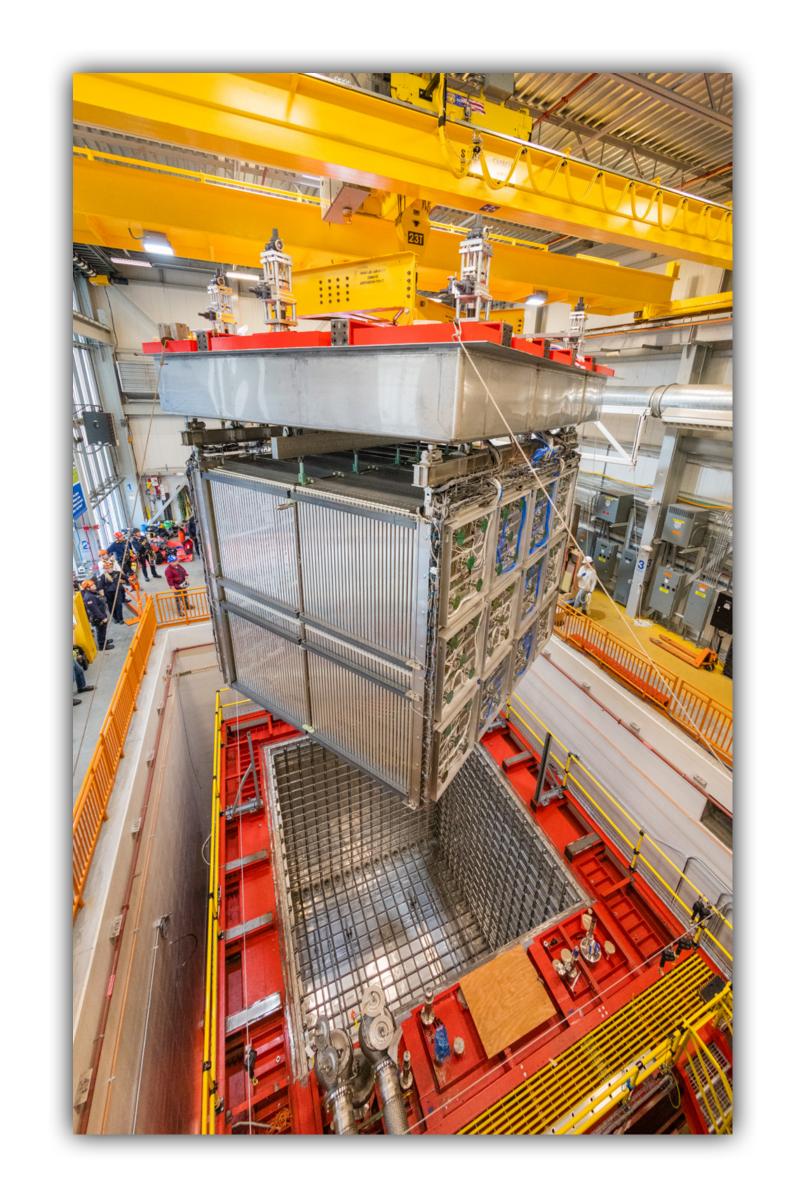


- The Short Baseline Neutrino Program aims to resolve the low energy excess seen at MiniBooNE (Phys. Rev. Lett. 121, 221801(2018))
- Multi-detector facility sitting in the Booster Neutrino Beam at Fermilab → Pion decay-in-flight source (PRD 79, 072002)
 - Same as MiniBooNE!
- By using the same detector technology, target nucleus, and beamline, the SBN program can constrain systematic
 uncertainties to the %-level



The Short Baseline Near Detector (SBND)

- Liquid Argon Time Projection Chamber (LArTPC)
- Near detector of the SBN Program 110 meters from the source
- Finished collecting our first run of neutrino data this summer (Run 1: December 2024 — July 2025)
 - ~3 million neutrino interactions!
- Physics goals:
 - 1. Constrain beam and cross section uncertainties for SBN sterile neutrino search
 - 2. Search for Beyond the Standard Model Physics and study rare processes
 - 3. Precision measurements of neutrino interactions on Argon (Cross sections)

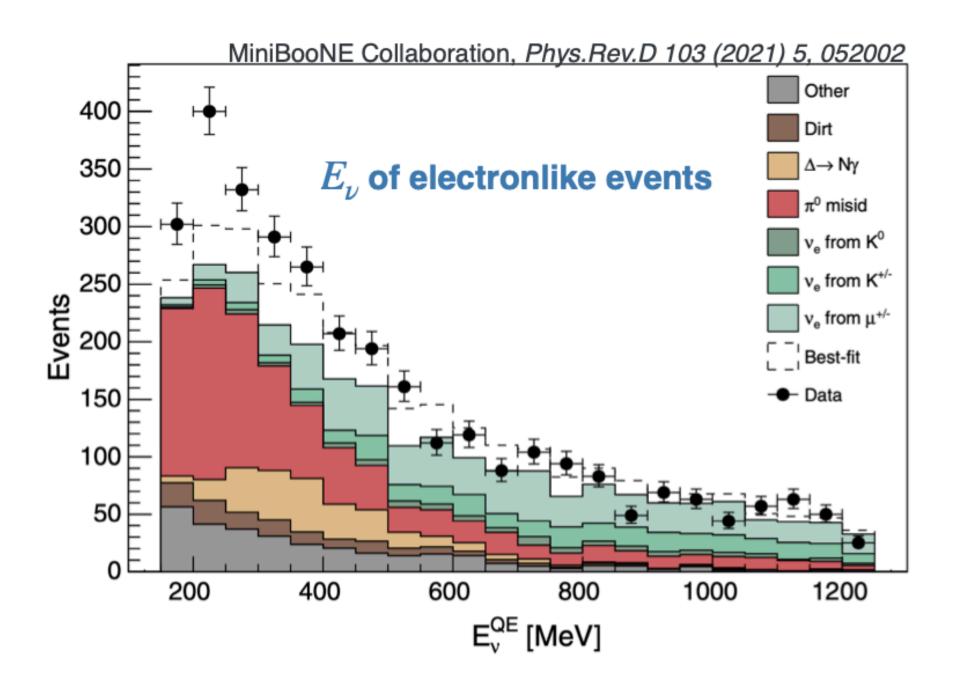




Physics Motivation



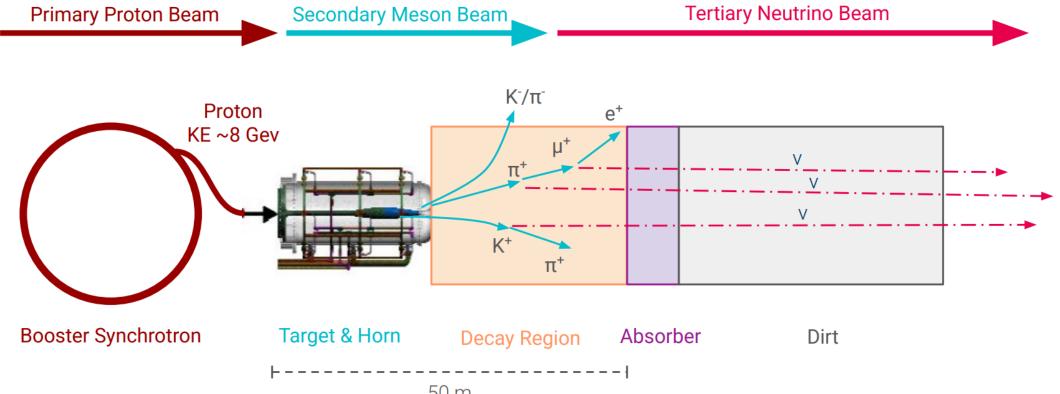
The SBN Oscillation Program



- Low energy excess in electron-like events was seen by **MiniBooNE**
- Used accelerator-based ν_u dominated source
 - The Booster Neutrino Beam (BNB)
- Motivates a search for sterile neutrinos at short baselines (~1eV scale in 3+1 model)

- SBN program sits in the same beamline and has both a near (SBND) and far (ICARUS) detector with the same detector technology
- LAr-TPC provides much better particle ID

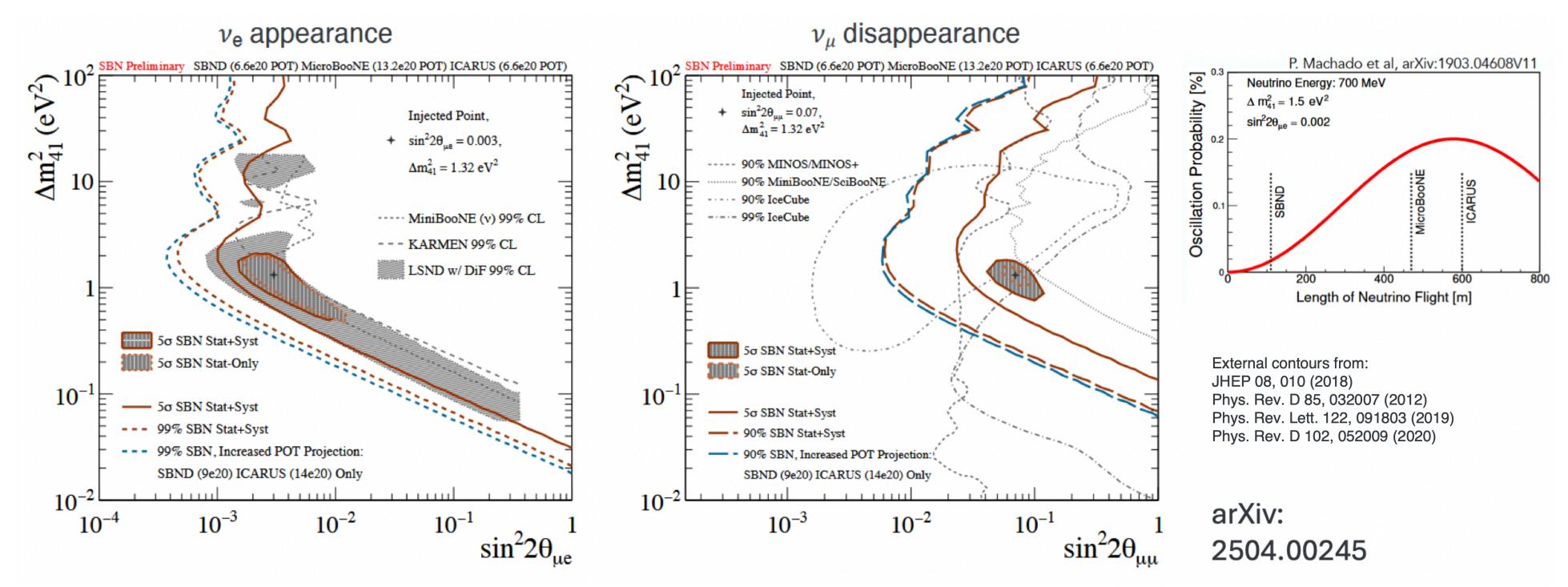
The Booster Neutrino Beam





The SBN Oscillation Program

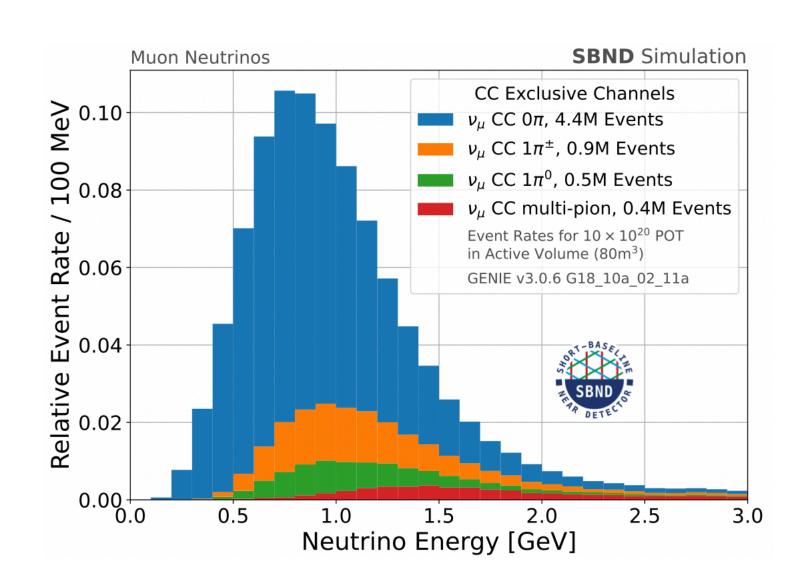


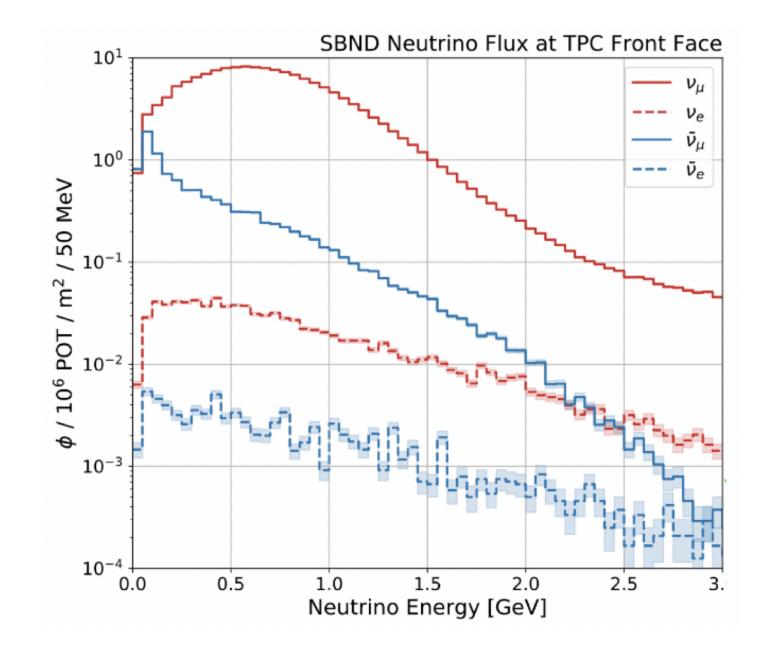


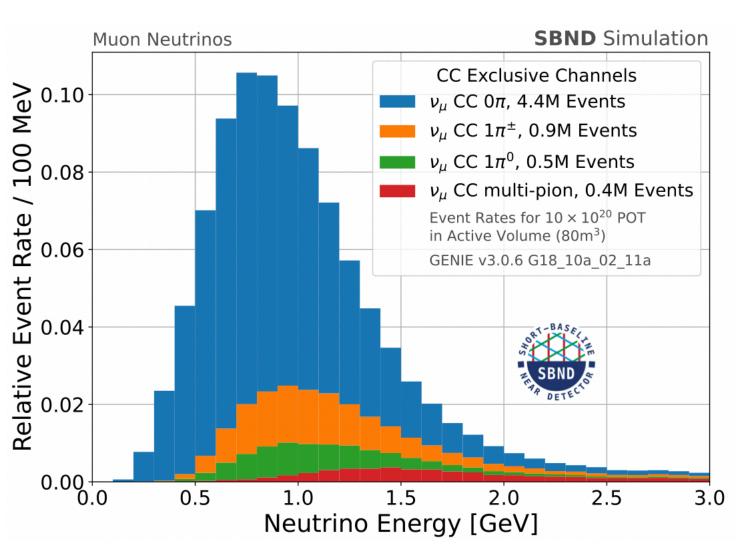
- SBND will constrain beam and cross section uncertainties as the near detector
- Need both ν_e appearance and ν_μ disappearance searches to conclusively address the sterile neutrino hypothesis



- SBND will collect unprecedented interaction statistics
 - Currently have the world's largest ν -Argon dataset
 - Estimate ~2 million ν_{μ} CC and ~15,000 ν_{e} CC interactions per year \rightarrow achieved in Run1 !
 - ~10 million neutrino interactions projected for the full data taking period of around 3 years



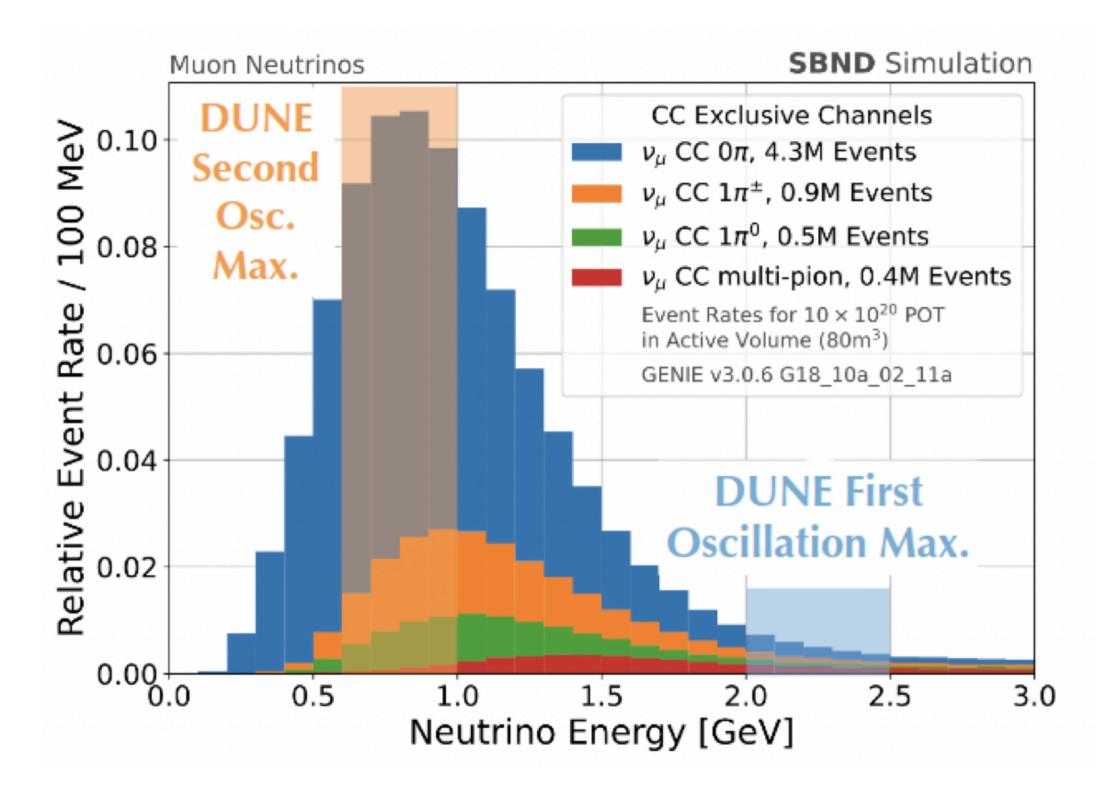


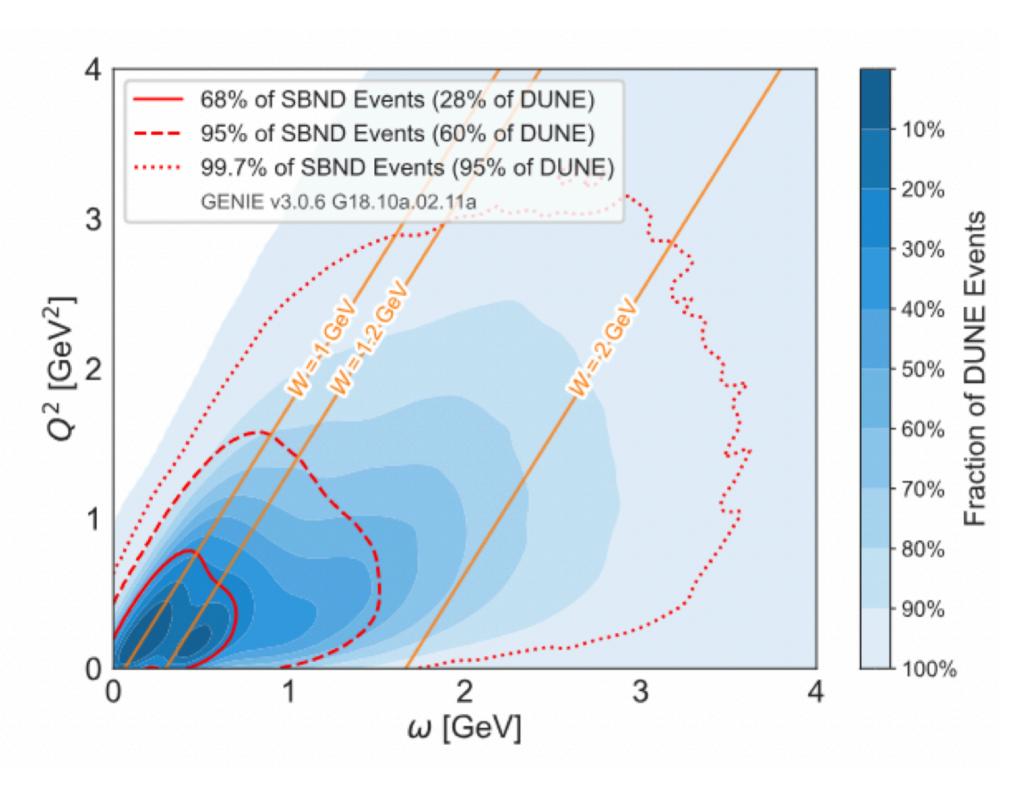




SBND Neutrino Flux: DUNE Coverage

- SBND has significant kinematic phase space overlap with DUNE
 - BNB flux peaks near the second oscillation maximum for DUNE (~0.8 GeV)
 - SBND can be used to constrain DUNE cross section uncertainties



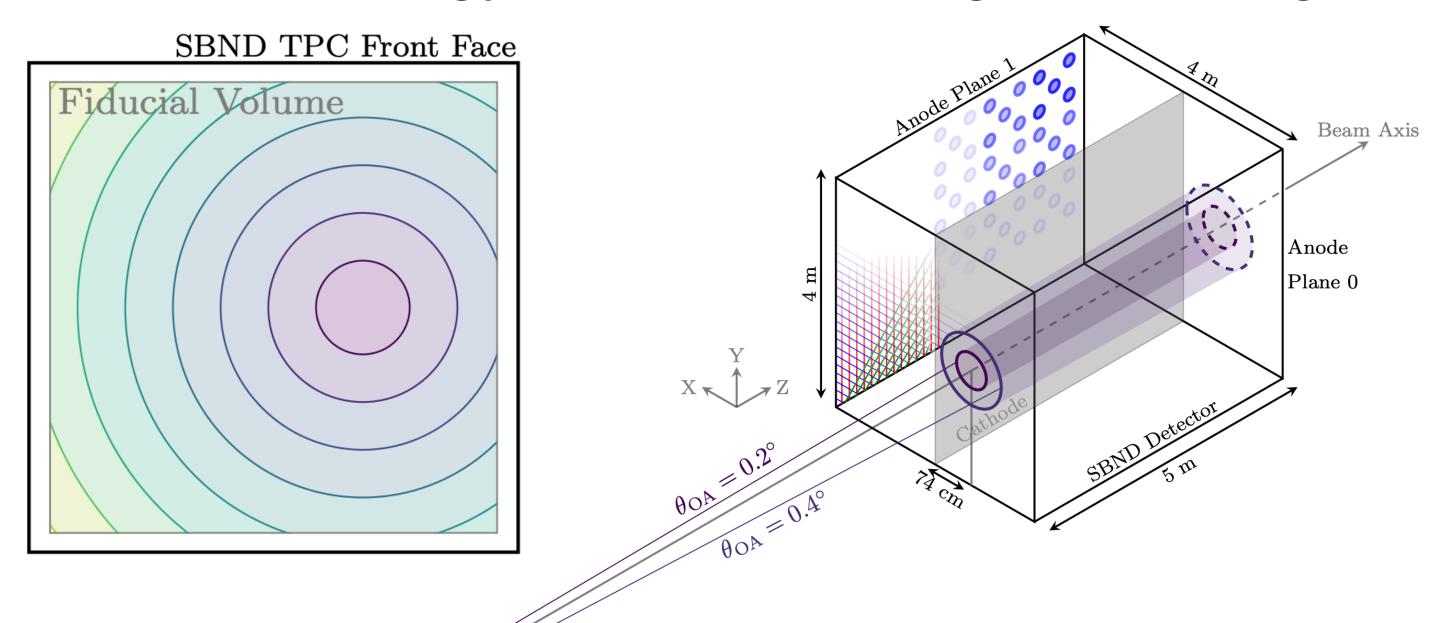


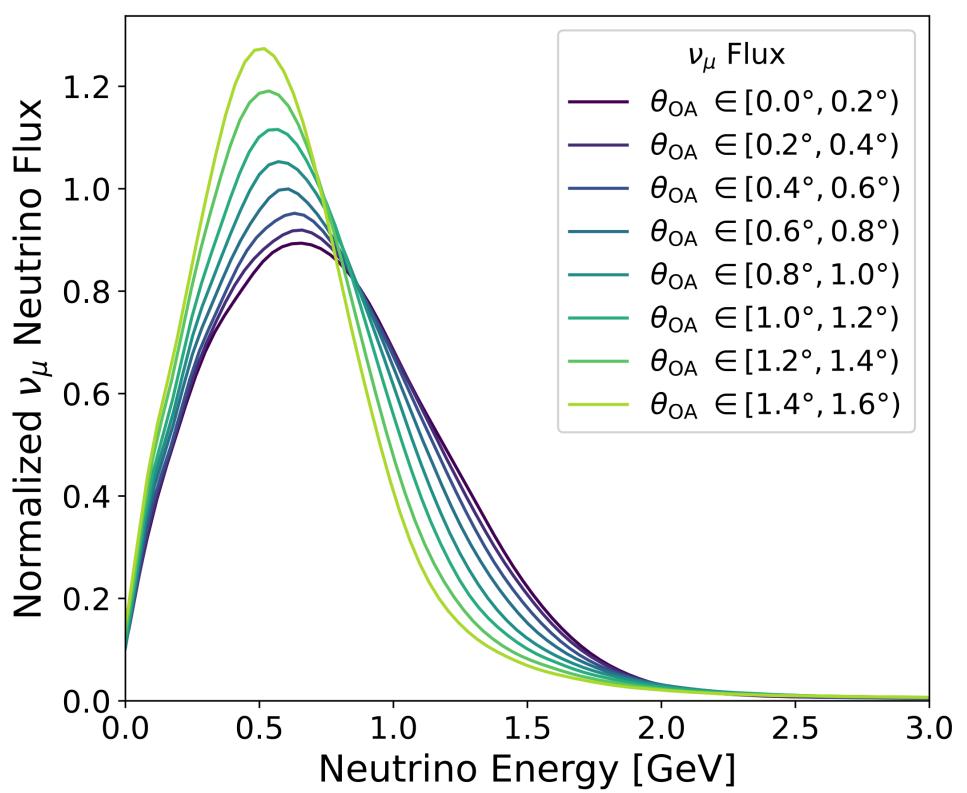
8



SBND Neutrino Flux: PRISM

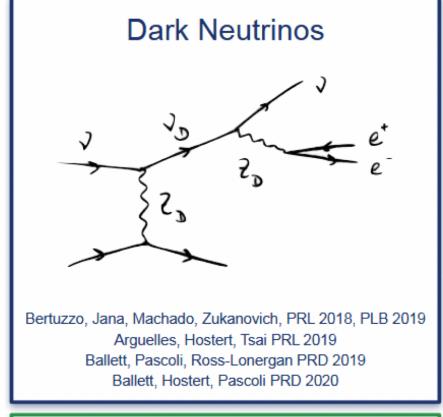
- Precision Reaction Independent Spectrum Measurement: arxiv:2508.20239
- SBND sits 74 cm off-axis and its close proximity to the beam allows it to probe an angular range of [0, 1.6] degrees
- Neutrino energy spectrum changes with angle

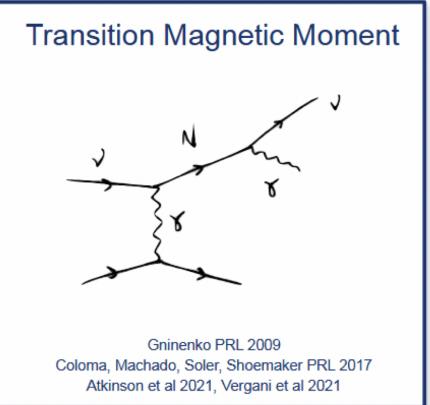


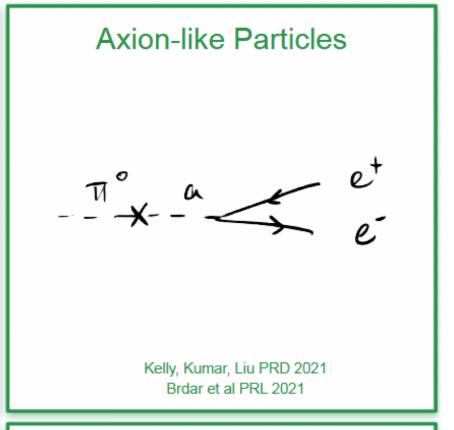


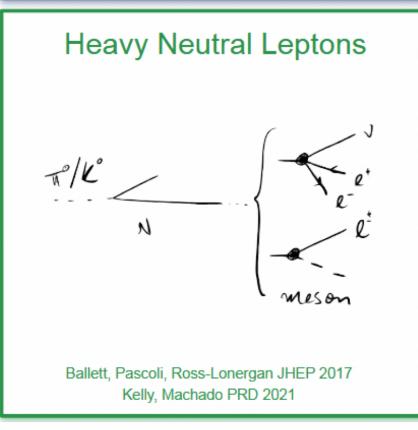


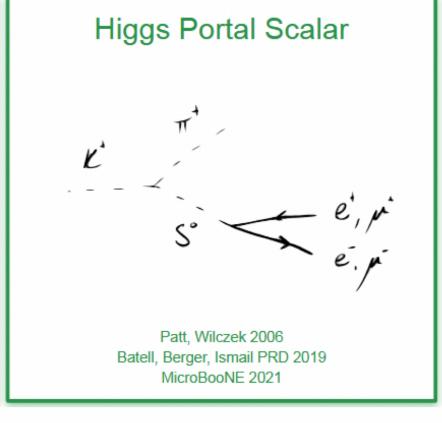
Beyond the Standard Model (BSM)











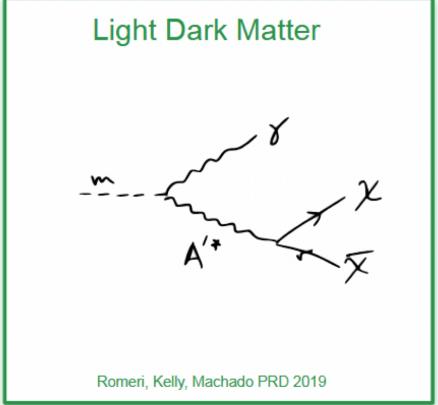
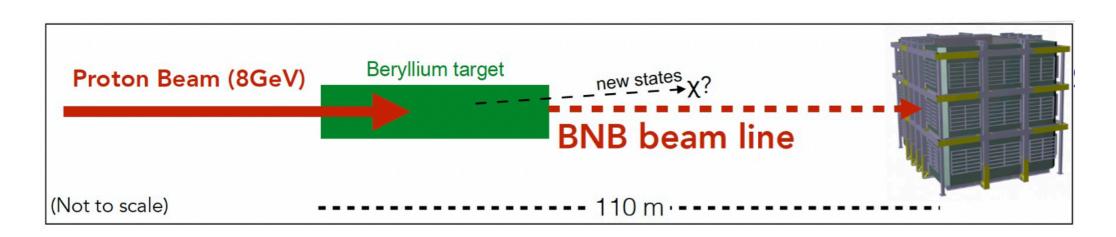


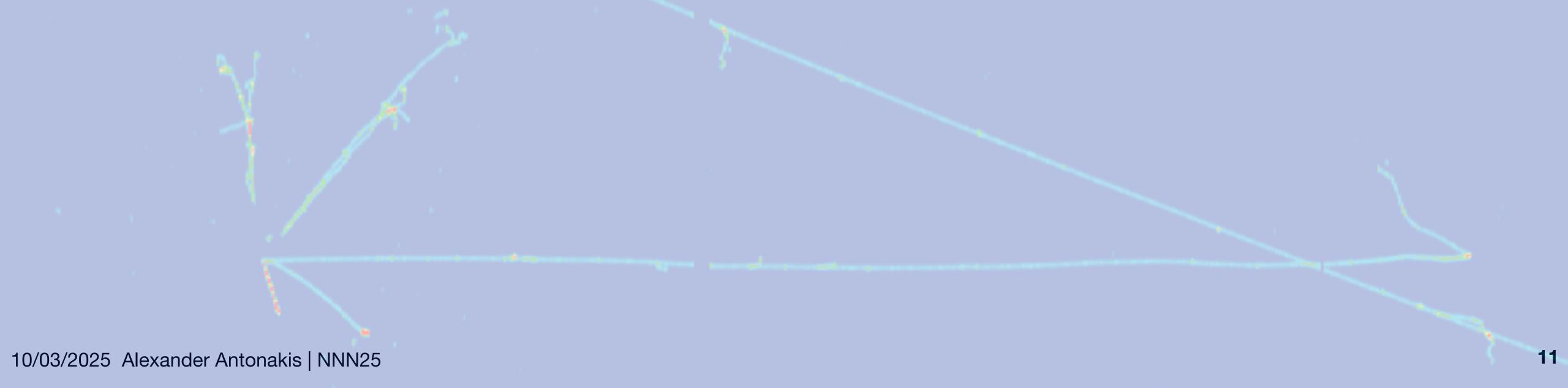
Image credit: Pedro Machado, Marco Del Tutto



- SBND's proximity to the beam and high statistics enables BSM studies
- Developing advanced timing reconstruction based on scintillation light will help separate massive long-lived particles from neutrinos based on time-of-flight
- Competitive sensitivities to many dark sector particles



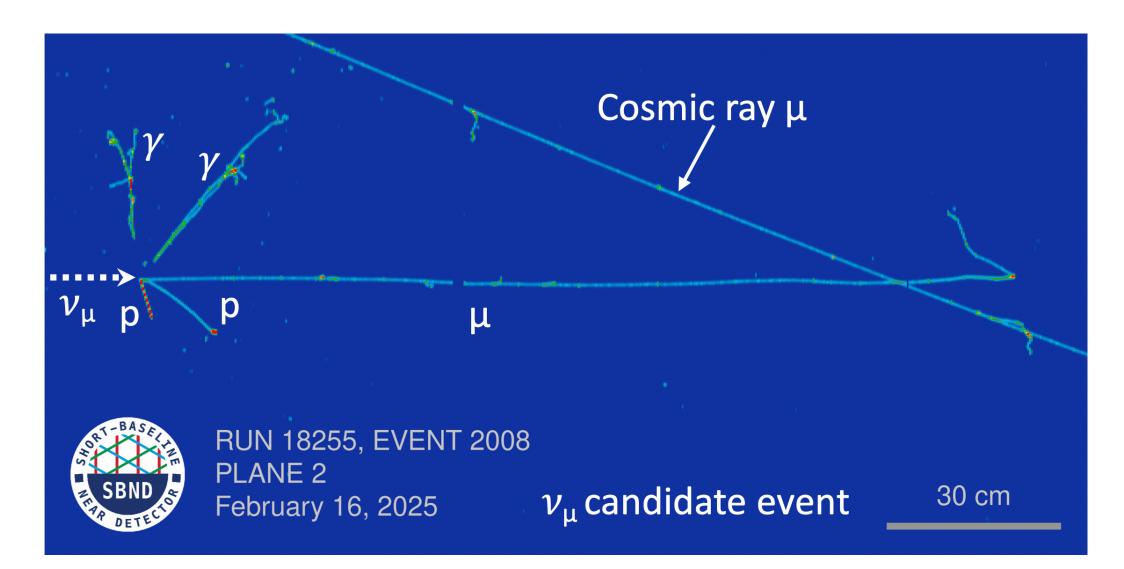
Detector Overview

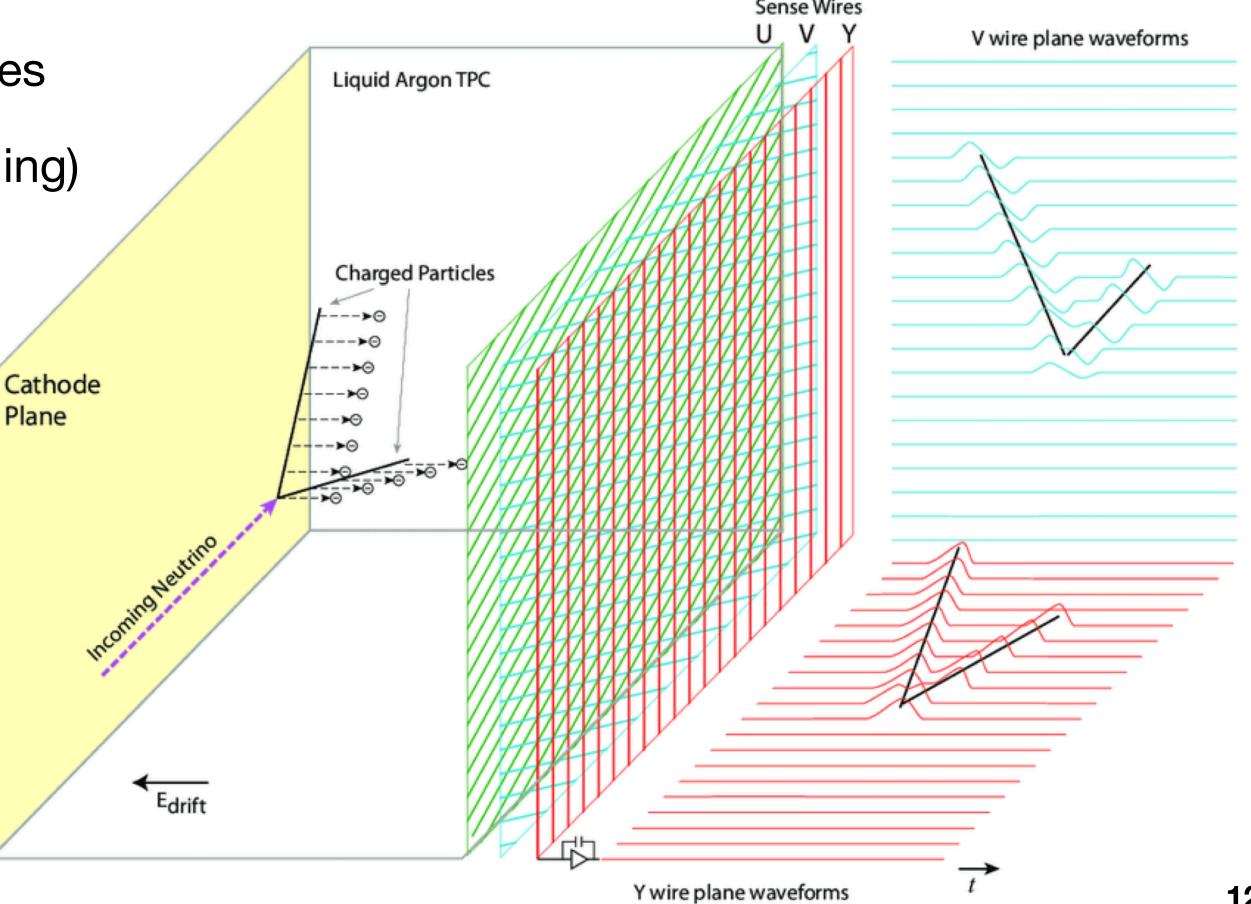




Liquid Argon Time Projection Chambers

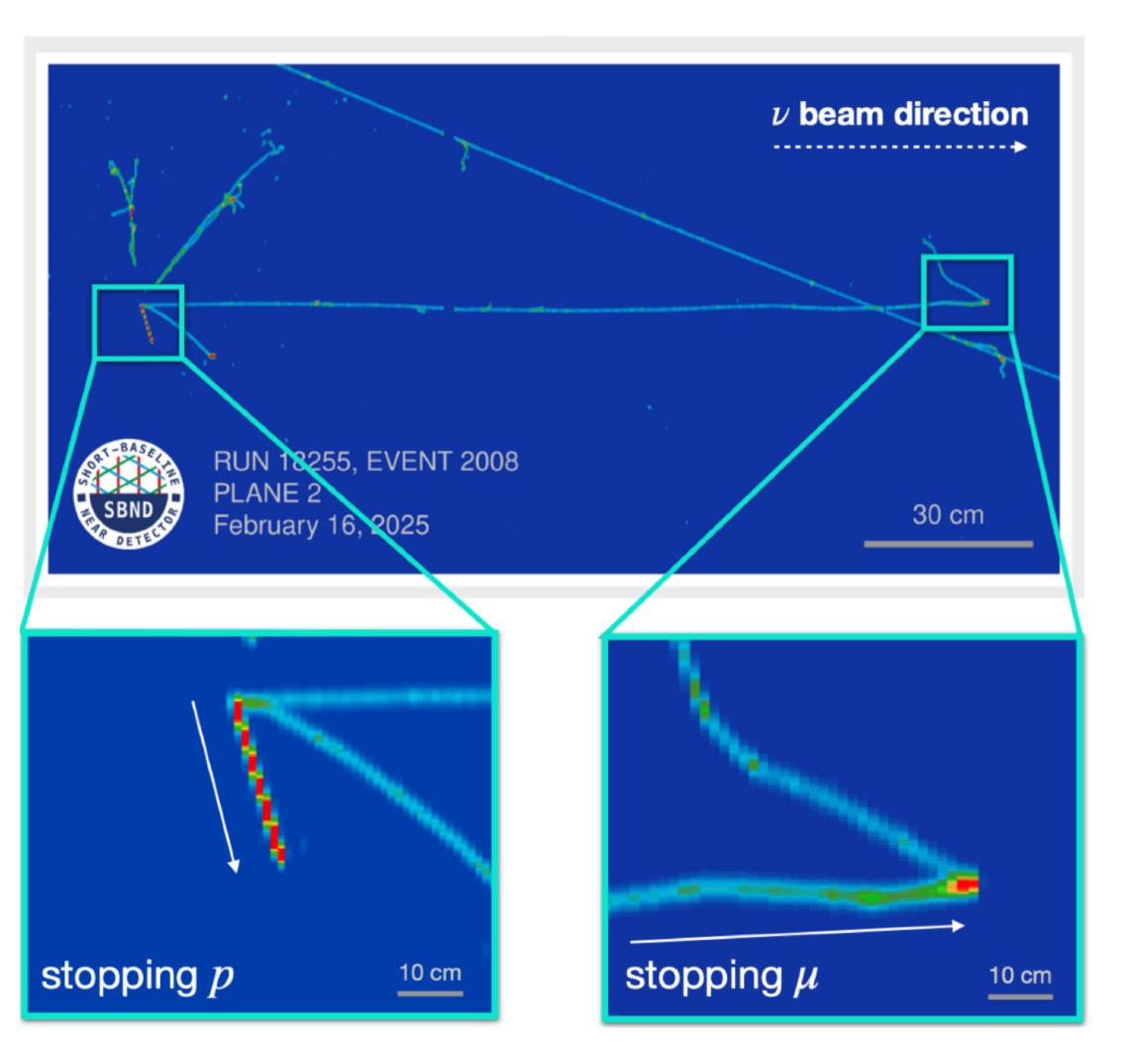
- SBND is a Liquid Argon Time Projection Chamber (LAr-TPC)
 - Charged particles deposit energy → producing ionization electrons and light
 - Ionization electrons drift towards wire readout planes
 - Scintillation light is recorded by PMTs (precision timing)
 - Millimeter-level 3D reconstruction





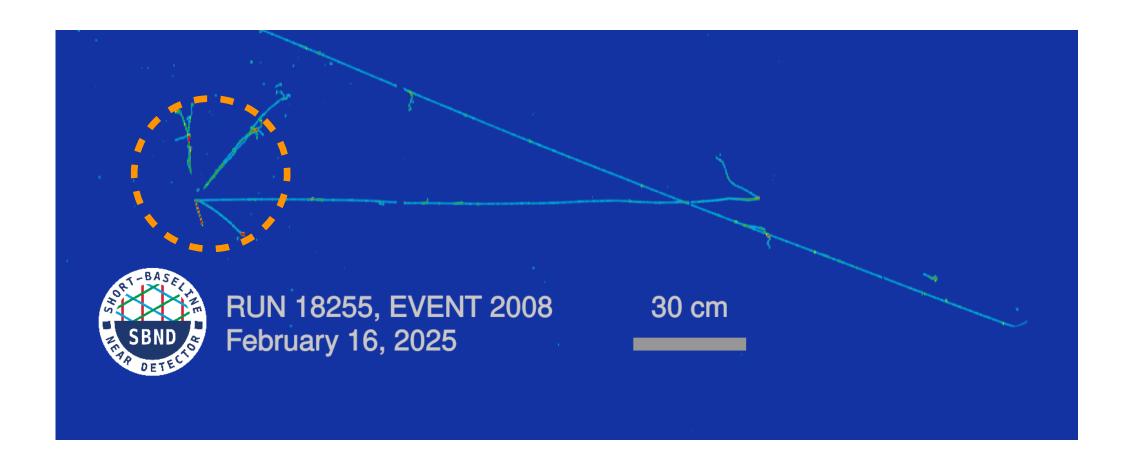


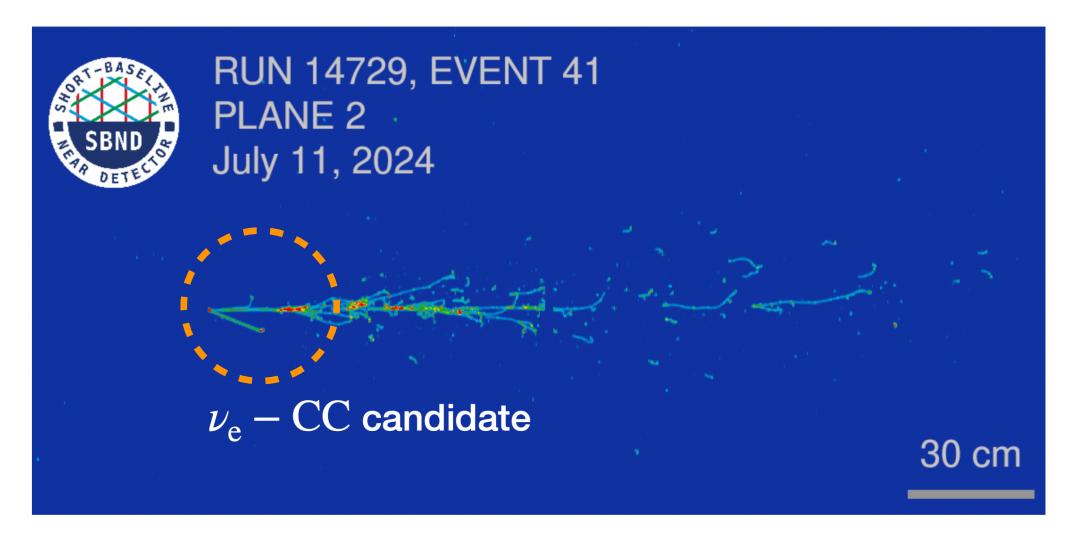
- LAr-TPCs provide calorimetry and 3D tracking with fine-grained information
 - Capable of identifying different species of particles
 - Particle Flow: parent vs daughter particles
 - Track vs Shower, Interaction Vertex, etc

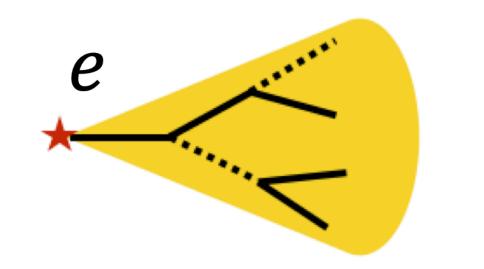


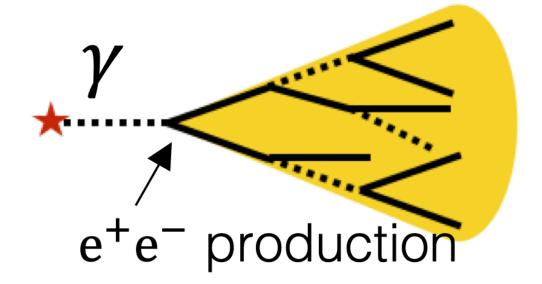


- LAr-TPCs provide calorimetry and 3D tracking with fine-grained information
 - Capable of identifying different species of particles
 - Particle Flow: parent vs daughter particles
 - Track vs Shower, Interaction Vertex, etc
 - Electron vs Photon Discrimination
 - Critical for oscillation program



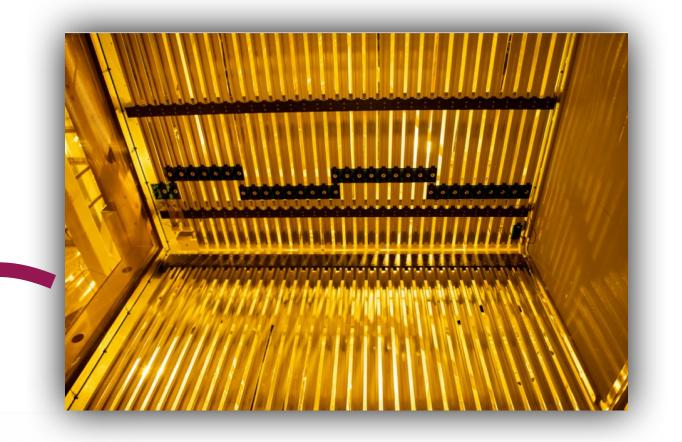








The SBND LAr-TPC



Field Cage for uniform 500 V/cm electric field



Biased at -100 kV

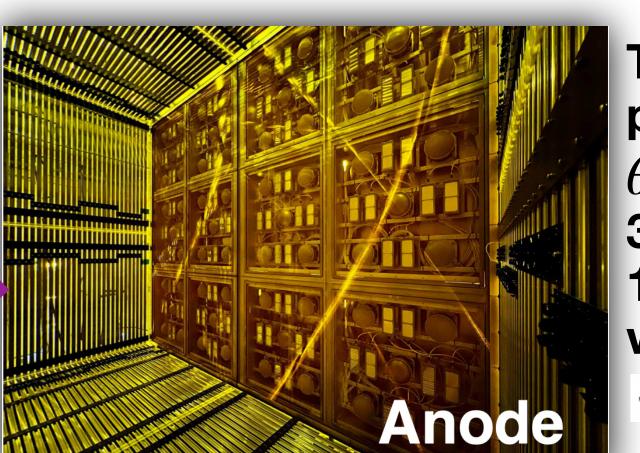


Cold electronics amplify and digitize at 89 K



TPC sits in a membrane cryostat same kind used in DUNE

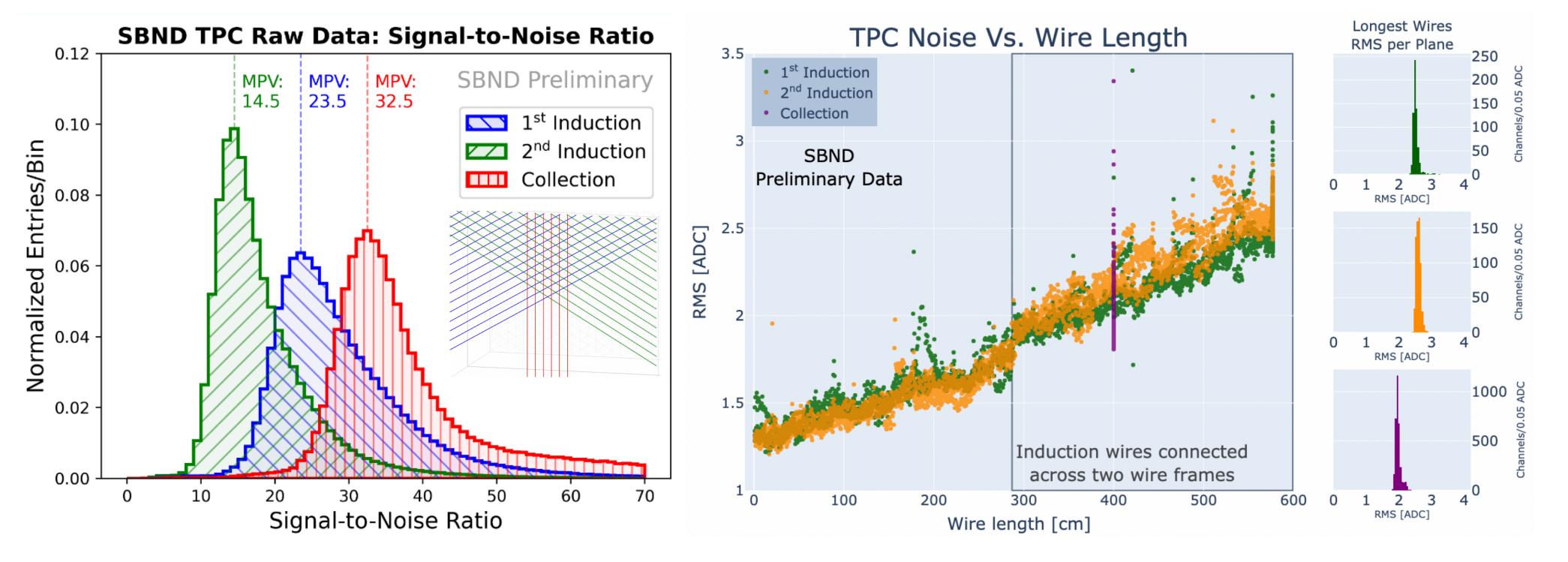
Held at ground

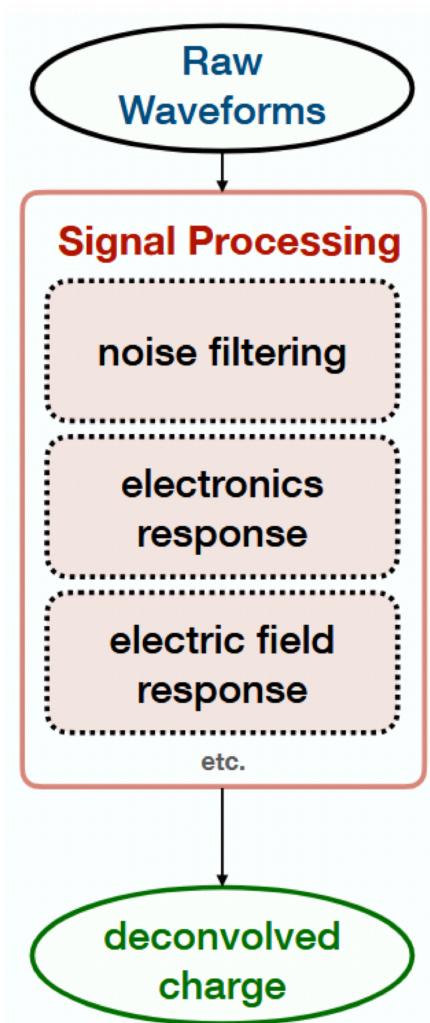


Three wire planes $\theta = 0^{\circ}, \pm 60^{\circ}$ 3mm spacing 11,264 total wires

JINST 15, P06033 (2020)

TPC Noise Performance

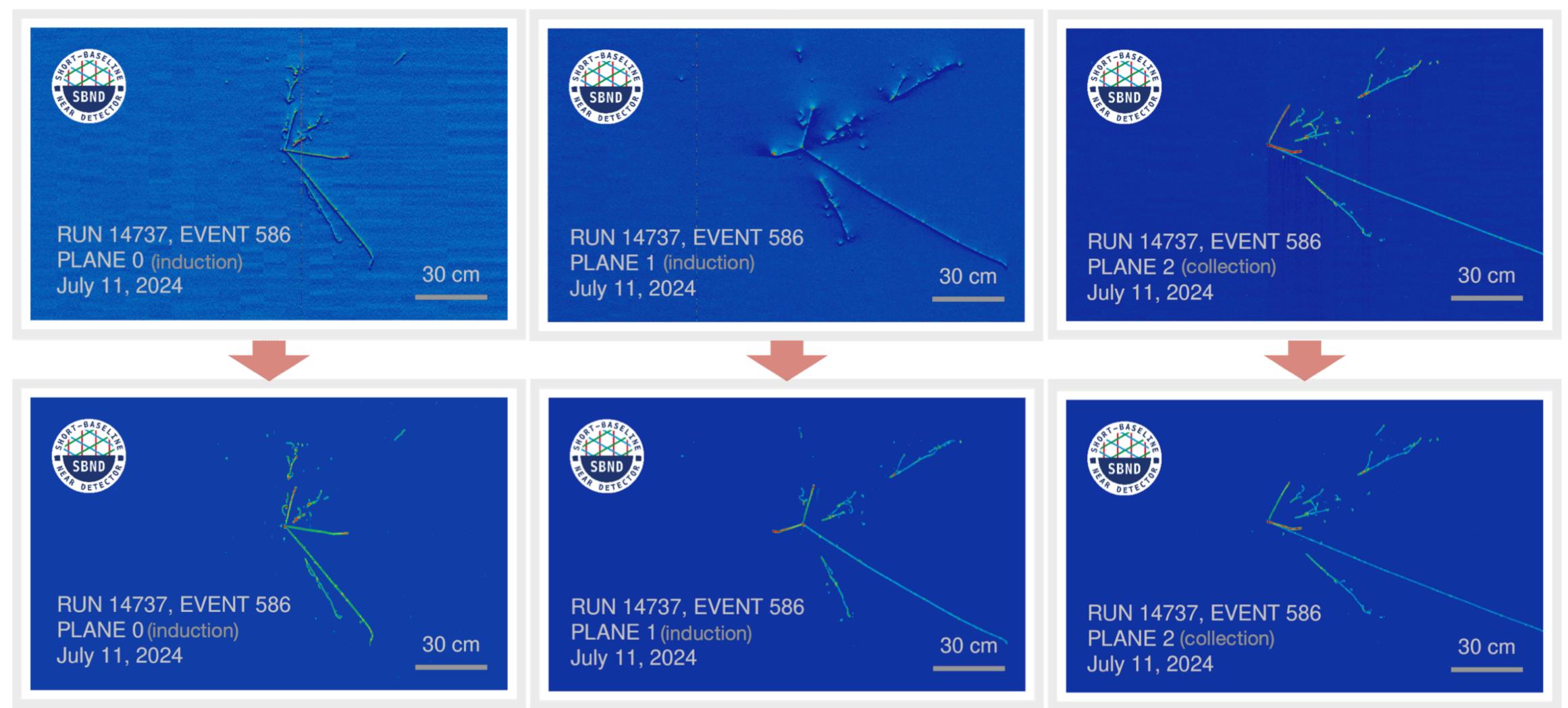




- For the longest wires, rms is less than 3 ADCs in general
- Excellent signal to noise ratio in raw data: 32.5 for the collection plane



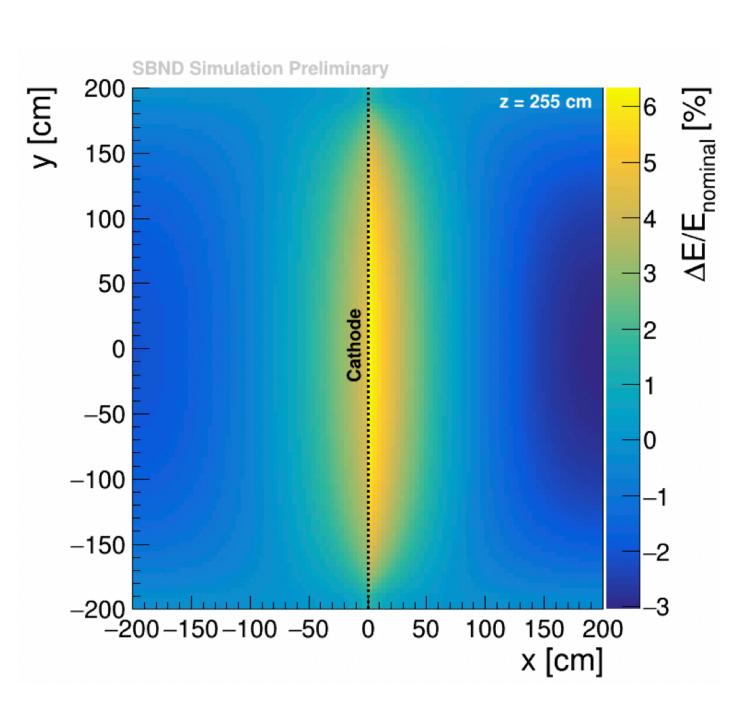
TPC Signal Processing

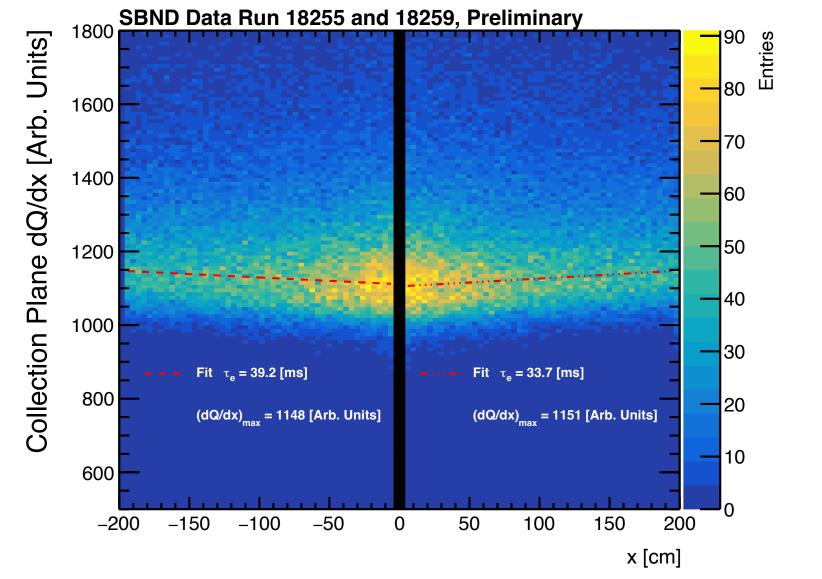


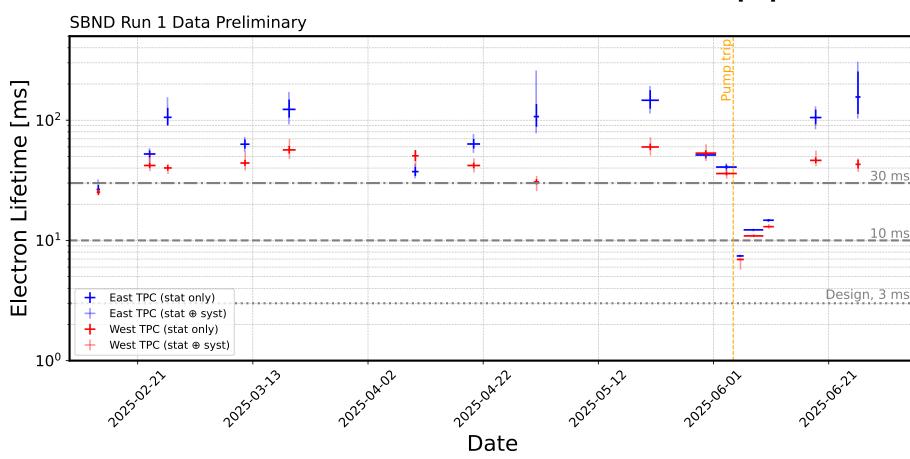


TPC Calibration

Space charge effect is well-calibrated: average effect is < 1cm across the detector

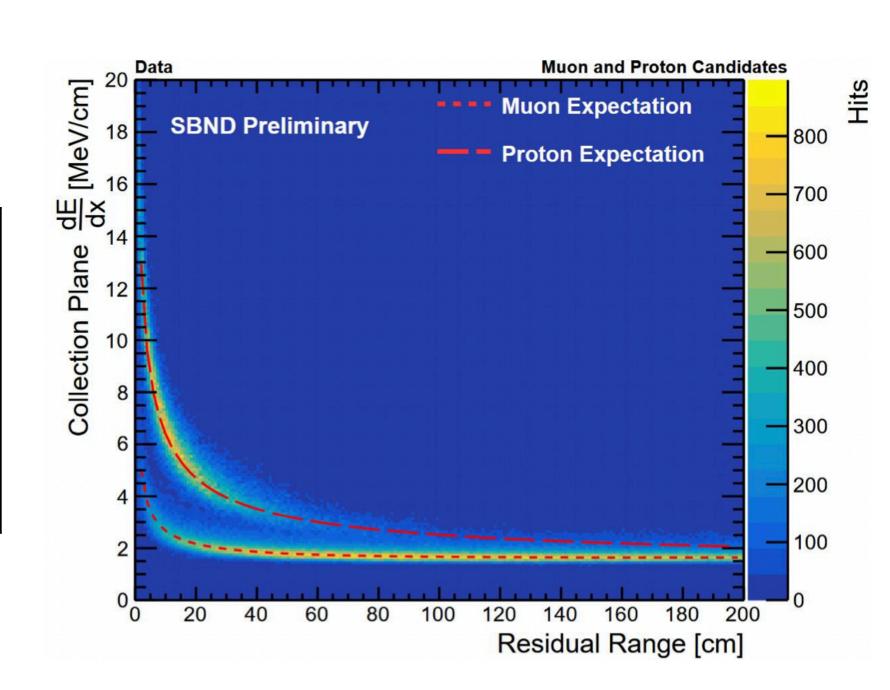






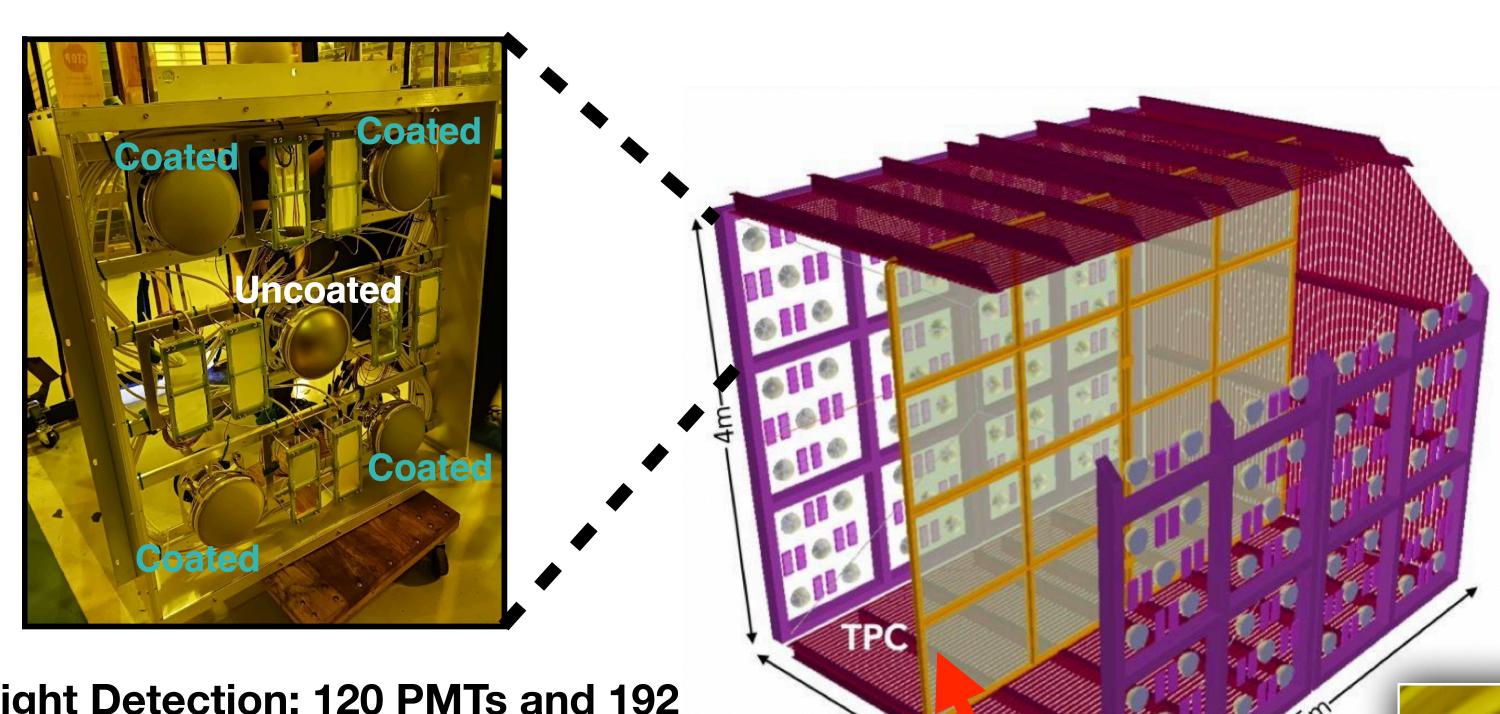
Electron lifetime is longer than 30 ms and stable!

Muon and proton dE/dx after calibration agree well with expected Landau-Vavilov theory





The SBND Photon Detection System (PDS)



TPB-coated cathode scintillation point vuv light point visight-spot" on cathode

 Light Detection: 120 PMTs and 192 X-ARAPUCAs

 Coated vs uncoated allows for independent spatial reconstruction

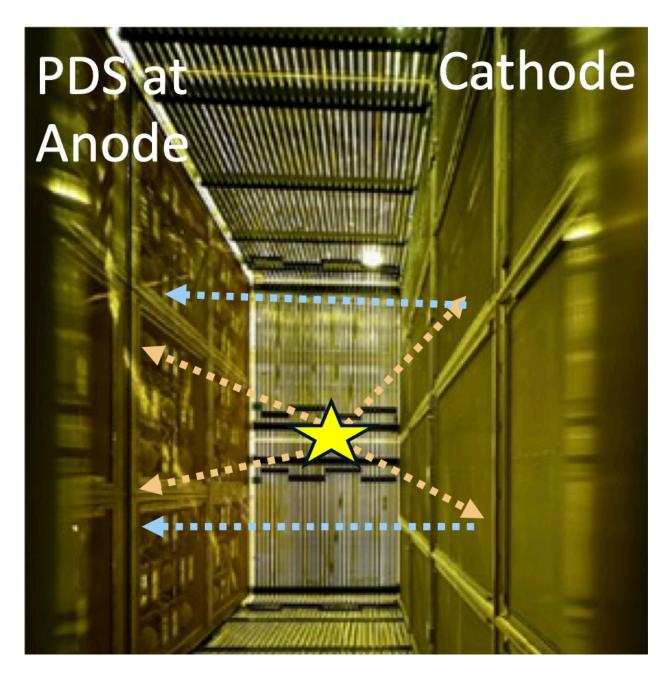
Timing Resolution is ~O(1ns)

Wavelength shifting reflective foils at cathode

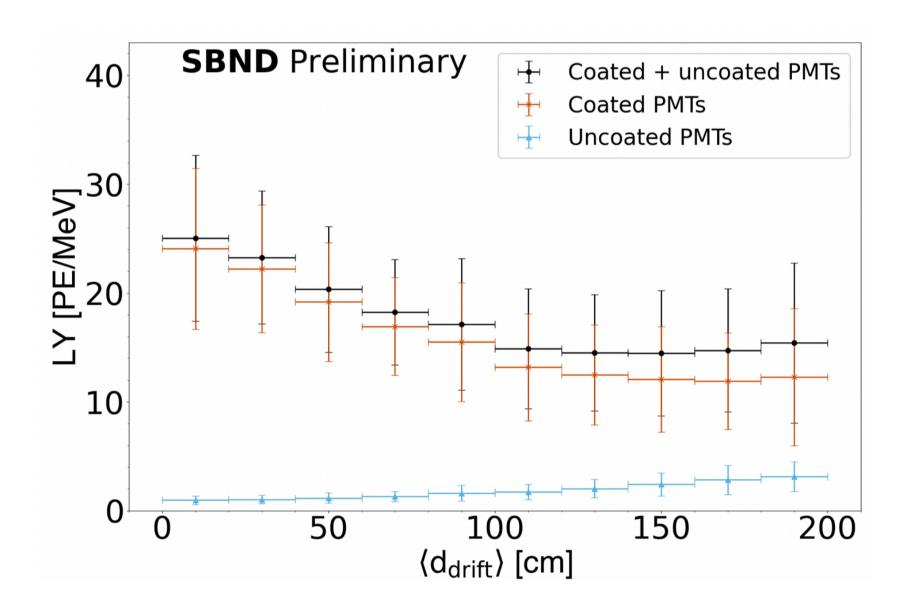
Increase uniformity of the light yield

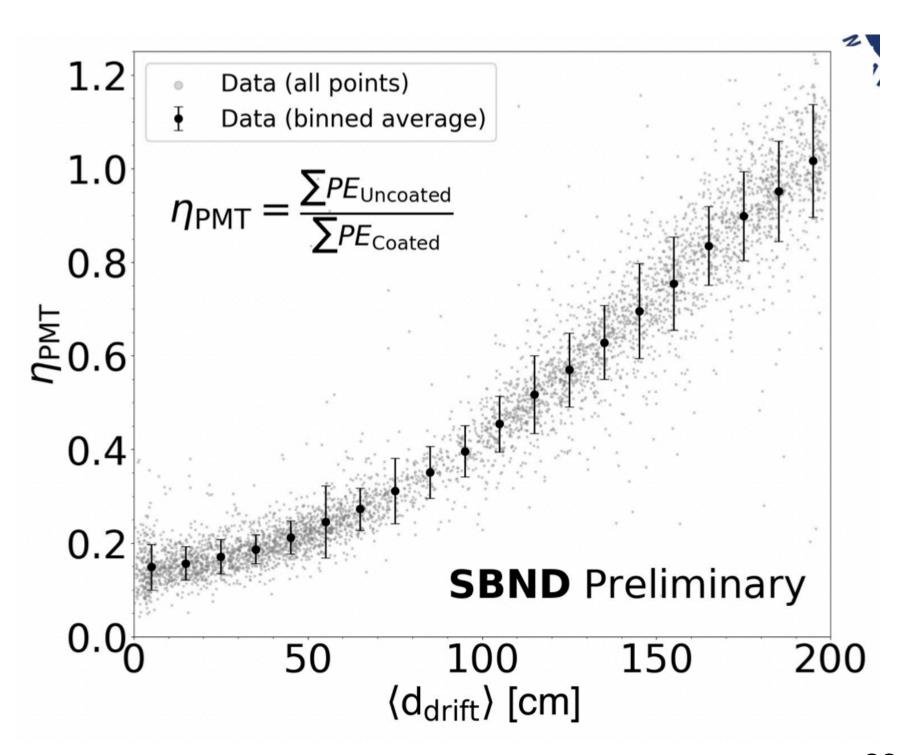


PDS Performance



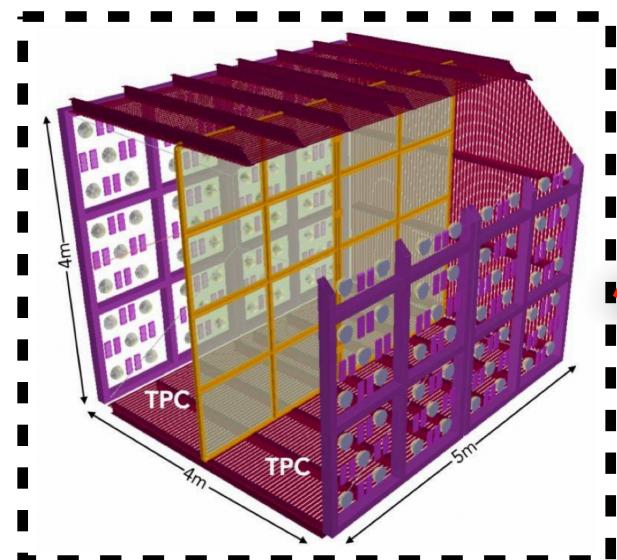
- Primary scintillation light is measured by the coated PMTs
- Uncoated PMTs measure the VUV scintillation light that reflects of the cathode
- Position reconstruction using the light: η_{PMT} is a proxy for drift distance
- The PDS has excellent timing resolution: ~O(1ns)
- Provides the basis for SBND's trigger system and can fully resolve the BNB spill structure (see next slides)





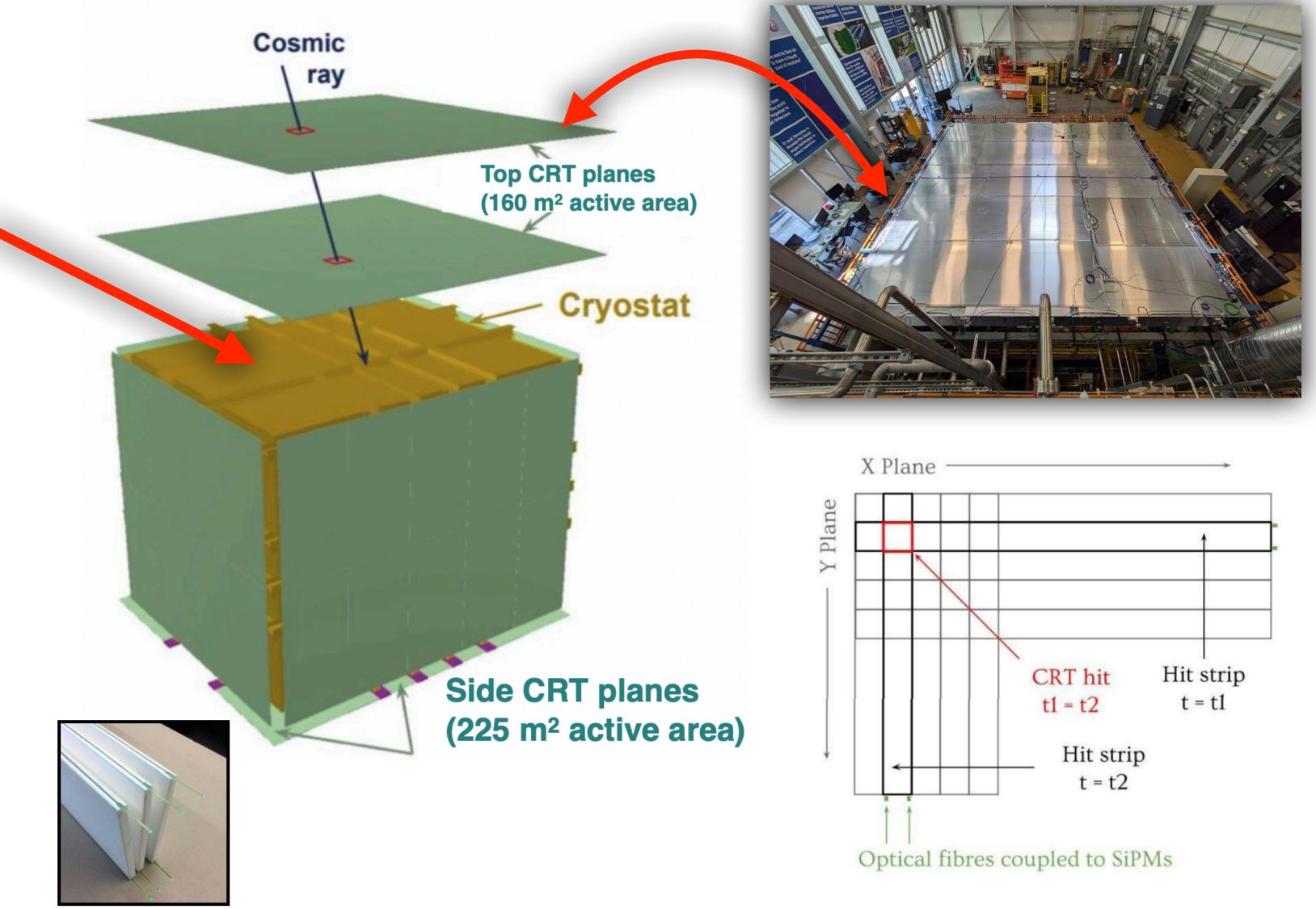


The SBND Cosmic Ray Tagger (CRT)



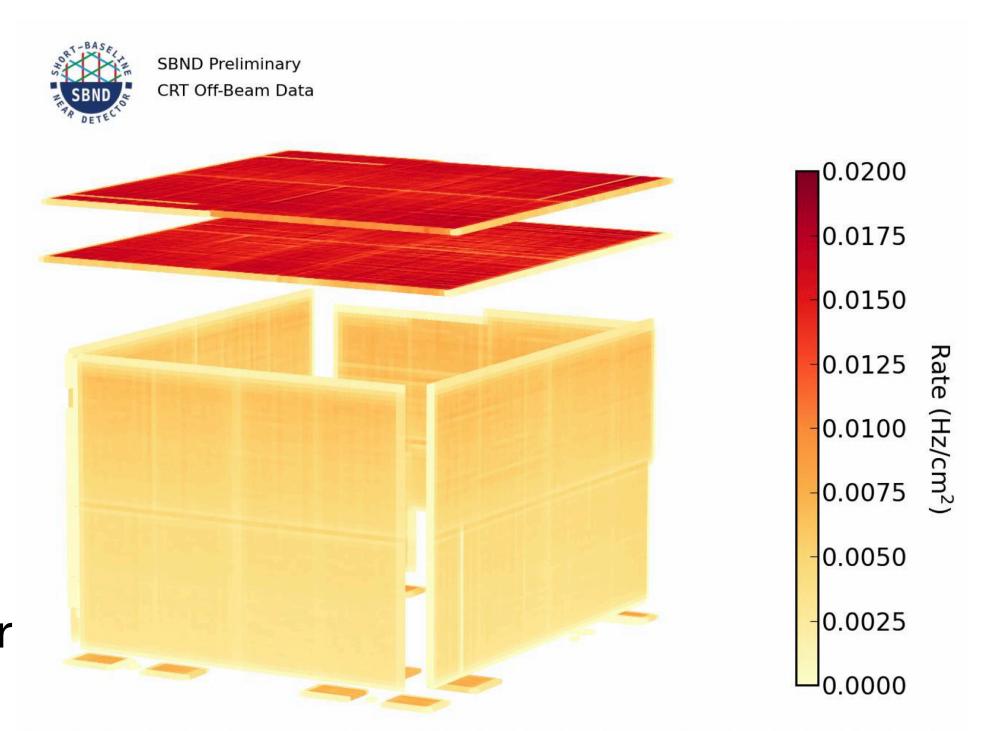
External particle tracker made of overlapping perpendicular scintillator bars:

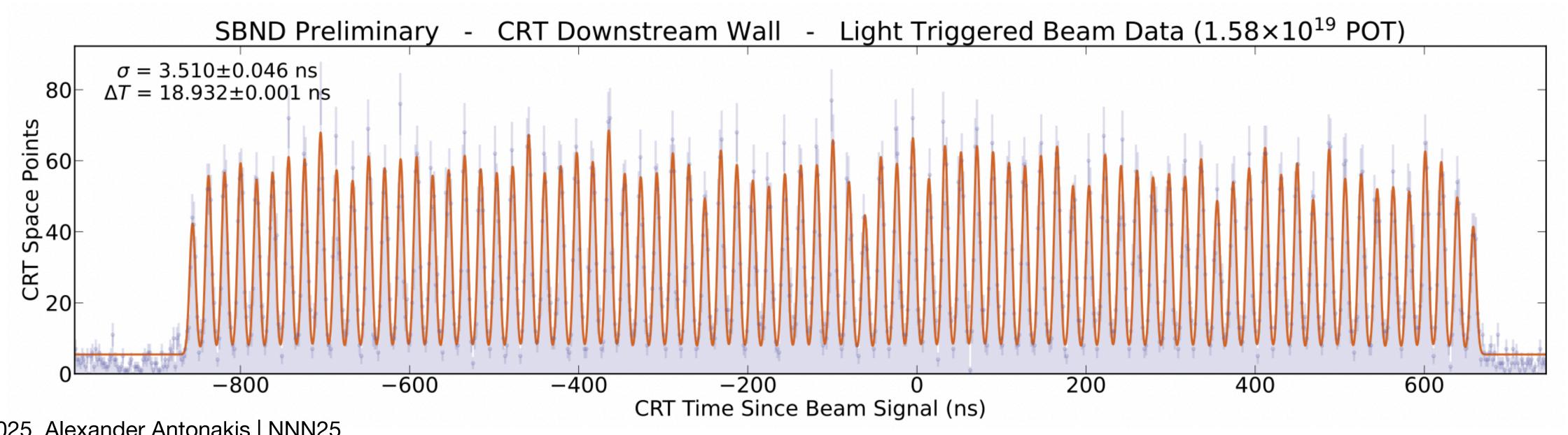
- Wavelength shifting optical fibers
- Two Silicon Photomultipliers per bar for light detection





- SiPM voltage and gain calibration give a highly uniform response with excellent hit efficiency (> 90%)
- CRT can reproduce all 81 BNB proton bunches with ~3.3 ns timing resolution
- The CRT will be an excellent tool for cosmic rejection and vetoing neutrino interactions that happen outside the detector







September 2022

October 2022



TPC and PDS Completed

Cryostat Completed at ND Building



TPC Transported to ND Building

TPC Lowered into Cryostat

March 2024

Detector Filled with LAr



July 2024

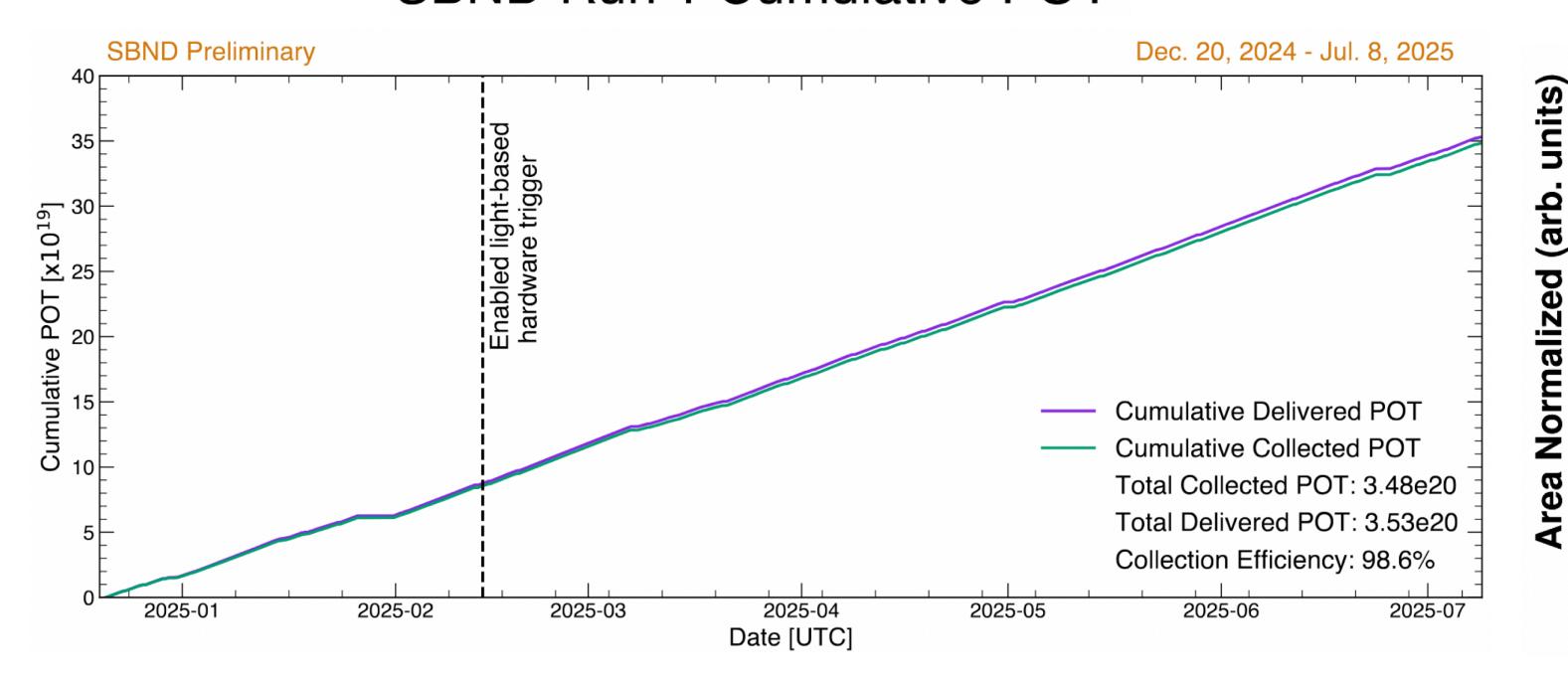
Physics data!

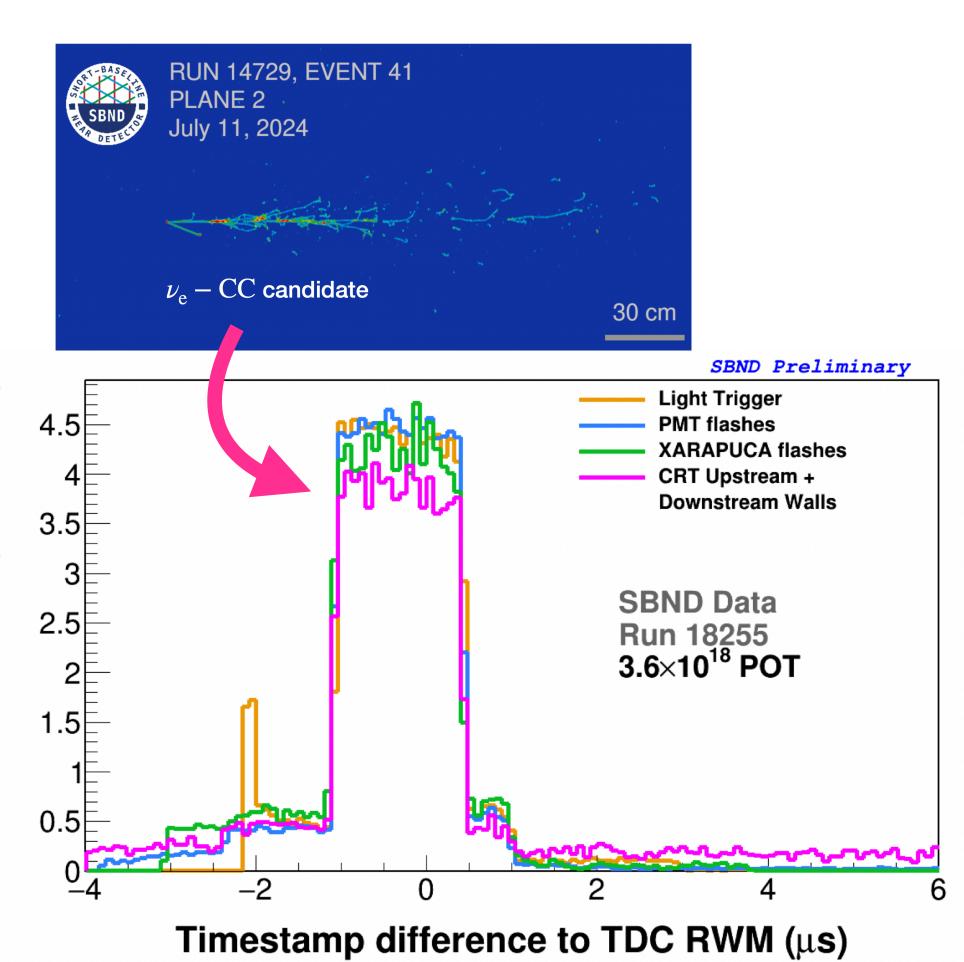


SBND Operations/Data Taking

- Run 1 data collection is complete (December 2024 July 2025): \sim 3.5 x 10^{20} POT
- > 3 million neutrino interactions collected so far!
- Already have the world's largest $\nu-$ Argon interaction dataset

SBND Run 1 Cumulative POT





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Status of First Analyses



SBND Cross Sections

- SBND's high statistics and LAr-TPC technology enables a broad cross section program
- Many analyses in progress:

Coherent Pion Production

- ν_{μ} CC Inclusive
- ν_{μ} CC1p0 π
- ν_{e} CC Inclusive

Previewed here and will be shown at NuInt 2025!

- ν_{μ} CC 2p0 π
- ν_{μ} CC π^0
- ν_u CC Shallow Inelastic Scattering
- ν_{μ} CC QE hyperon production
- Nuclear Cluster Production (deuterons, tritons, helions and alphas)
- Neutrino-electron elastic scattering

 $NC1\pi^01\pi^{\pm}$

• ν_{μ} CC1 π^{\pm}

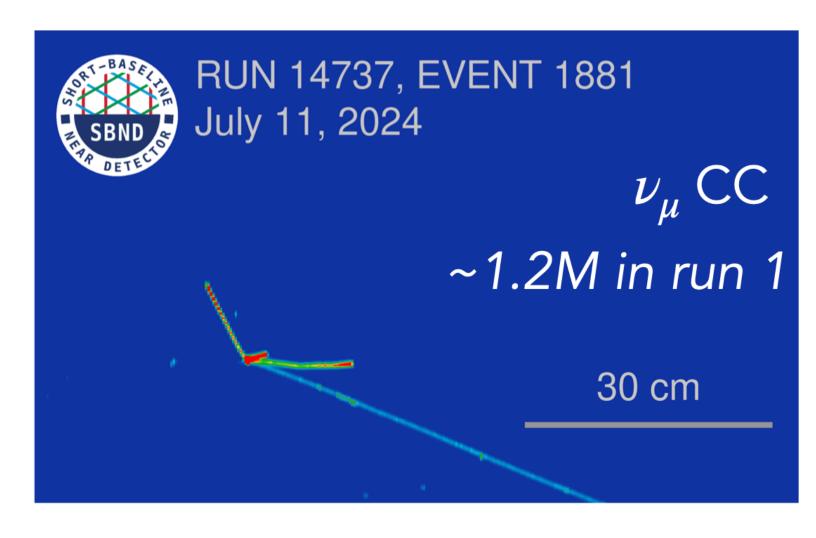
Resonance Production of Eta mesons

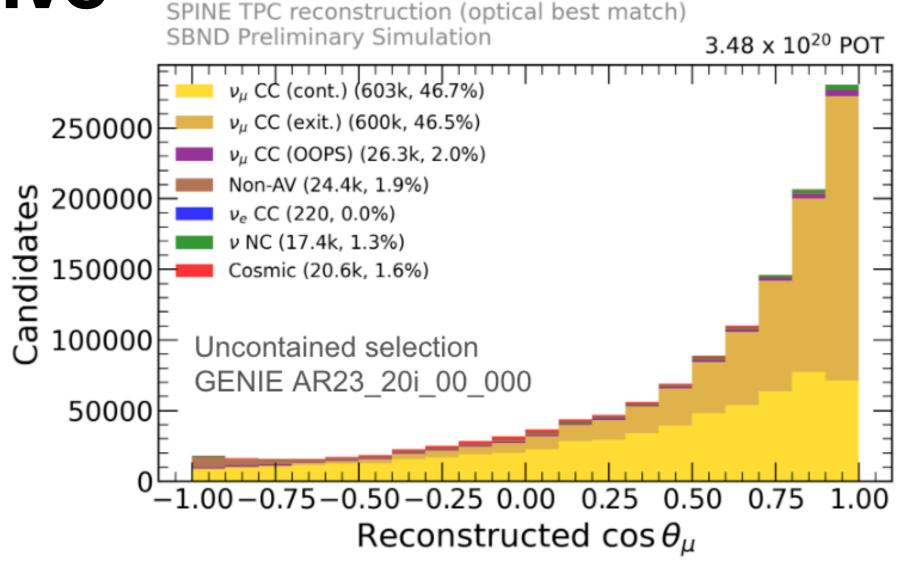


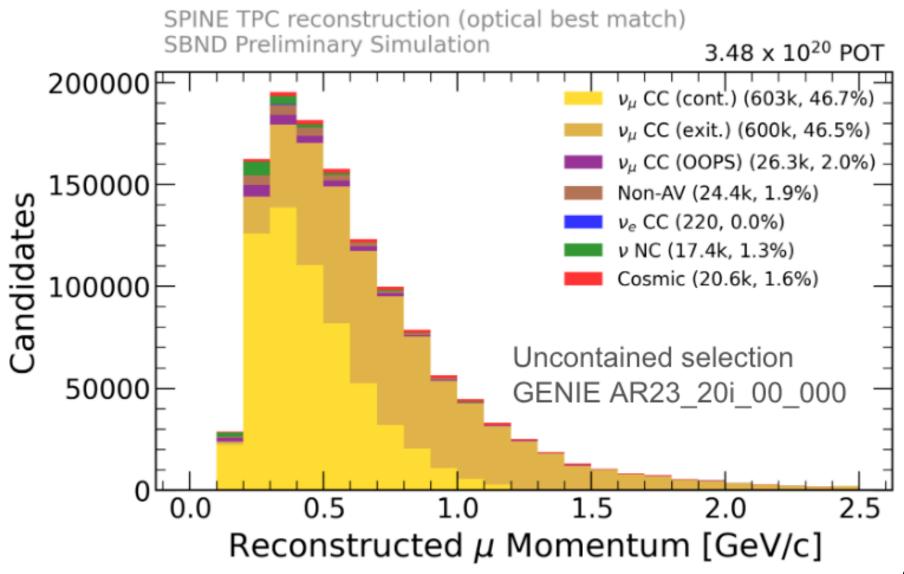
SBND Cross Sections: ν_{μ} CC Inclusive

- Early measurement can establish detector performance
 - Directly compare reconstruction algorithms
 - Excellent cosmic rejection as a surface detector
- Large statistics and low model bias in muon kinematics
- Multi-dimensional differential measurements
- High purity and efficiency

Purity = 93% Efficiency = 79%







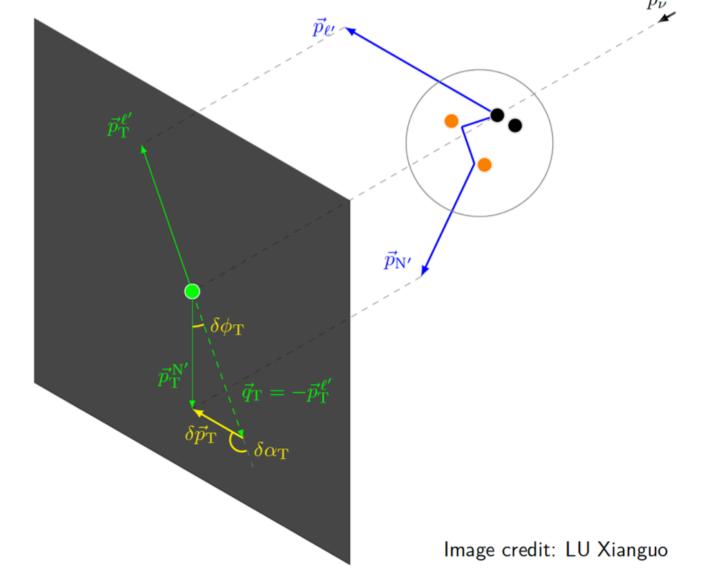


SBND Cross Sections: ν_{μ} CC1p0 π

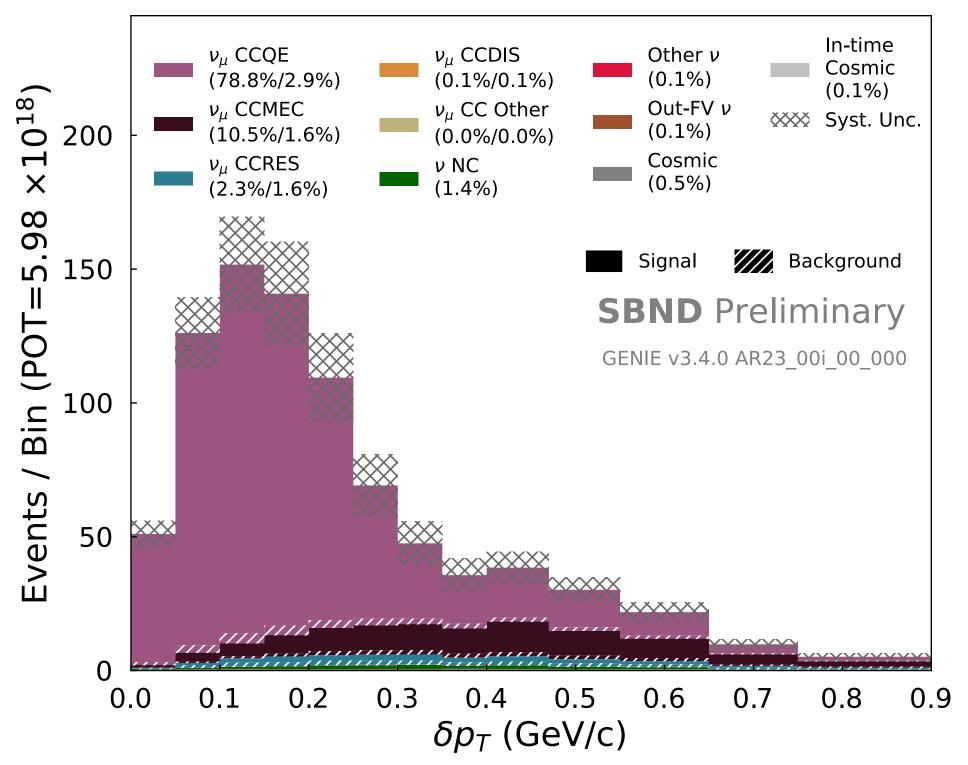
- Targets the quasi-elastic mode and enables observation of low Q^2 effects
- Measurements in muon and proton kinematics as well as Transverse Kinematic Imbalance variables (TKI)

Allows for the study of nuclear effects

Purity = 92.1% Efficiency = 20.3%





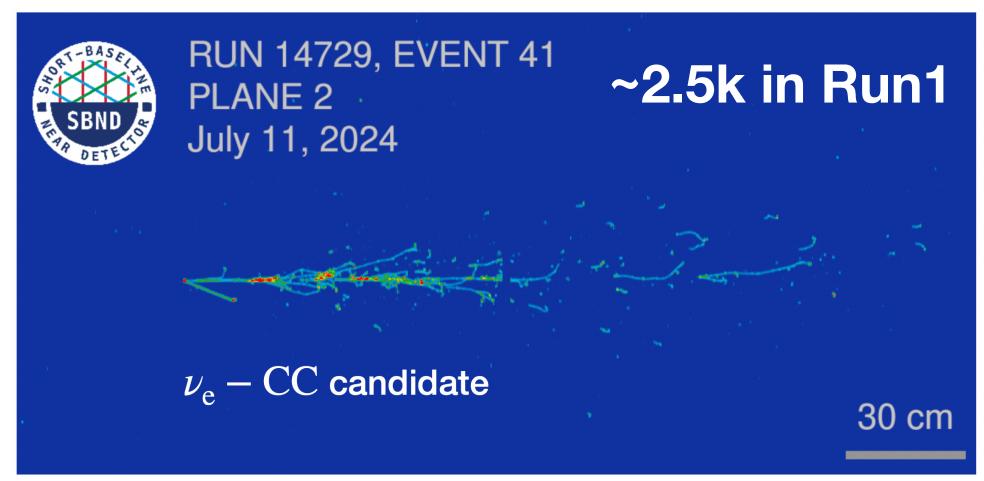


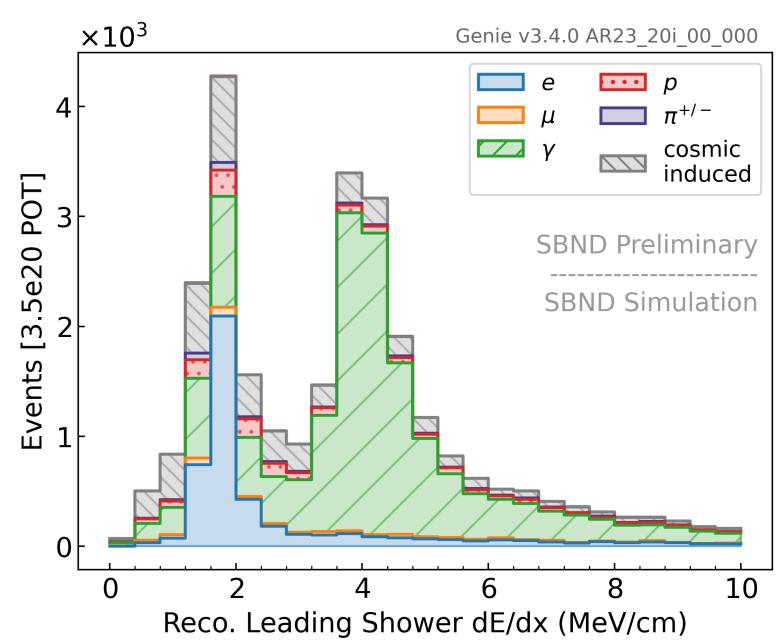


SBND Cross Sections: ν_e CC Inclusive

- Despite small contribution to the BNB flux, still expect ~15,000 events/year
- Important for oscillation physics which uses the LAr-TPCs ability to distinguish between electrons and photons
- Variables such as the shower opening angle, dE/dx and distance between vertex and shower start help to select electrons
- Expect to have ~2.5k ν_e events after selection in Run1 dataset

Purity = 62%, Efficiency = 27%

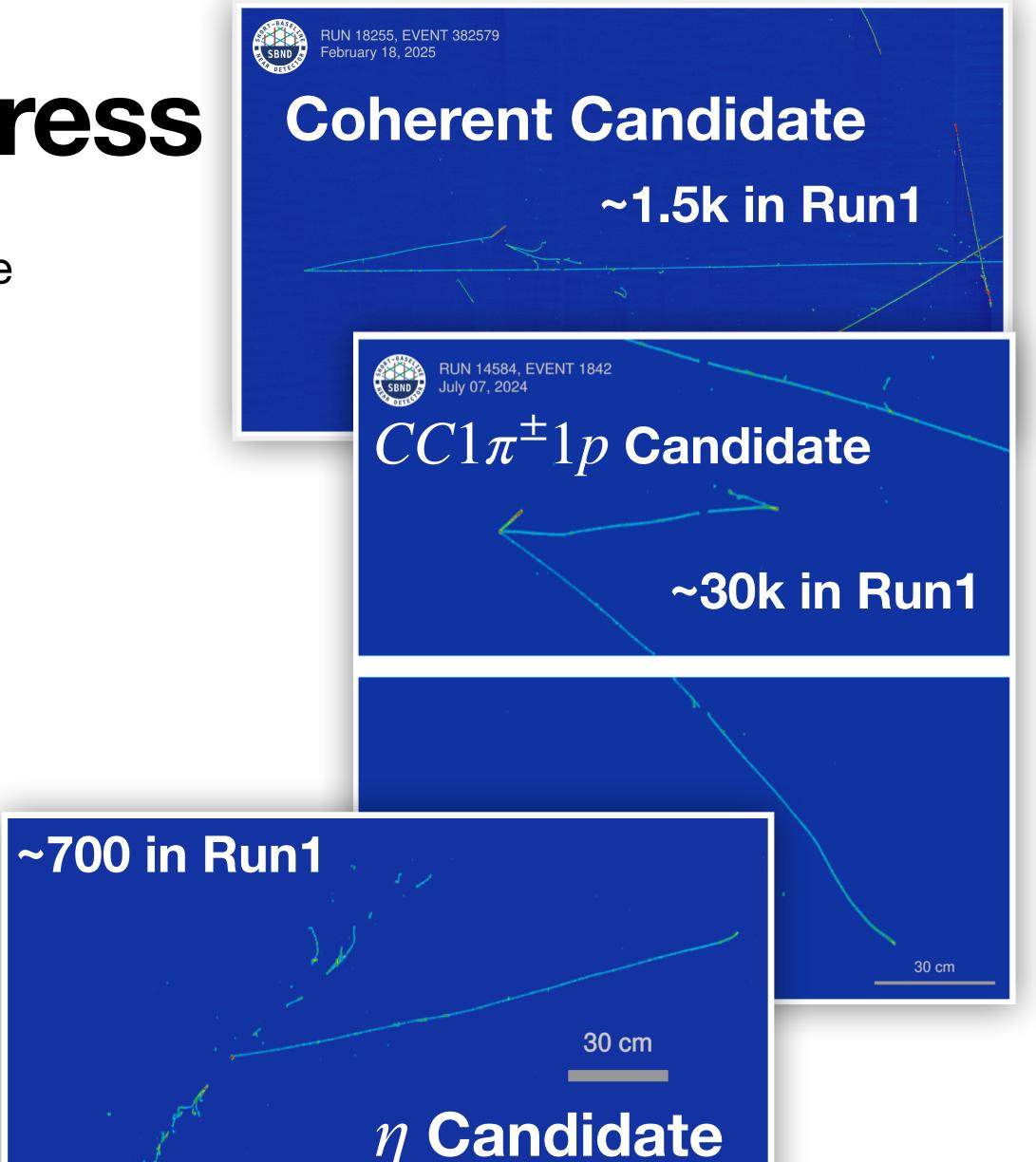






Other Analyses in Progress

- Longer term, SBND will be able to leverage the statistics to probe multi-dimensional cross sections and study rare processes
- Here are some select examples of ongoing analyses:
 - Coherent Pion Production
 - ~800 signal events with 6 months of data
 - ν_{μ} CC $1\pi^{\pm}$
 - Mostly probes resonance events and exudes shower topologies
 - Resonance Production of Eta Mesons
 - Probe resonances beyond the Delta baryon
 - 2 photon decay signal definition
 - Will quickly accumulate world leading statistics for a high precision measurement



RUN 18255, EVENT 300505

30

February 17, 2025



- SBND has been taking physics data since December 2024 and we completed Run 1 in July 2025!
- We already have the world's largest dataset of $\nu-$ Argon interactions
- Detector calibration and reconstruction shows great performance
- Wide physics program including the short-baseline oscillation program, detailed cross sections, rare processes and beyond the standard model searches
- First physics results being shown next week at NuInt 2025!

Thank You!



The SBND Collaboration

updated May 2025

Including both scientific and technical personnel

*Spokespeople

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University of Edinburgh: A. Hamer, L. Kotsiopoulou, M. Nebot-Guinot, H. Parkinson, M. Reggiani-Guzzo, A. Szelc

Federal University of ABC, Brazil: C. Moura

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Instituto Tecnológico de Aeronáutica, Brazil: F. Marinho, L. Paulucci

Fermilab: R. Acciarri, W. Badgett, L. Bagby, V. Basque, M. Betancourt, D. Caratelli, F. Cavanna, O. Dalager, M. Del Tutto, V. Di Benedetto, S. Dixon, S. Dytman, S. Gardiner, M. Geynisman, H. Greenlee, S. M. Kancharla, C. James, T. Junk, W. Ketchum, M.J. Kim, J.Y. Li, L. Liu, P. Machado, M. Micheli, D. Montanari, J. Mueller, T. Nichols, B. Norris, M. Nunes, S. Oh, O. Palamara*, J. Paton, V. Pandey, Z. Pavlovic, D. Pushka, G. Putnam, A. Schukraft, S. Shetty, M. Stancari, A. Stefanik, T. Strauss, D. Torretta, M. Toups, L. Wan, P. Wilson, L. Yates, J. Zennamo

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University of Minnesota: A. Furmanski, N. Pallat Mount Holyoke College: S. Balasubramanian

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Queen Mary University, UK: Y. Dabburi, N. McConkey

Rutgers University: K. Lin, A. Mastbaum

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University of Sussex: R. Darby, C. Griffith, S. Kr Das

Syracuse University: A. Filkins, R. Rajagopalan, M. Soderberg

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Tufts University: O. Alterkait, Z. Imani, T. Wongjirad

Virginia Tech: C. Mariani, G. Moreno-Granados, P. Roy

University of Warwick: A. Chappell, X. Lu, J. Marshall, A. Wilkinson

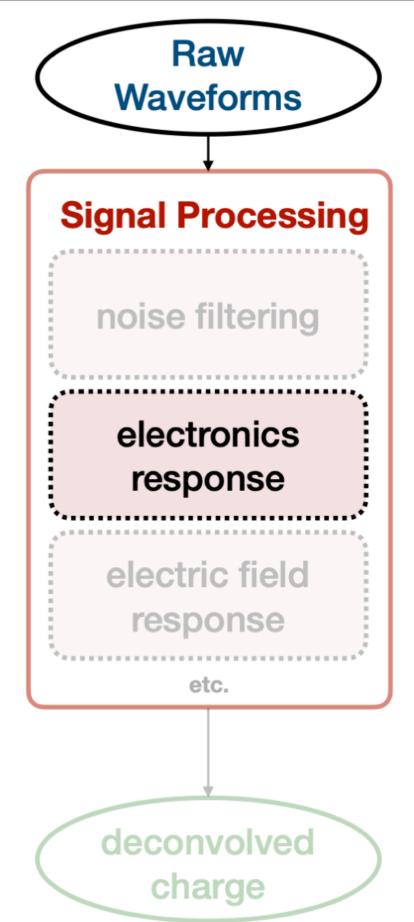


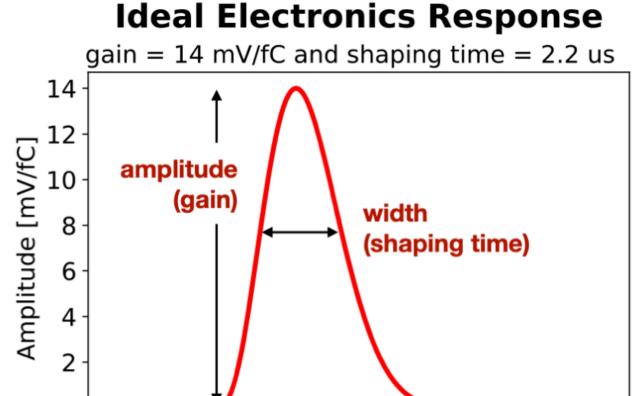
Backup

33

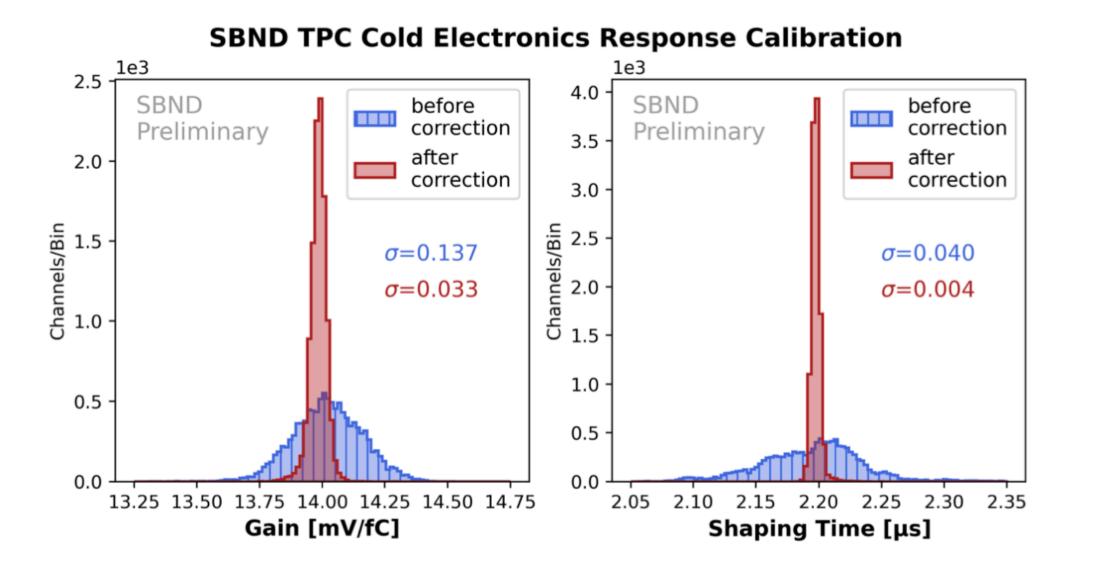


Signal Processing





time [μ s]



- Cold electronics are responsible for amplifying and shaping signals for each wire deconvolved charge
- Clear improvement in response uniformity after corrections from internal calibration source

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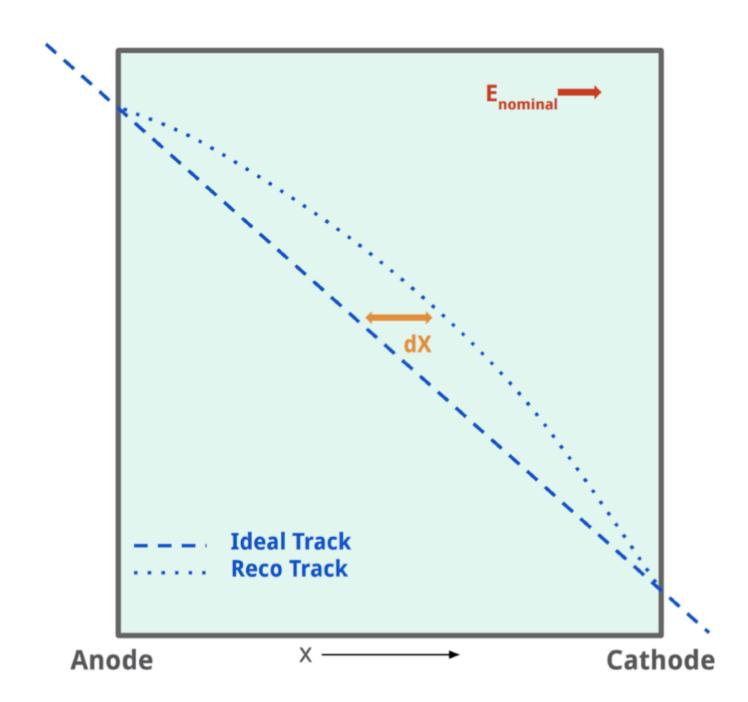
12

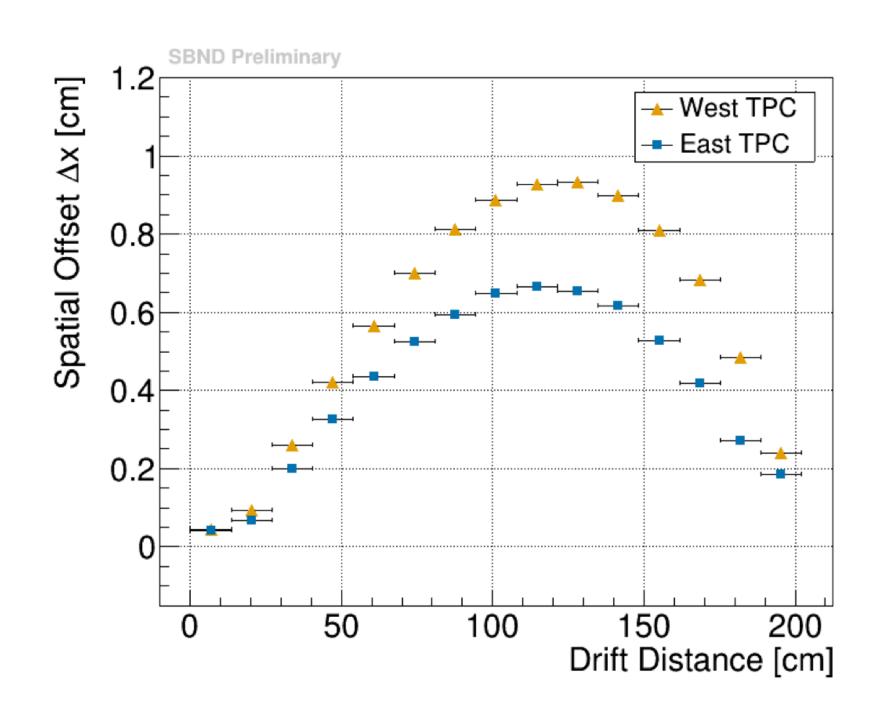
10



Space Charge Effect

- \bullet One of the first steps in calibrating our TPC is characterizing the electric field, including distortions that come from accumulated Ar^+ ions, which is called the space charge effect
- The offsets due to space charge are on average <1cm across the detector, reflecting electric field distortions up to 6% of the nominal field

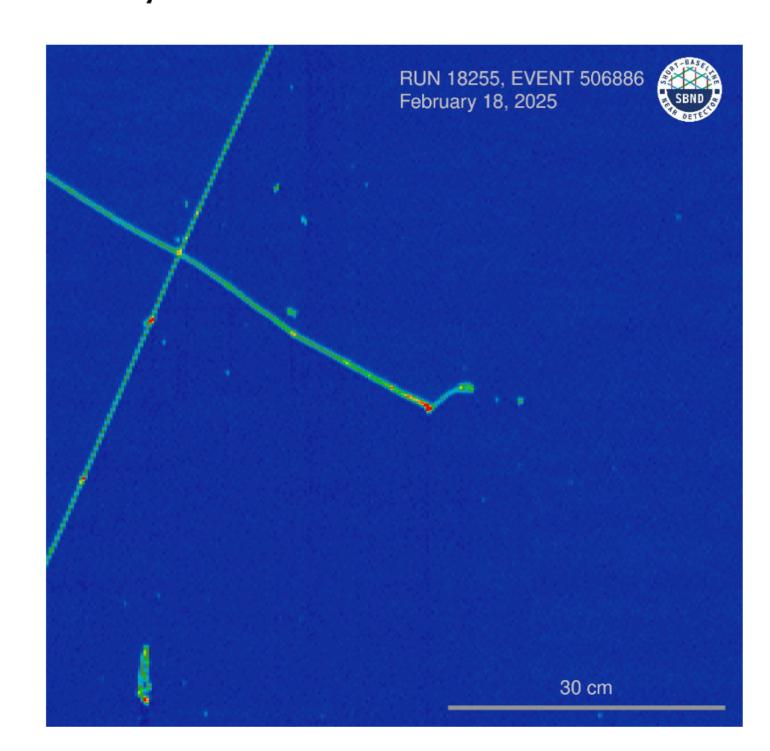


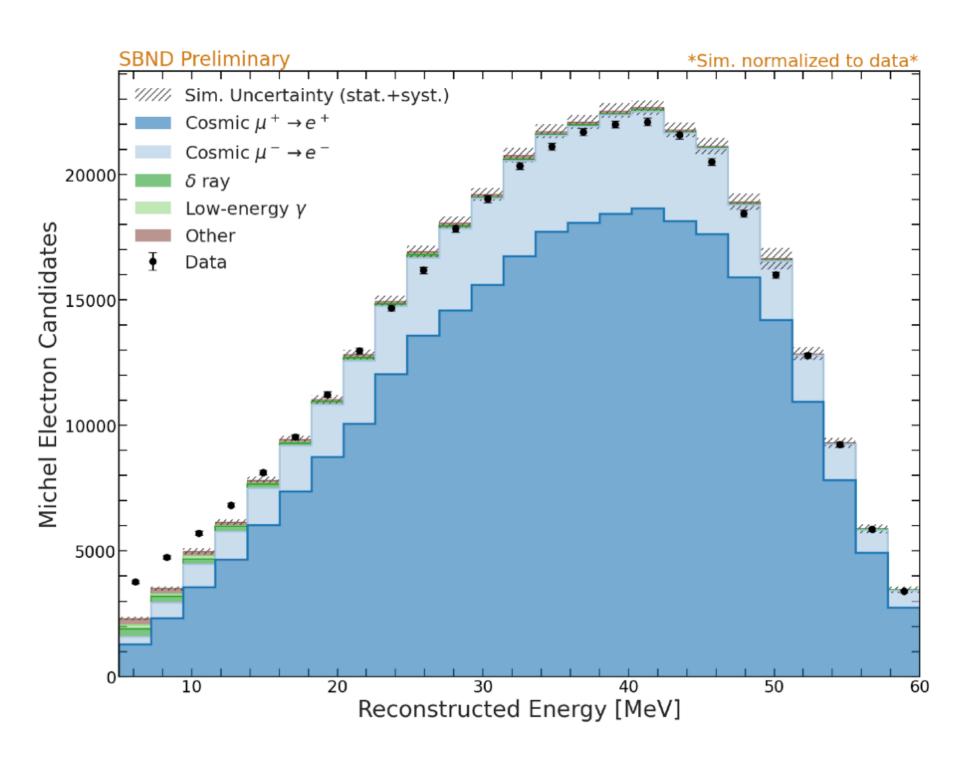




Michelle Spectrum

- Another valuable standard candle is the low-energy Michel electrons produced when muons decay at rest in the detector, which have a well-understood spectrum
- The Michel spectrum shows good agreement with simulation, indicating that we are able to accurately reconstruct these low-energy electrons in data







Flash Match (best match, whole detector) + SPINE (ML) TPC reconstruction SBND Preliminary Simulation

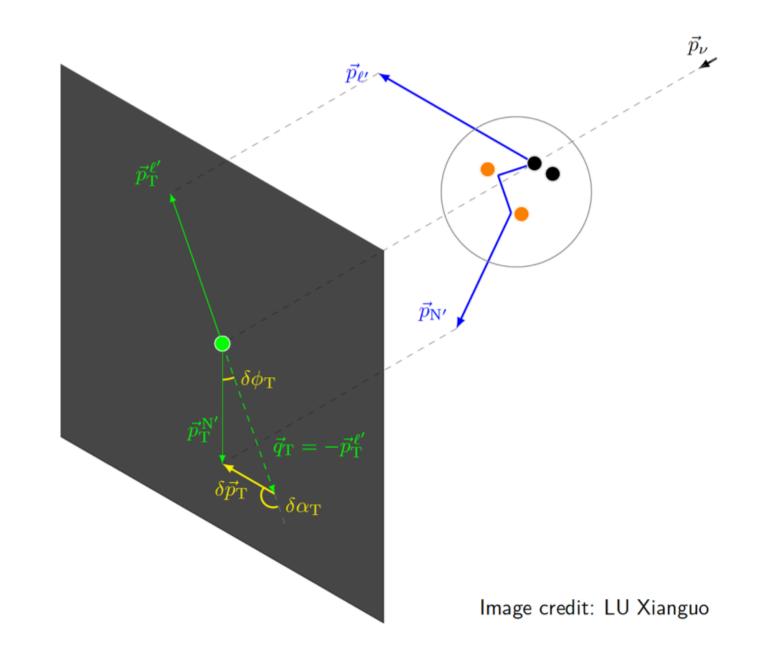
Cut	Purity [%]	Efficiency [%]	Differential Efficiency [%]	Differential Purity [%]
No cut	3.05	100.0	n/a	n/a
Flash Match (best match, whole detector)	26.29	94.4	94.4	23.25
Fiducial	53.01	91.44	97.04	26.72
Has Muon	77.98	83.59	92.15	24.96
Start dE/dx	86.74	80.23	96.64	8.76
Flash Score	93.12	79.45	99.23	6.38

FIG. 7: Purity and efficiency after each cut of the uncontained selection.



v_{μ} CC 1p0 π Measurement

- Final state kinematics
 - One muon with $P_{\mu} \in [220, 1000]$ MeV/c
 - One proton with $P_p \in [300, 1000]$ MeV/c
 - No charged pion with $P_{\pi} > 70$ MeV/c
 - No neutral pions
 - Other particles not listed are allowed
- Measurement kinematic variables
 - μ and p momentum and direction relative to the beam
 - TKI variables: δp_T , $\delta \alpha_T$, $\delta \phi_T$
- Events required to be fully contained within the detector volume for this first analysis



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v_{μ} CC 1p0 π Event Selection

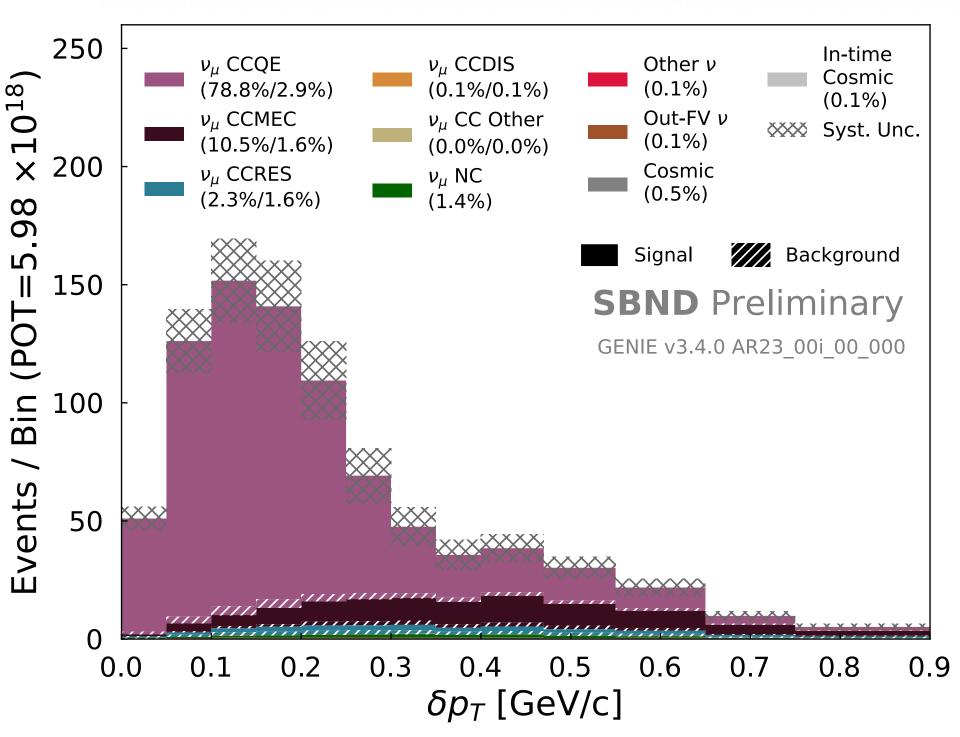
- TPC-only event selection
 - Cosmic rejection
 - Select neutrino-like events
 - SBND benefits from lower cosmic background rates than other SBN detectors
 - Efficiency: 78.7%, signal purity: 20.7%, CCQE purity: 33.2%
 - 2-track topology selection
 - Select clean events with only two particle trajectories emerging from a common vertex
 - Efficiency: 26.3%, signal purity: 58.9%, CCQE purity: 62.7%
 - Muon-proton candidate selection
 - Identify events where the two tracks are classified as muon-like and proton-like using calorimetry-based PID scores
 - Efficiency: 20.3%, signal purity: 92.1%, CCQE purity: 81.7%



SBND Cross Sections: ν_{μ} CC 1p0 π

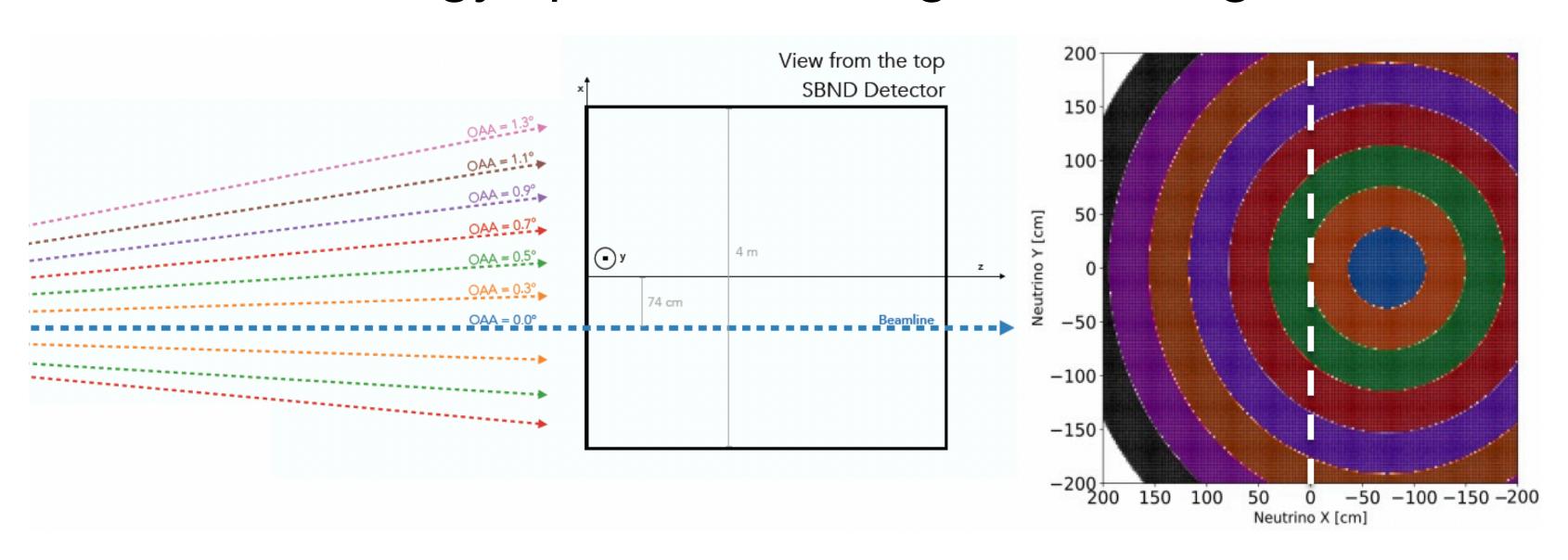
- Targets the quasi-elastic mode and enables observation of low Q^2 effects
- Measurements in muon and proton kinematics as well as Transverse Kinematic Imbalance variables (TKI)
- Allows for the study of nuclear effects
- Estimated 2 million signal events after 3 years of data taking
- 88 % purity at 38% efficiency

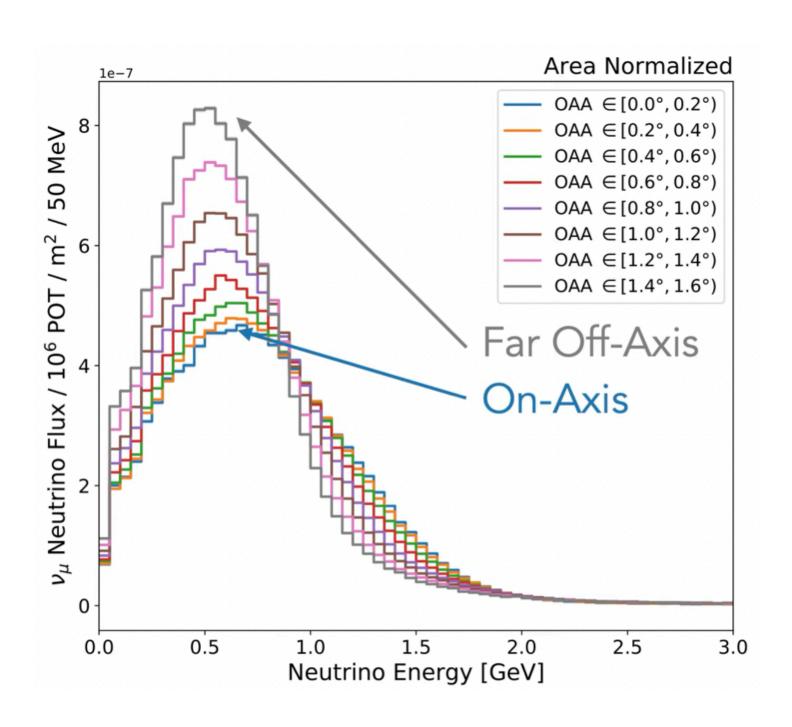






- Precision Reaction Independent Spectrum Measurement: https://arxiv.org/abs/2508.20239
- SBND sits 74 cm off-axis and its close proximity to the beam allows it to probe an angular range of [0, 1.6] degrees
- Neutrino energy spectrum changes with angle

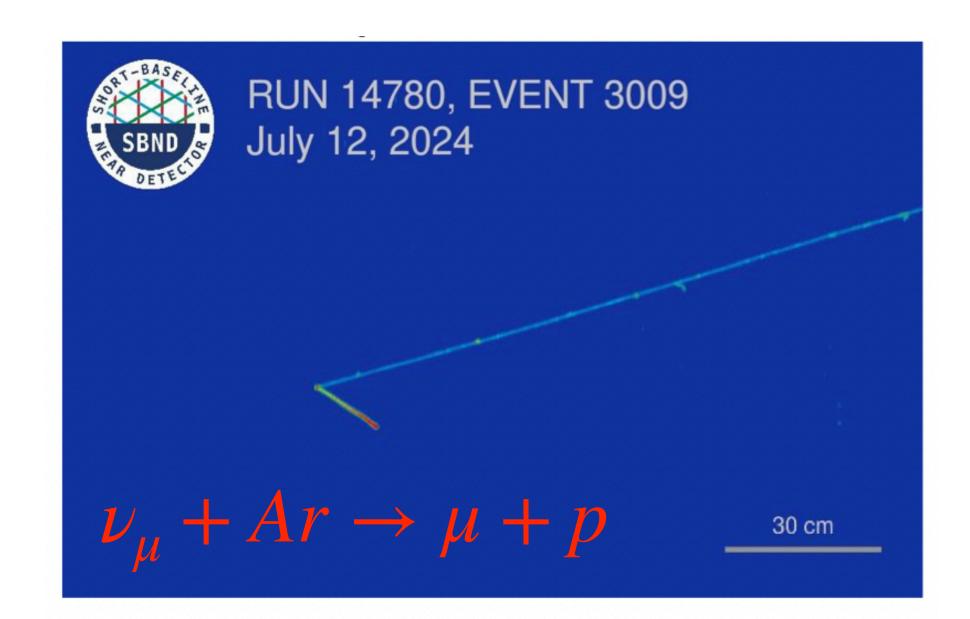


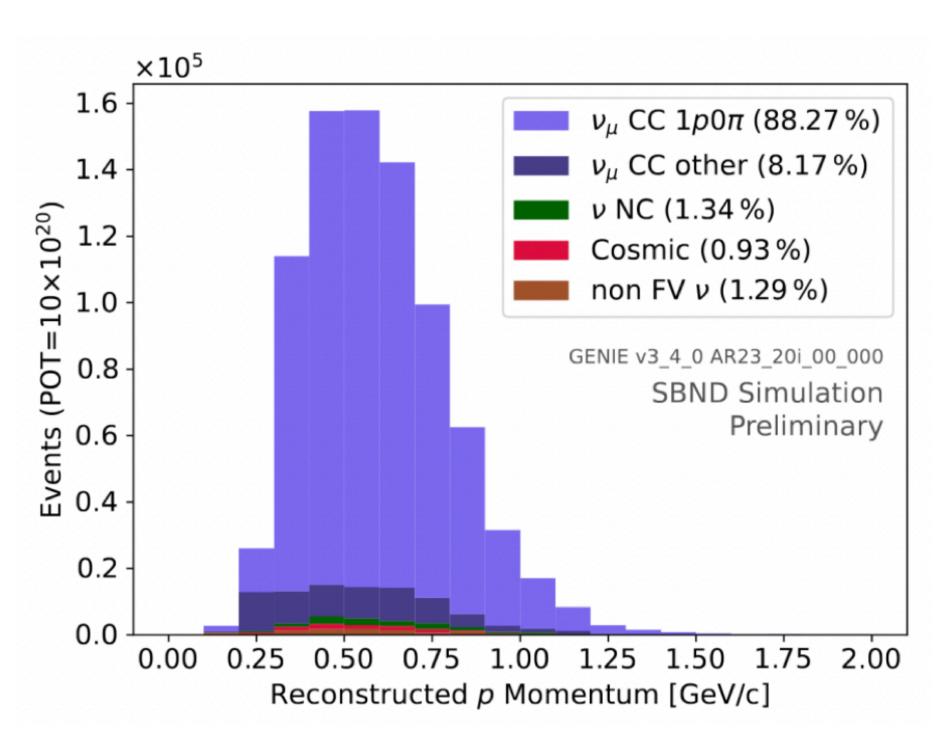




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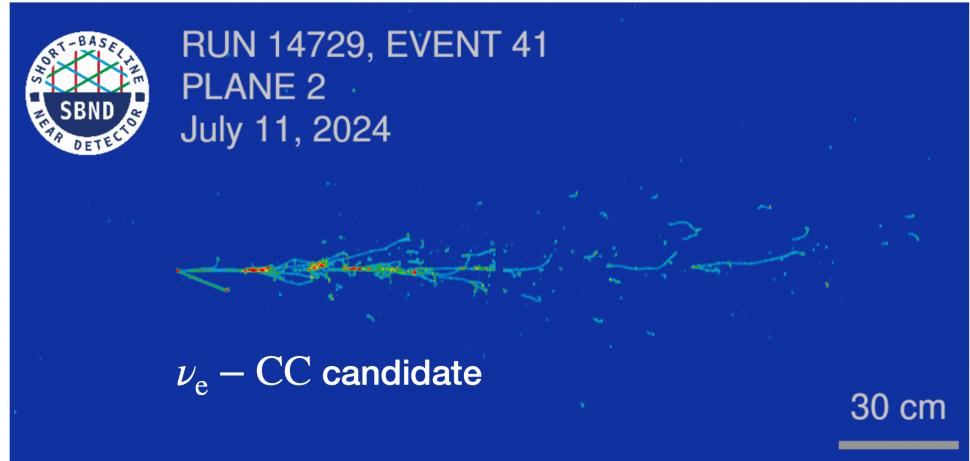


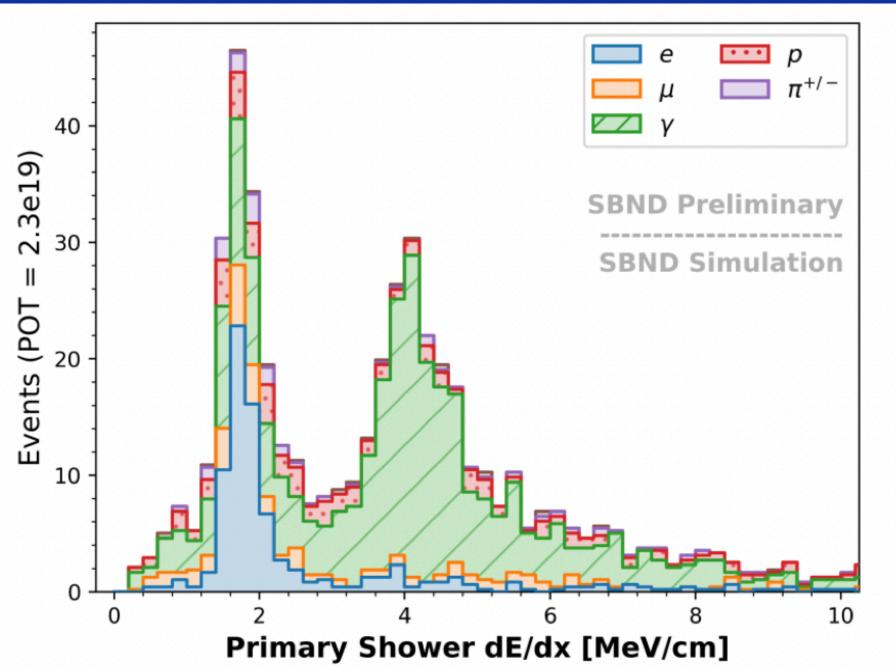




SBND Cross Sections: ν_{e} CC Inclusive

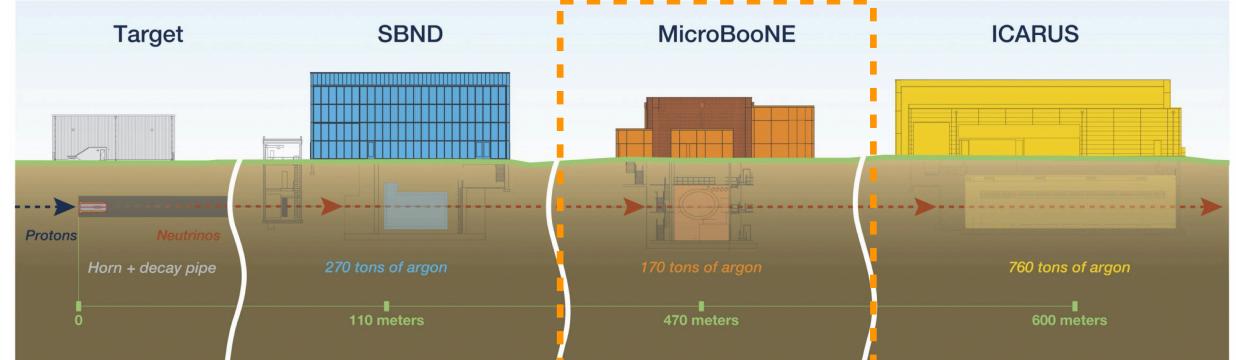
- Despite small contribution to the BNB flux, still expect ~15,000 events/year
- Important for oscillation physics which uses the LAr-TPCs ability to distinguish between electrons and photons
- Variables such as the shower opening angle, dE/dx and distance between vertex and shower start help to select electrons
- 78% purity and 31% efficiency for $E_e > 500\,$ MeV

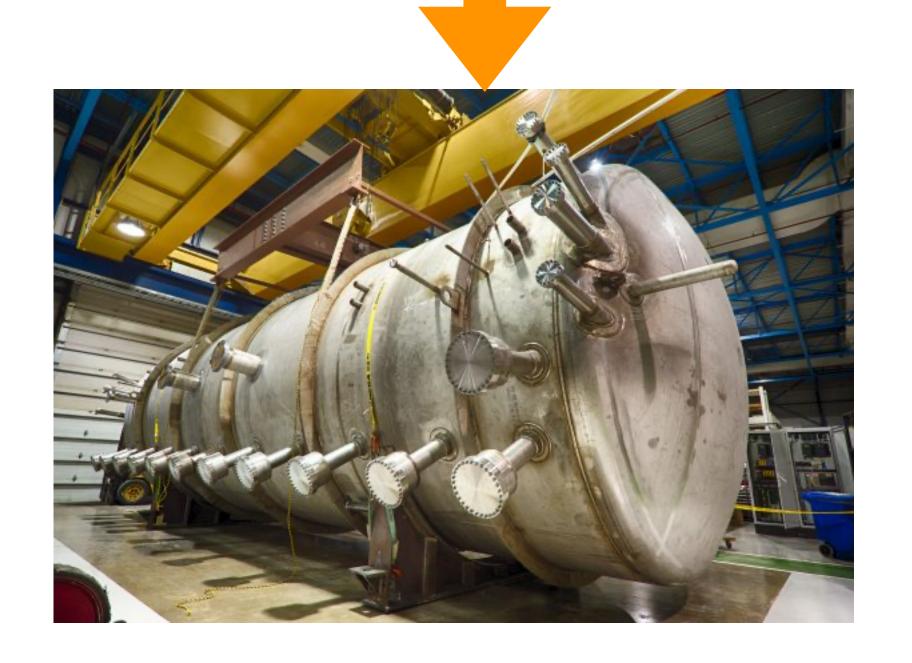






- LAr-TPCs provide calorimetry and 3D tracking with fine-grained information
 - Capable of identifying different species of particles
 - Particle Flow: parent vs daughter particles
 - Track vs Shower, Interaction Vertex, etc.
 - Generational Knowledge
 - LAr-TPCs are a well-understood technology and we benefit from the experience of many past experiments (ICARUS, ArgoNeuT, MicroBooNE ...)

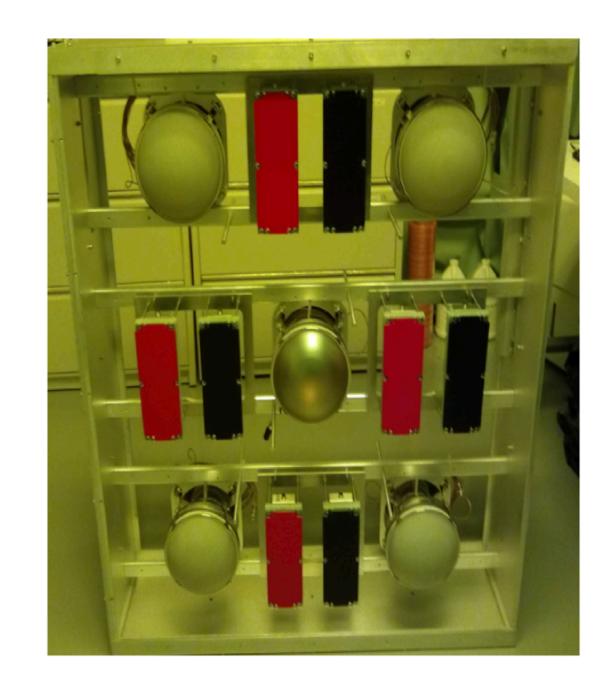




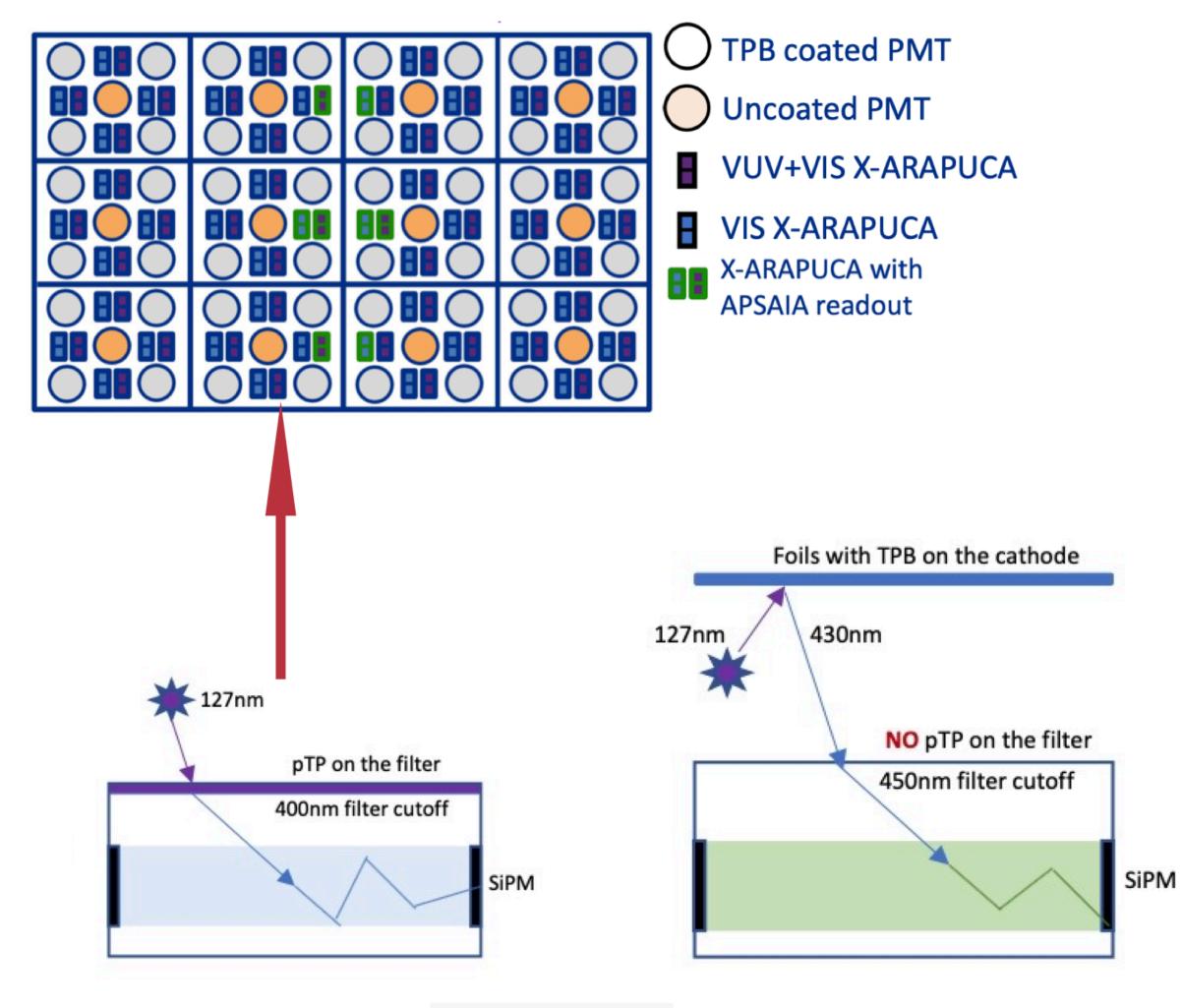


SBND Photon Detection System (PDS)

Modular system behind the APAs: "looking" inside the TPC with 24 modules (12 per side)



Each module has 5 8" PMTs (120 total) and 8 X-ARAPUCAS (192 total).



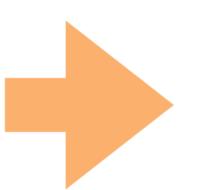
X-ARAPUCA's



Alternative BSM Stuff

Beyond the Standard Model searches at SBND

- High-intensity proton beams (high-intensity neutrino beams)
- SBND's proximity to the target



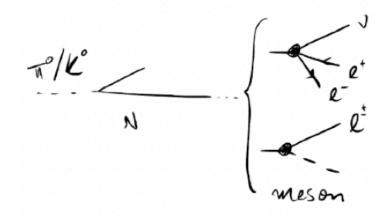
- Access to various BSM scenarios as alternative to MiniBooNE's anomaly.
- Probe signatures for new physics scenarios in the neutrino sector and beyond.

Dark Neutrinos



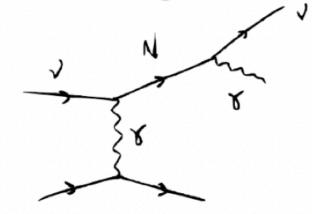
Bertuzzo Jana Machado Zukanovich PRL 2018, PLB 2019
Arguelles Hostert Tsai PRL 2019
Ballett Pascoli Ross-Lonergan PRD 2019
Ballett Hostert Pascoli PRD 2020

Heavy Neutral Leptons



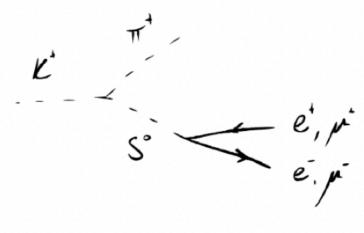
Ballett Pascoli Ross-Lonergan JHEP 2017 Kelly Machado PRD 2021

Transition Magnetic Moment



Gninenko PRL 2009 Coloma Machado Soler Shoemaker PRL 2017 Atkinson et al 2021 Vergani et al 2021

Higgs Portal Scalar



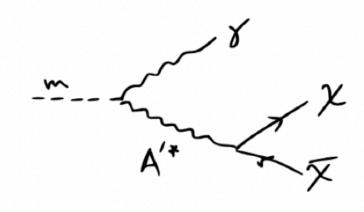
Pat Wilczek 2006 Batell Berger Ismail PRD 2019 MicroBooNE 2021

Axion-like Particles



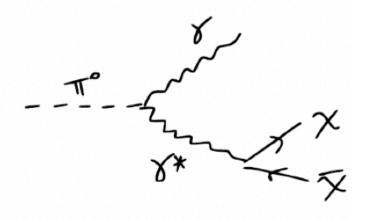
Kelly Kumar Liu PRD 2021 Brdar et al PRL 2021

Light Dark Matter



Romeri Kelley Machado PRD 2019

Millicharged Particles



Magill, Plestid, Pospelov, Tsai, PRL 2019 Harnik Liu Palamara, JHEP 2019

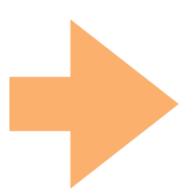
Diagrams: P. Machado Slide M. del Tutto, R. Jones



Alternative BSM Stuff

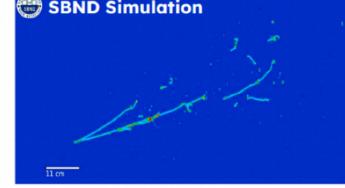
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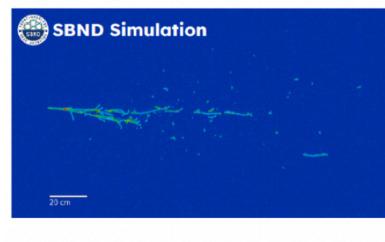
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Dark Neutrinos SBND Simulation



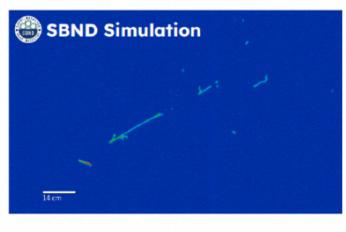
Bertuzzo Jana Machado Zukanovich PRL 2018, PLB 2019 Arguelles Hostert Tsai PRL 2019 Ballett Pascoli Ross-Lonergan PRD 2019 Ballett Hostert Pascoli PRD 2020

Heavy Neutral Leptons



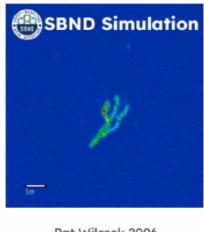
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Transition Magnetic Moment



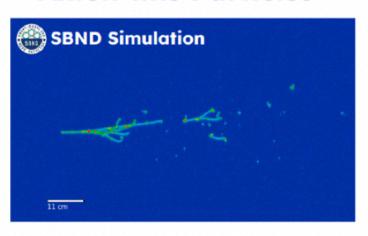
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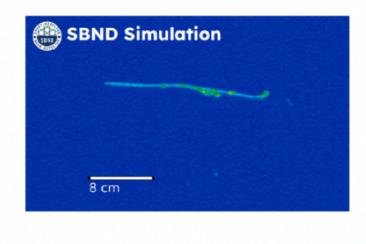
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Axion-like Particles



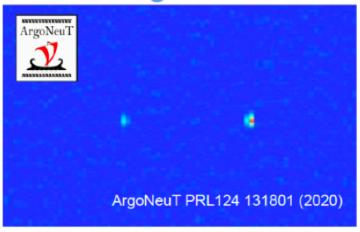
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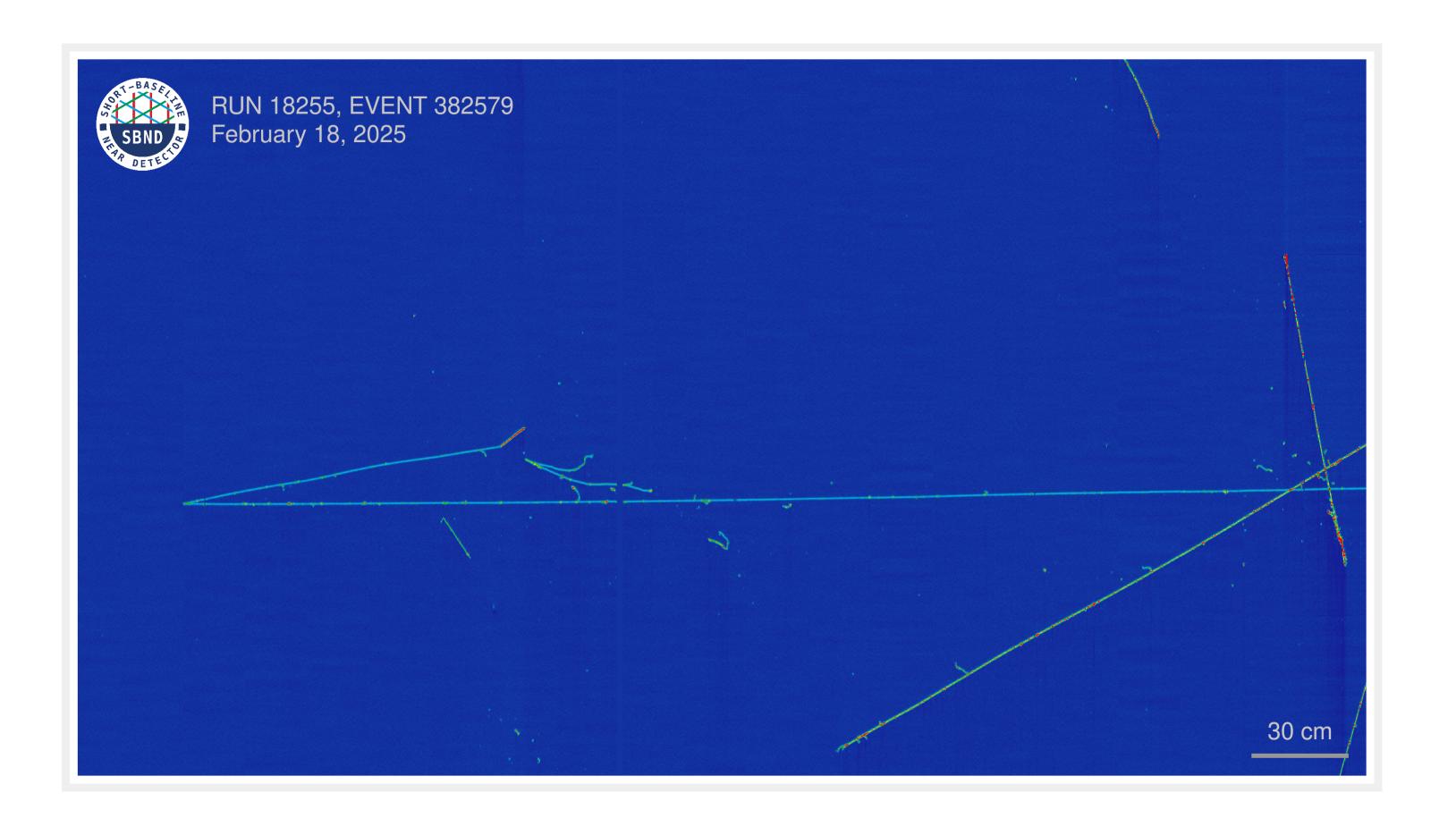


Magill, Plestid, Pospelov, Tsai, PRL 2019 Harnik Liu Palamara, JHEP 2019

> Diagrams: P. Machado Slide M. del Tutto, R. Jones



Coherent Candidate





Eta Meson Display

