



Nucleosynthesis in classical novae: nuclear physics and astrophysics uncertainties

Spencer Reyno
Saint Mary's University

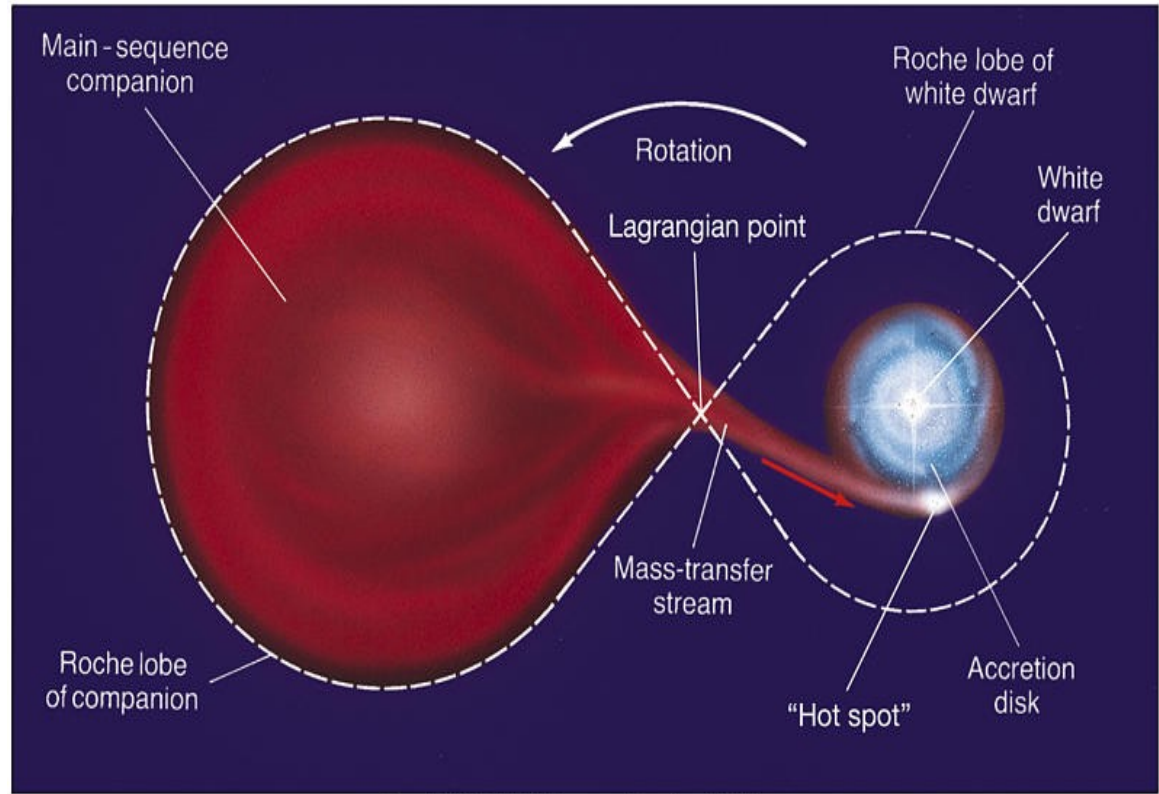
Canadian Astroparticle Summer Student Talk Symposium
August 12th, 2026



**Saint Mary's
University**
SCIENCE

Classical Novae

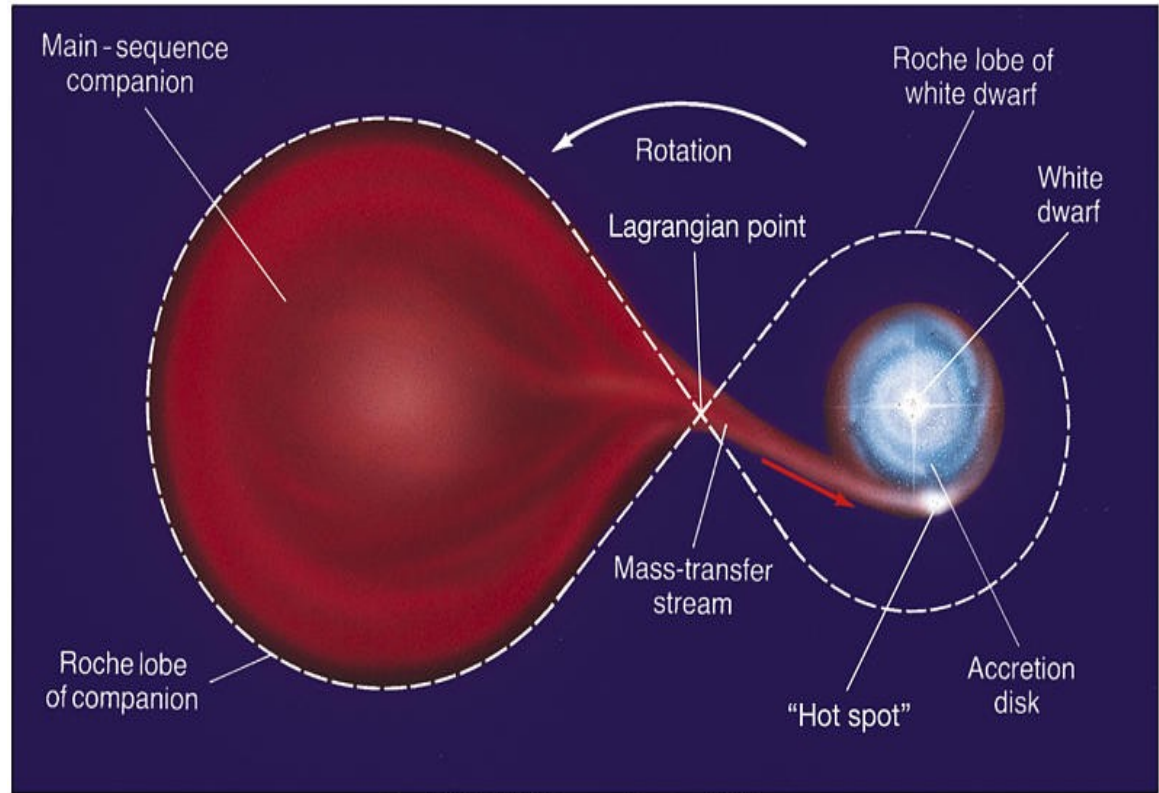
Classical Novae



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Classical Novae

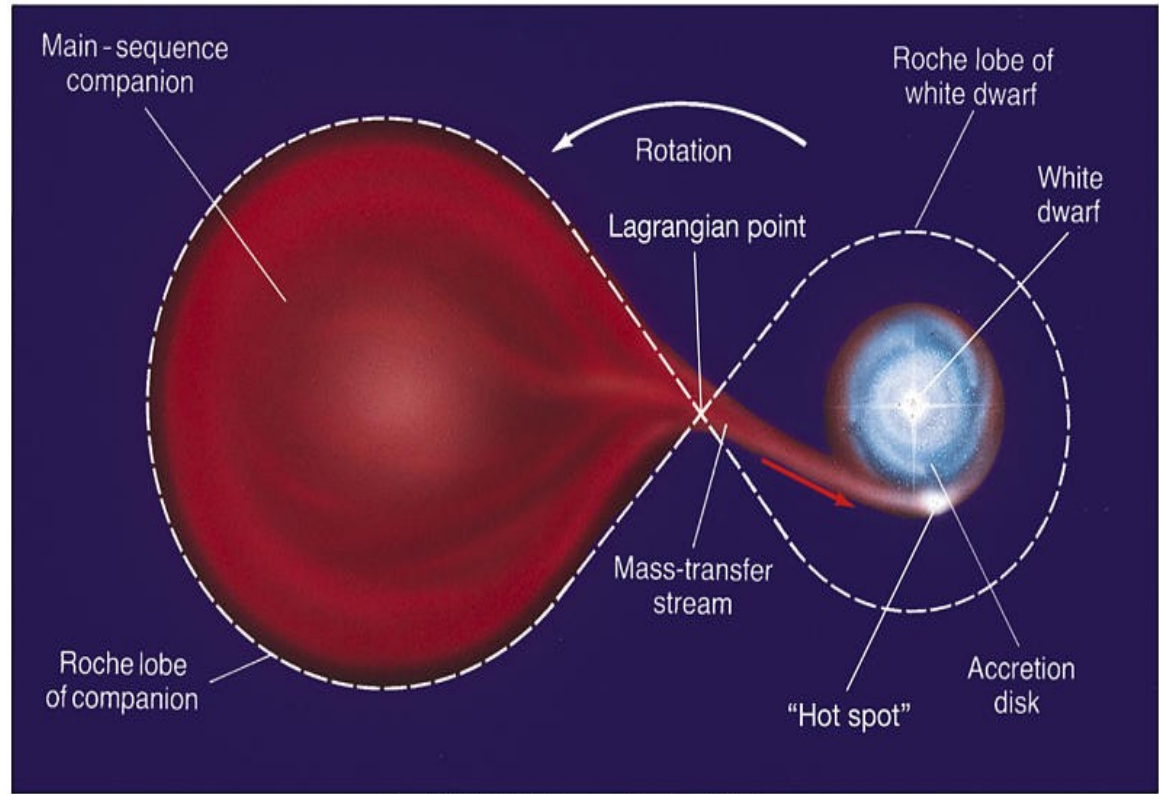
- White dwarf and non-degenerate companion



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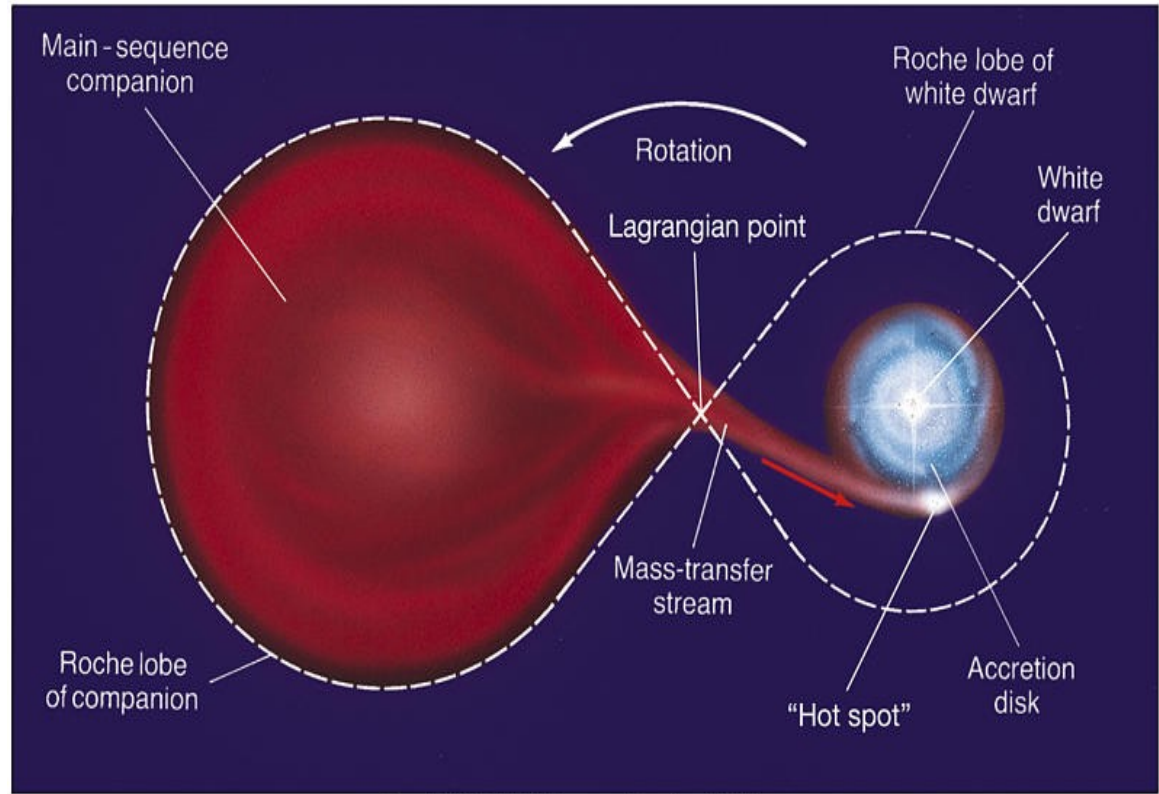
- White dwarf and non-degenerate companion
- Surface material from non-degenerate companion accretes onto white dwarf



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Classical Novae

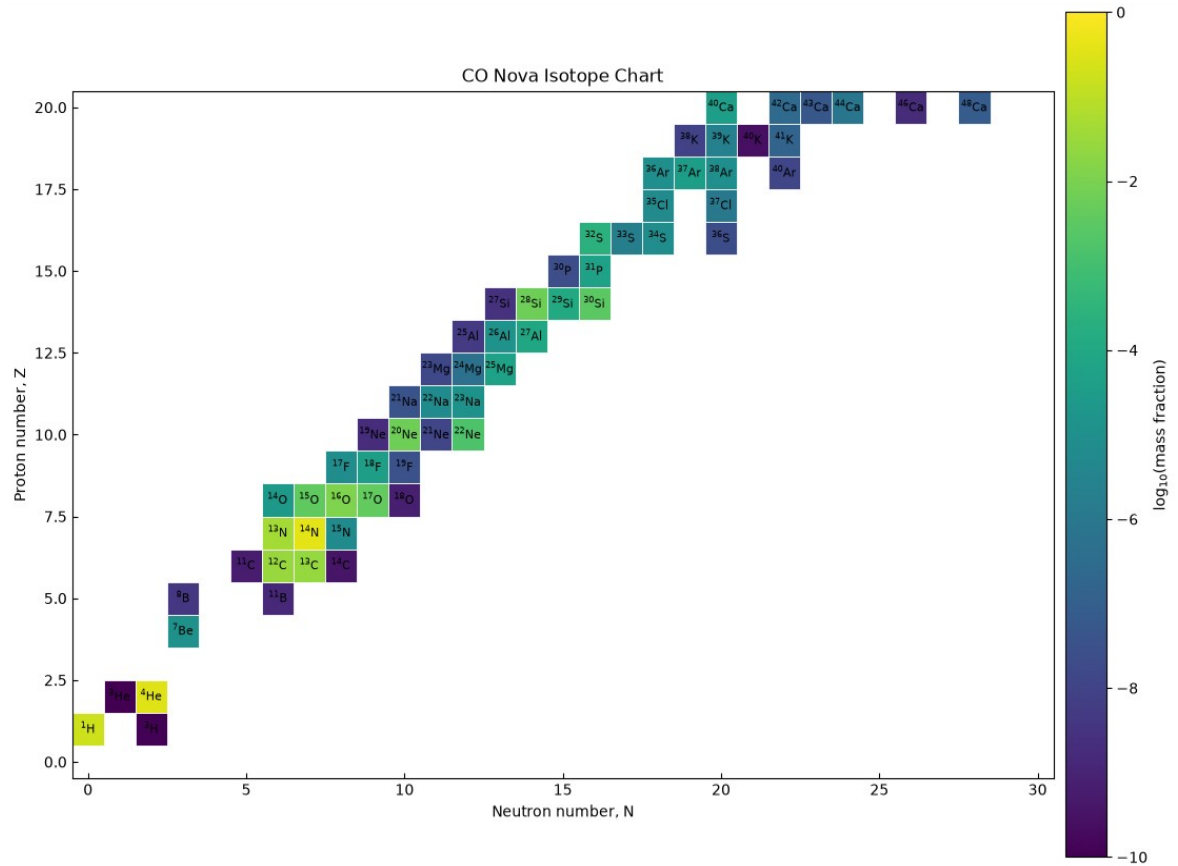
- White dwarf and non-degenerate companion
- Surface material from non-degenerate companion accretes onto white dwarf
- Thermonuclear runaway (TNR) occurs, and matter is expelled



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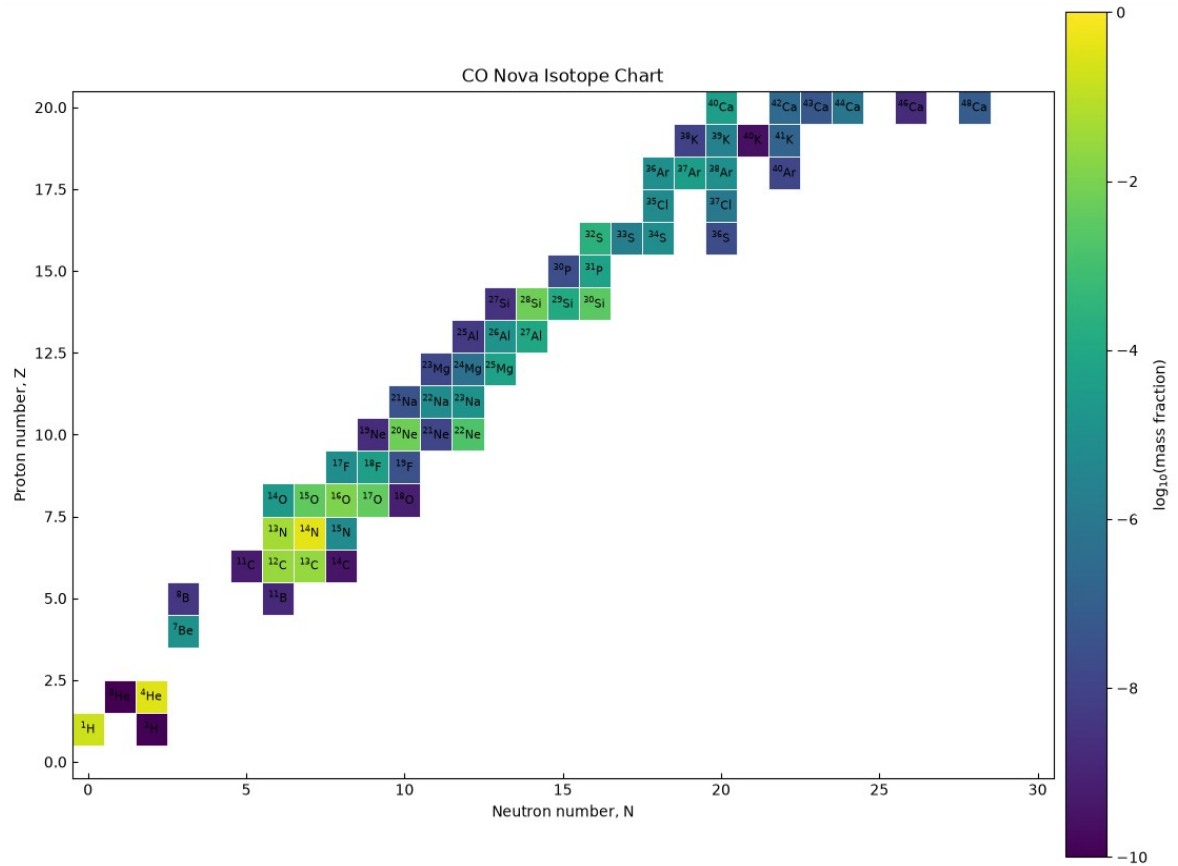
Classical Novae

- Nuclear reaction network present



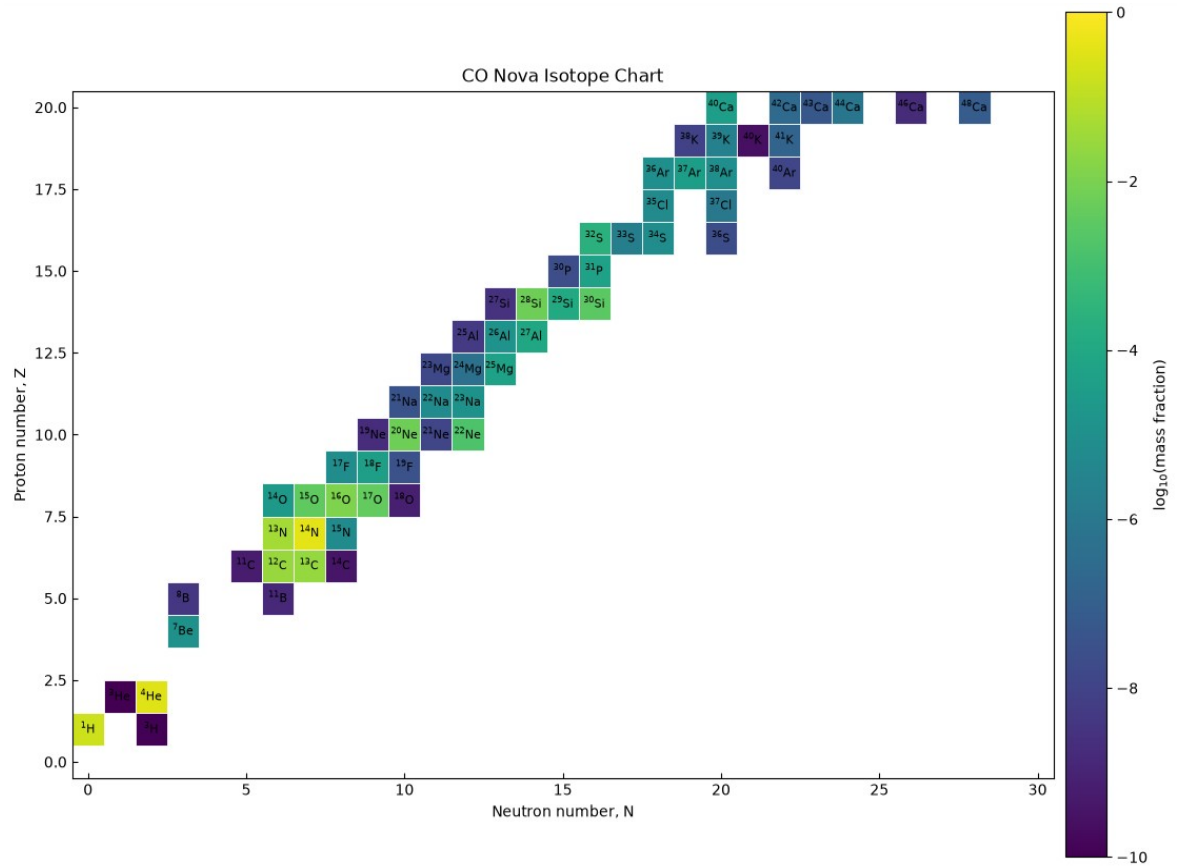
Classical Novae

- Nuclear reaction network present
- Isotopes up to $Z=20$ (Ca) produced



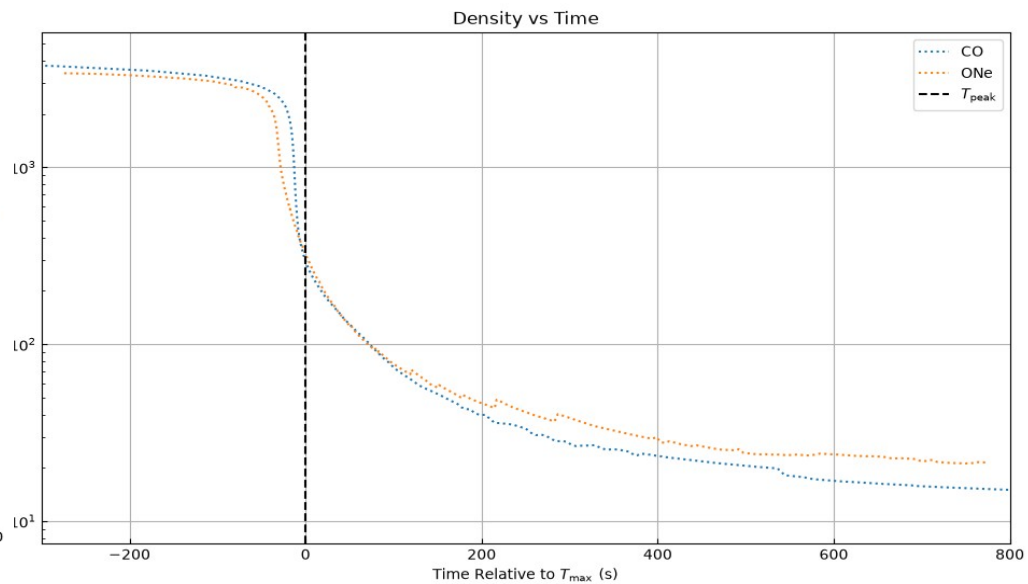
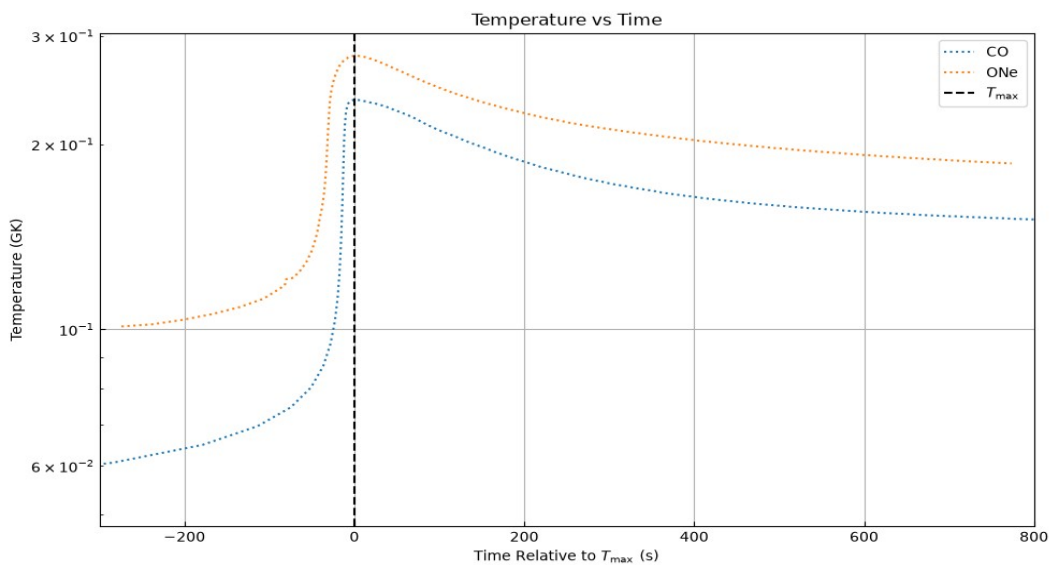
Classical Novae

- Nuclear reaction network present
- Isotopes up to $Z=20$ (Ca) produced
- Rare radioactive isotopes included

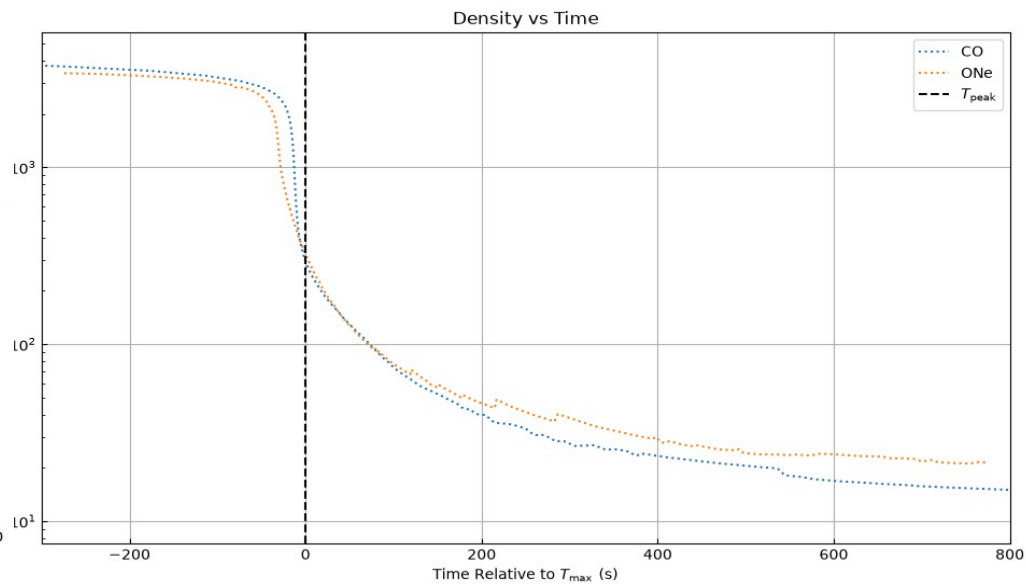
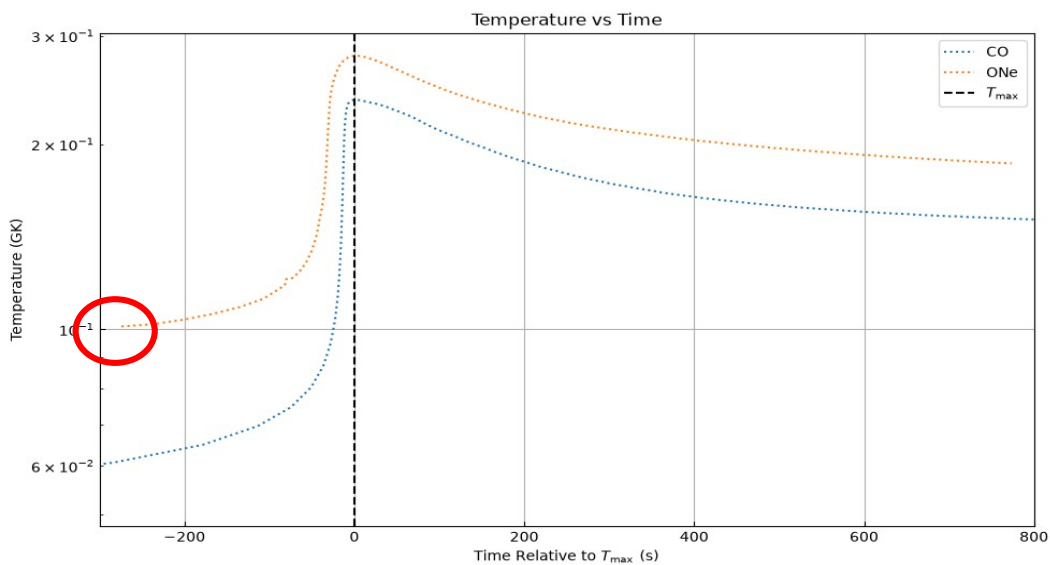


CO vs. ONe Novae

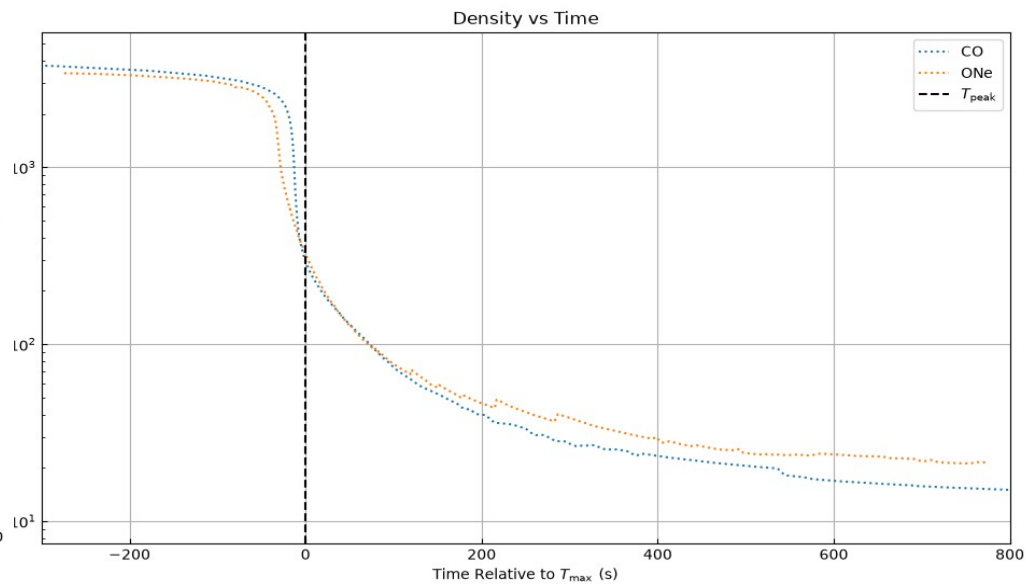
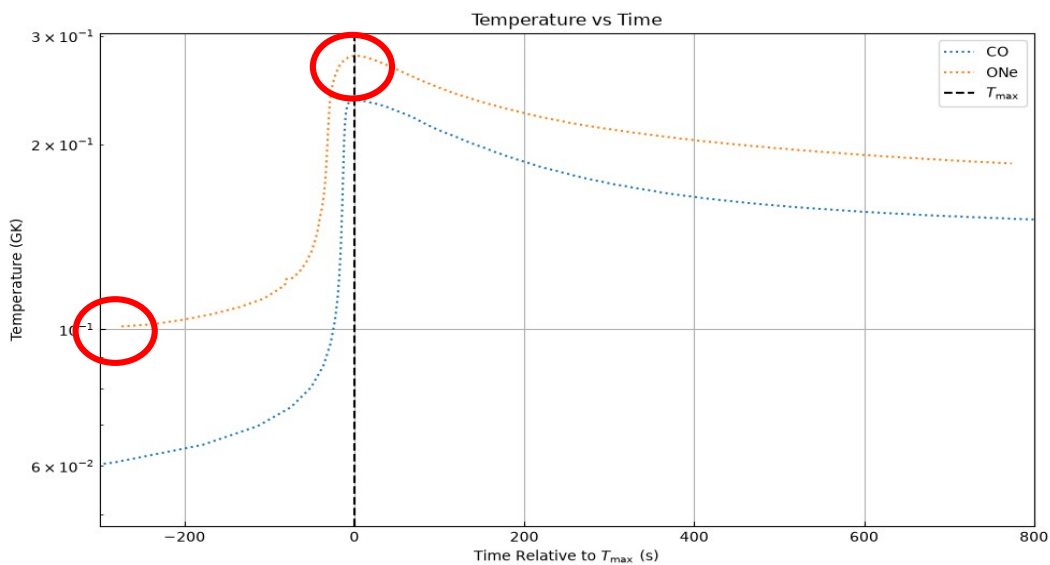
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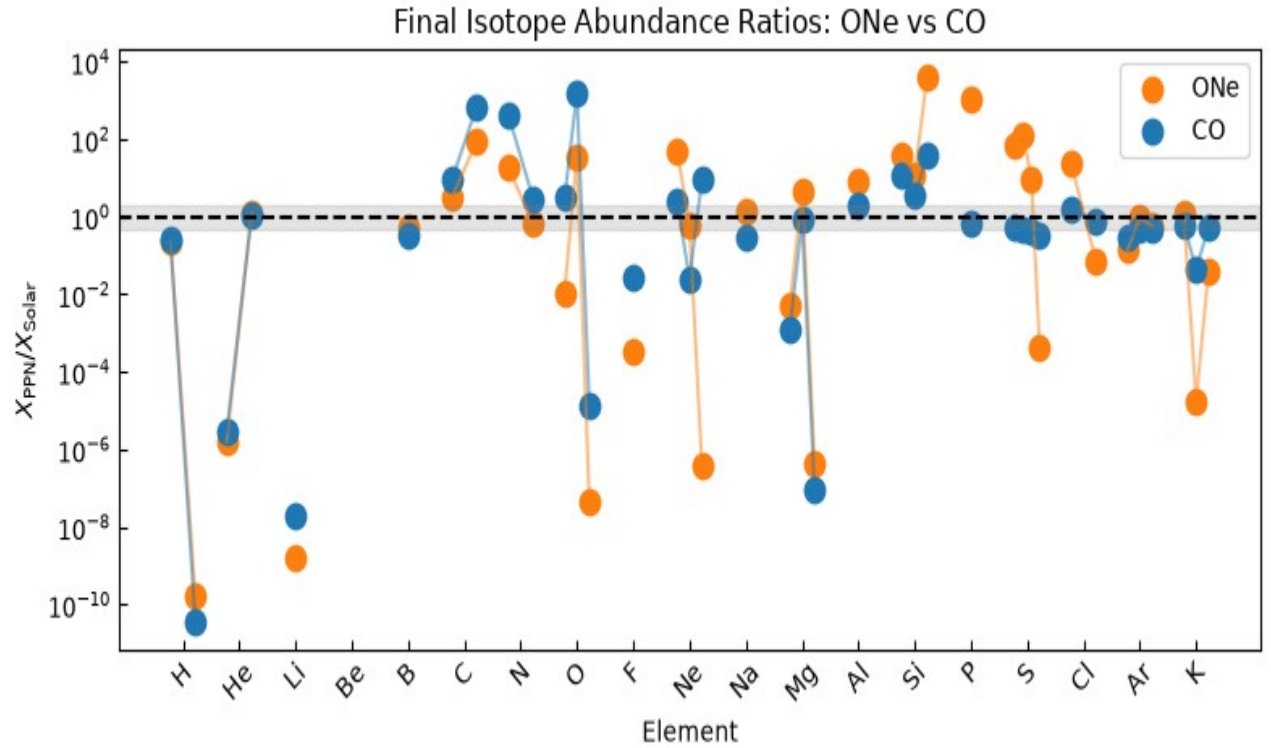


CO vs. ONe Novae



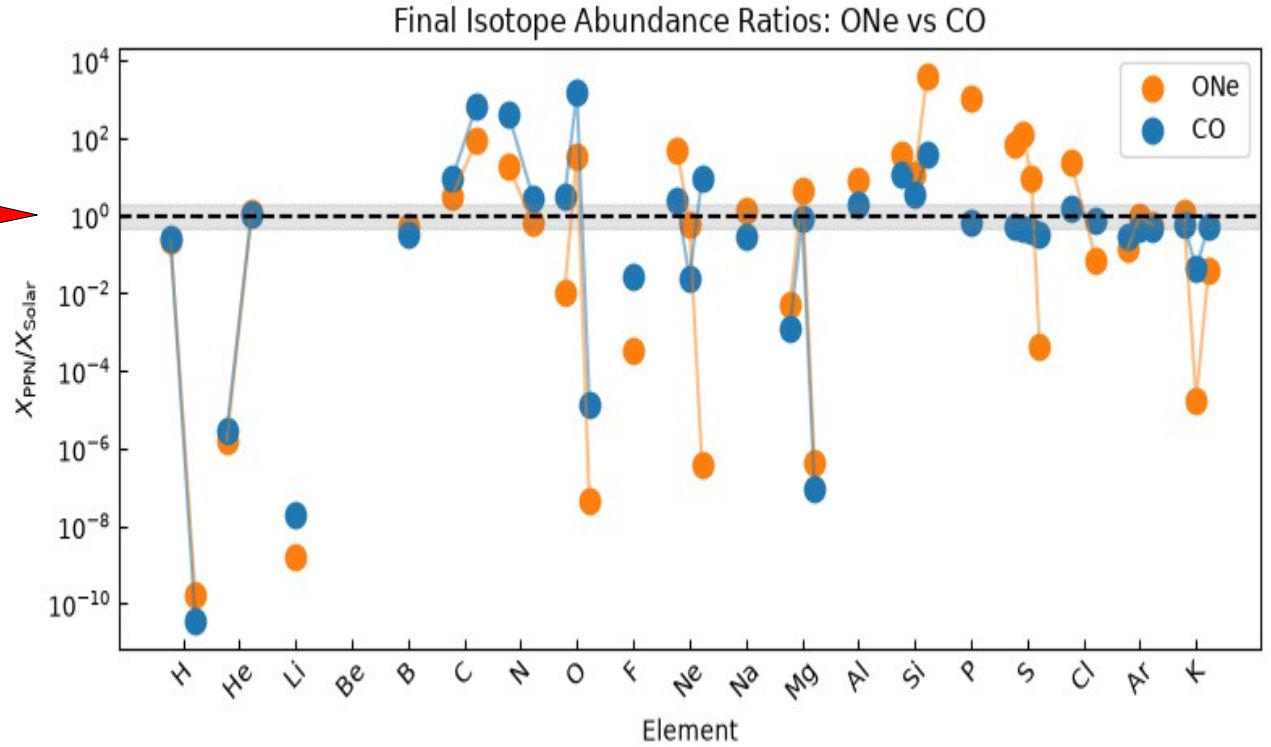
Solar Abundance Ratios

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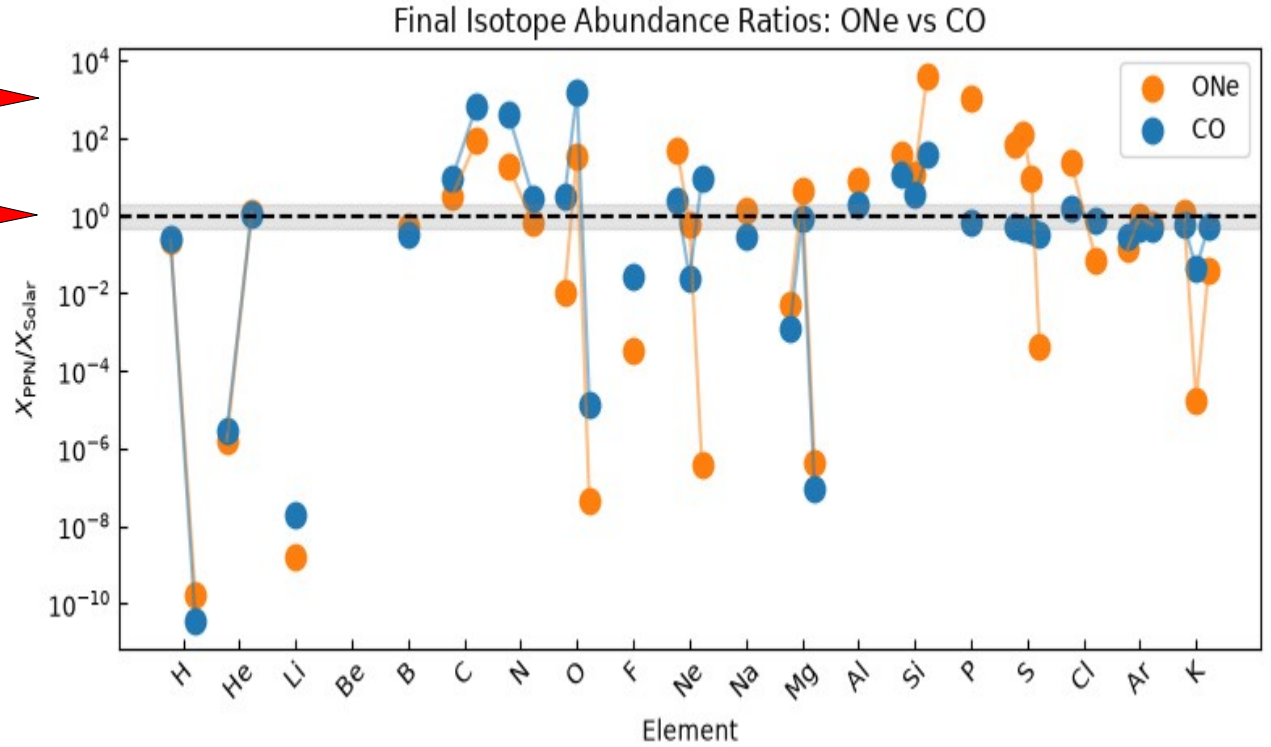
Solar abundance



Solar Abundance Ratios

Over-production 

Solar abundance 

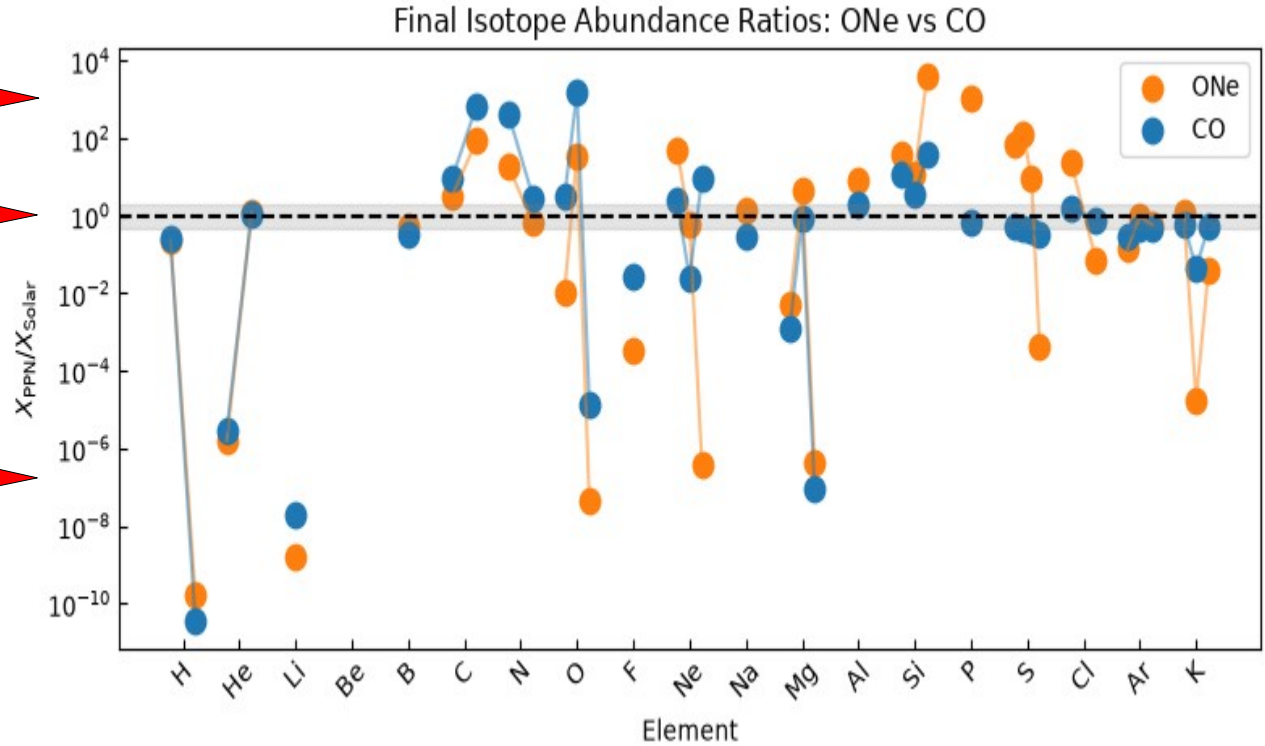


Solar Abundance Ratios

Over-production →

Solar abundance →

Under-production →

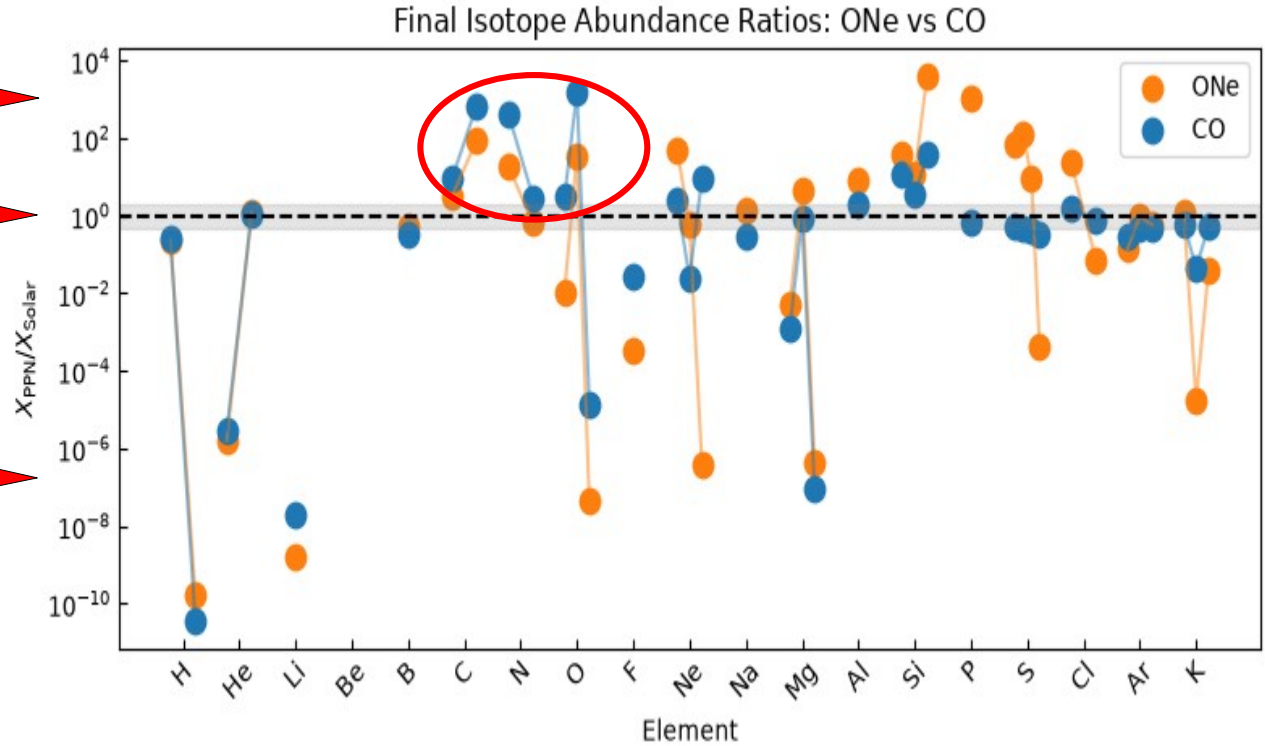


Solar Abundance Ratios

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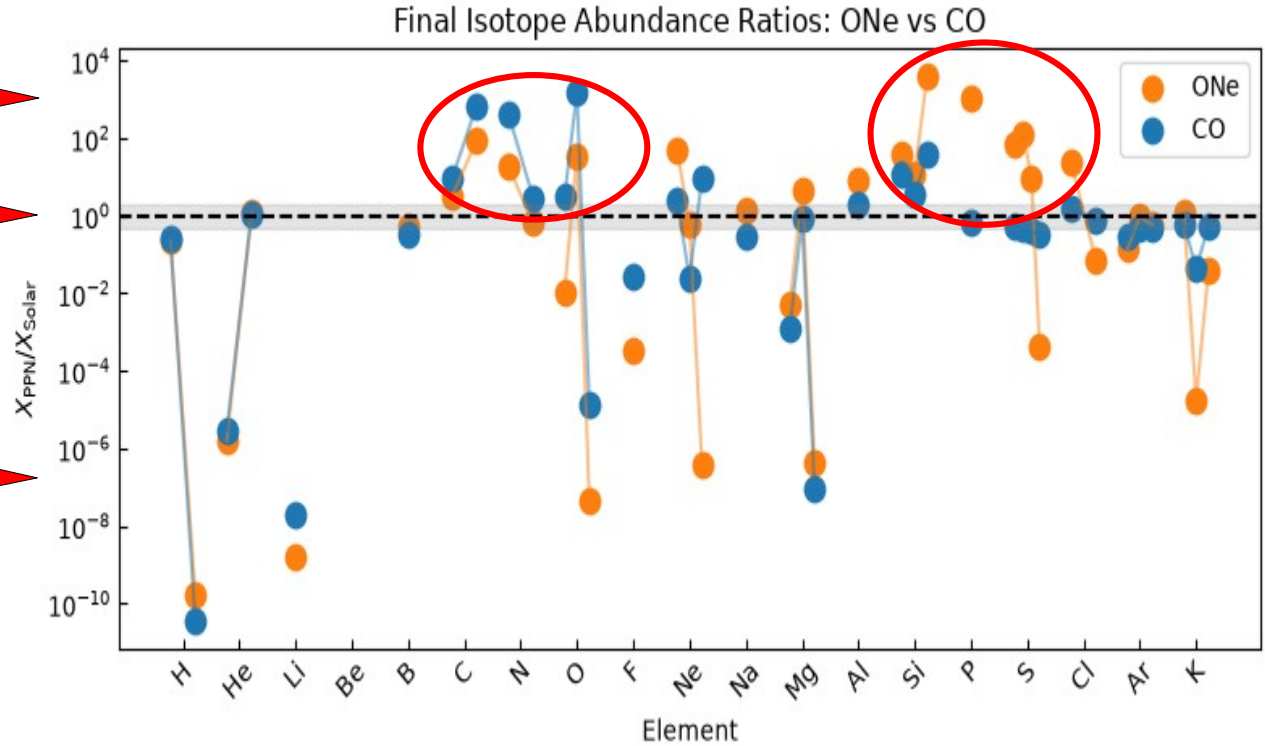
J. José & M. Hernanz, J. Phys. G: Nucl. Part. Phys. 34, R431 (2007)
S. Starrfield et al., Publ. Astron. Soc. Pac. 128, 051001 (2016)

Solar Abundance Ratios

Over-production →

Solar abundance →

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Experiments



Nuclear
Astrophysics

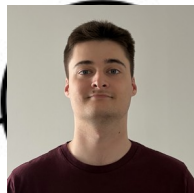
Simulations



Observations



Experiments

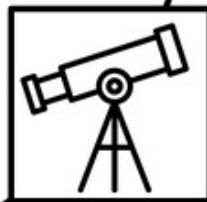


Nuclear
Astrophysics

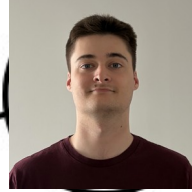
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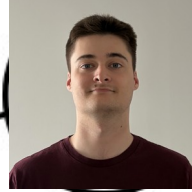
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Observations

Reaction Rate Uncertainties

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- Simulations of classical novae indicate the presence of rare radioactive isotopes

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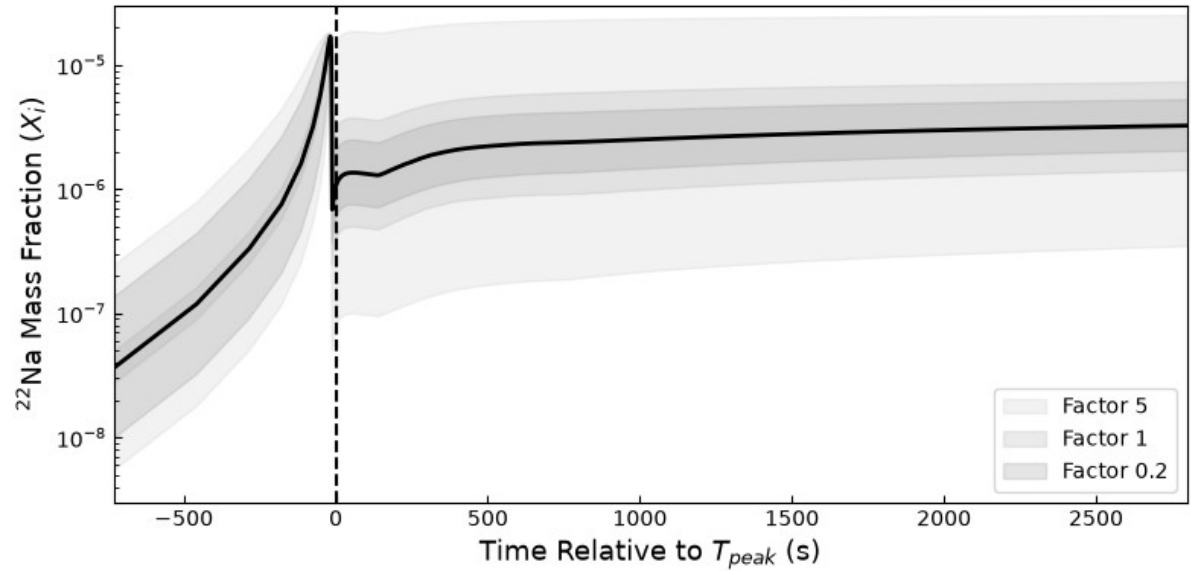
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Reaction Rate Uncertainties

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- Said reactions not well studied; have poorly constrained uncertainties

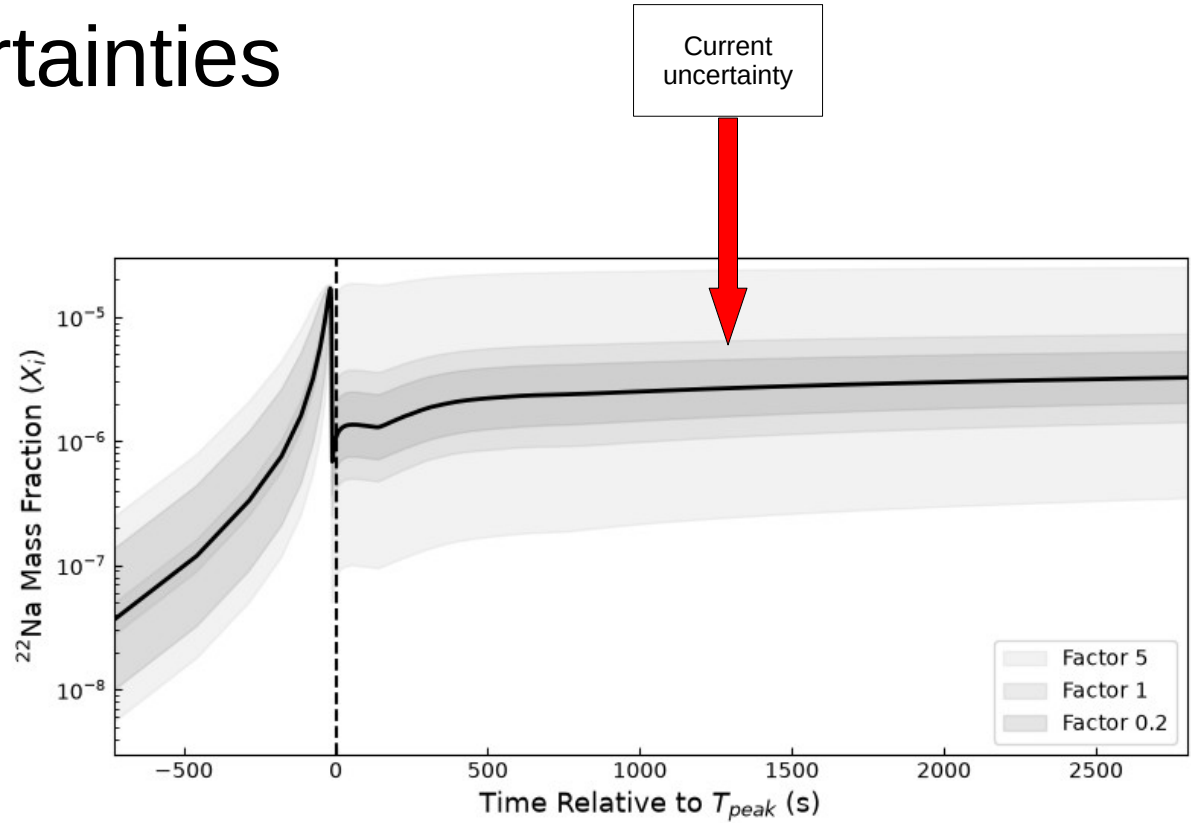
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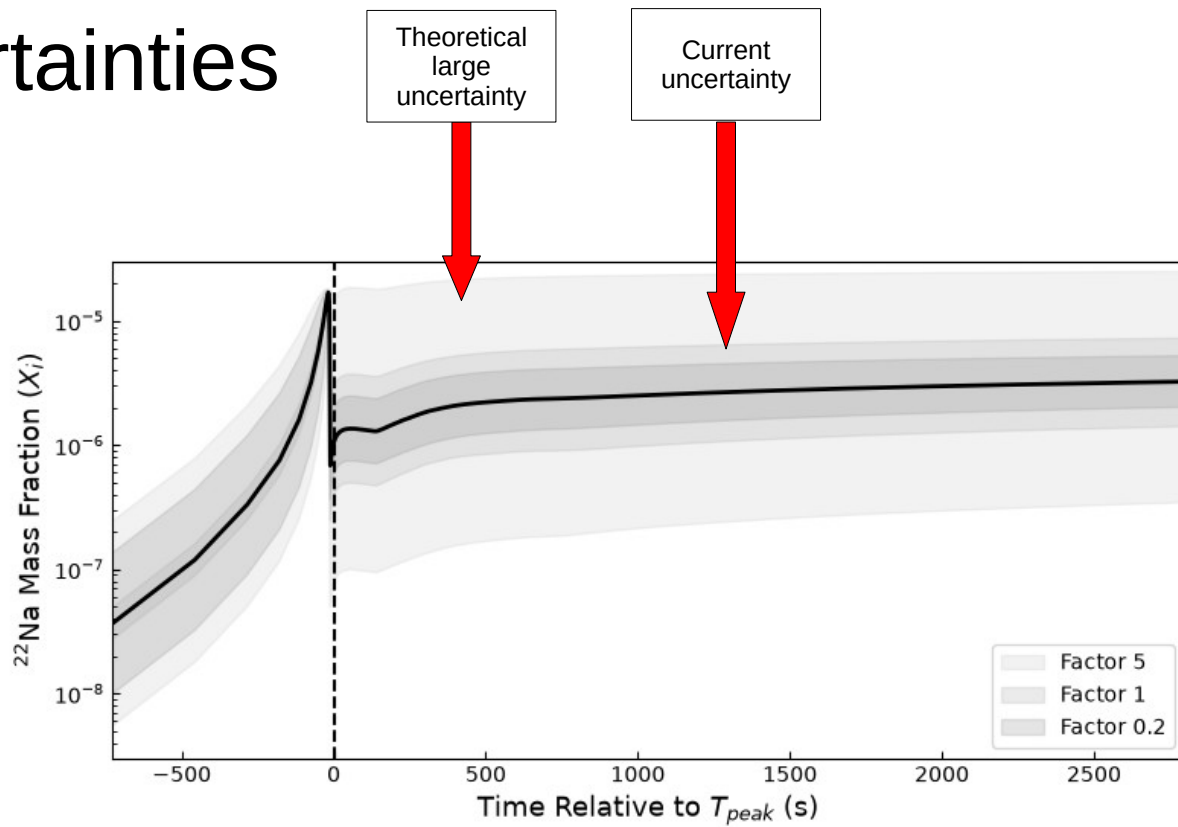
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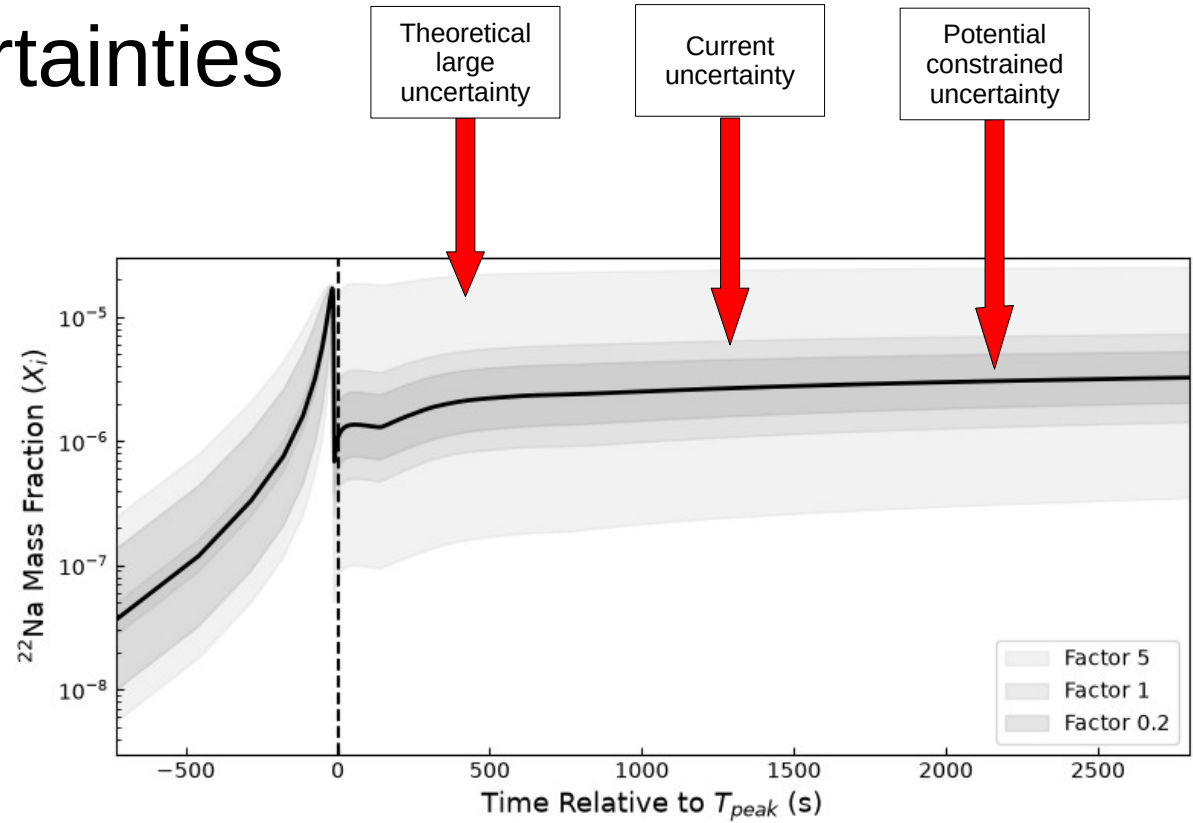
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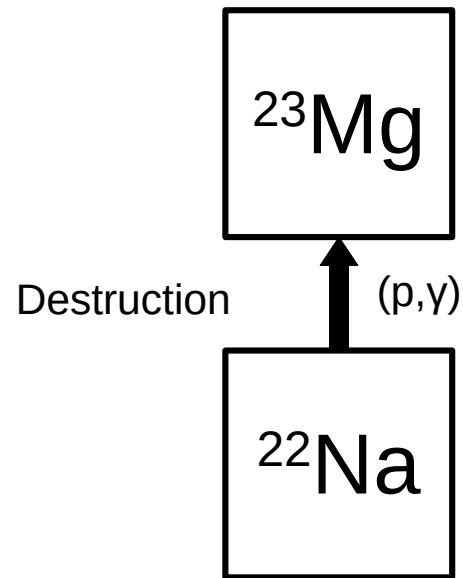
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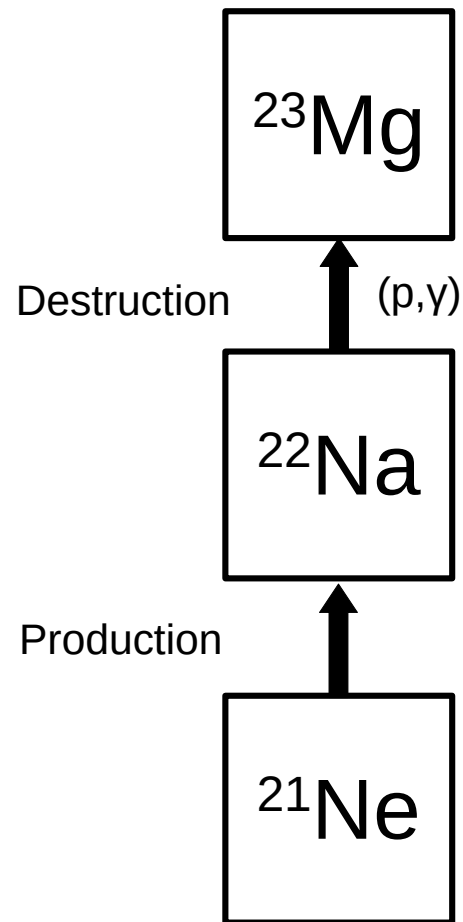
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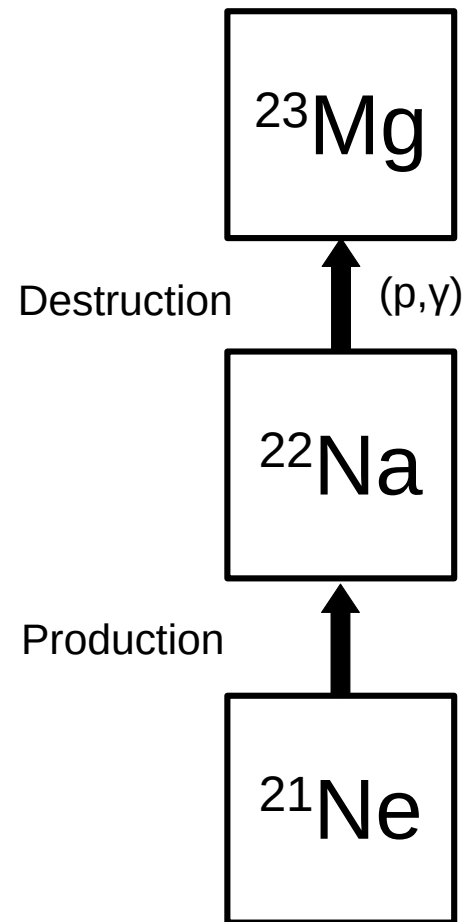
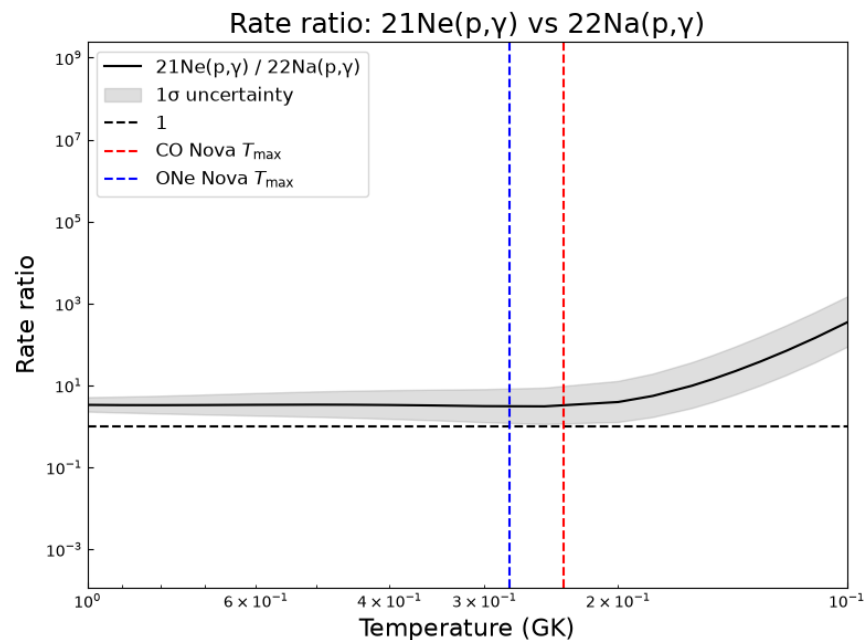
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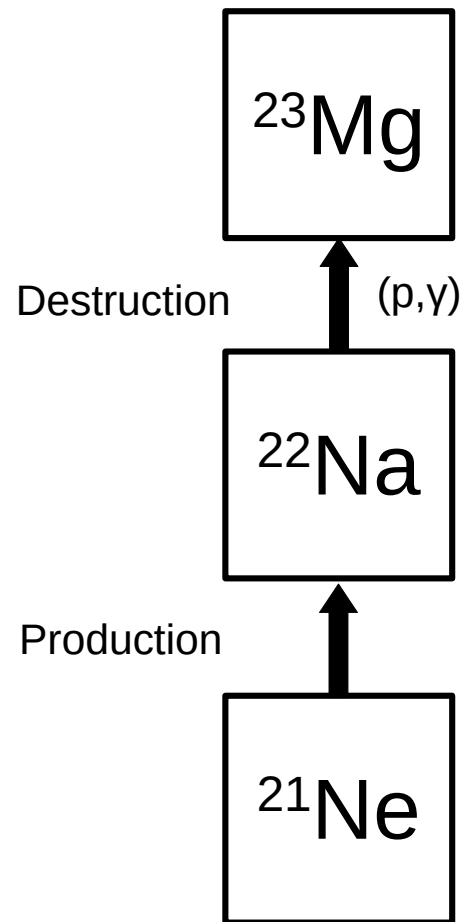
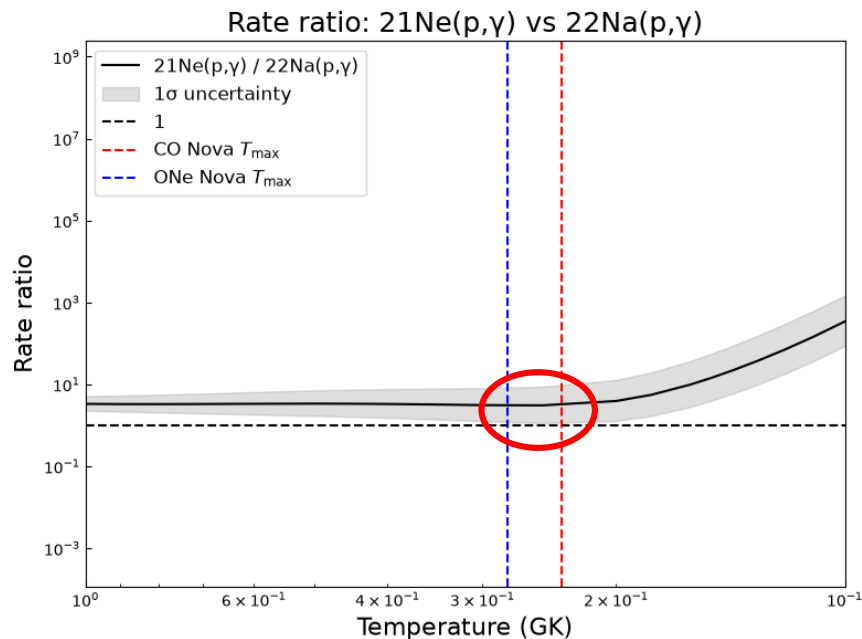
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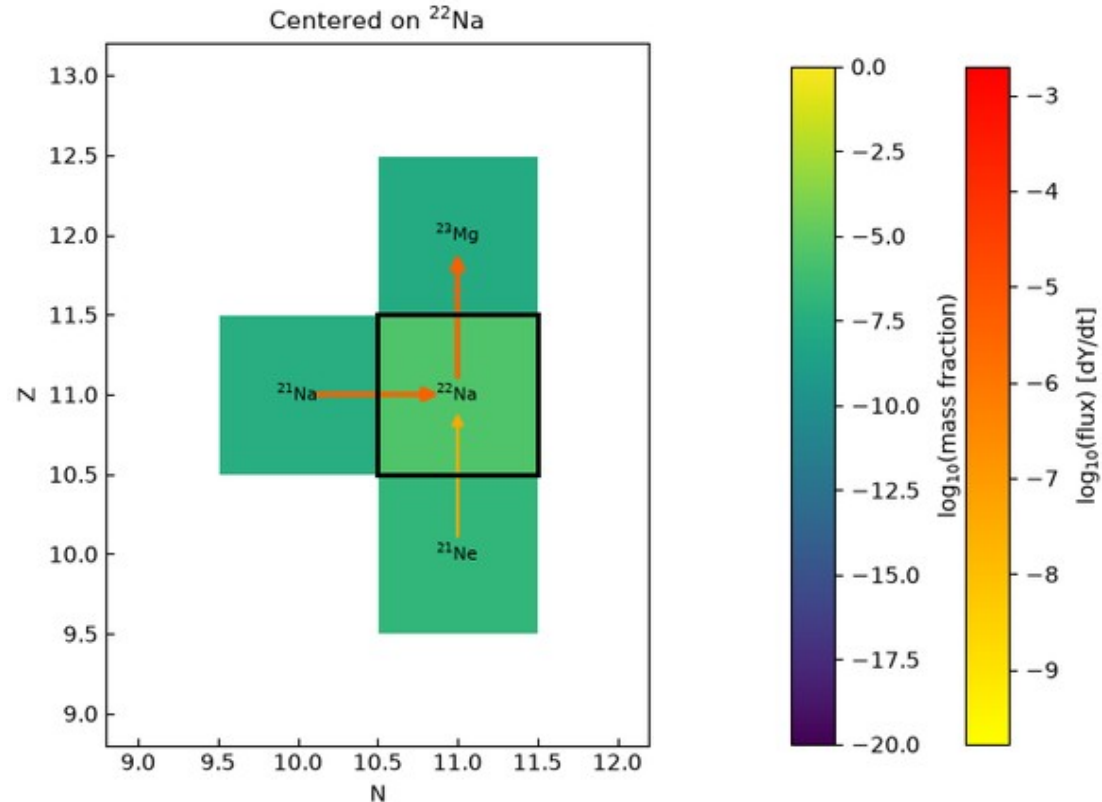
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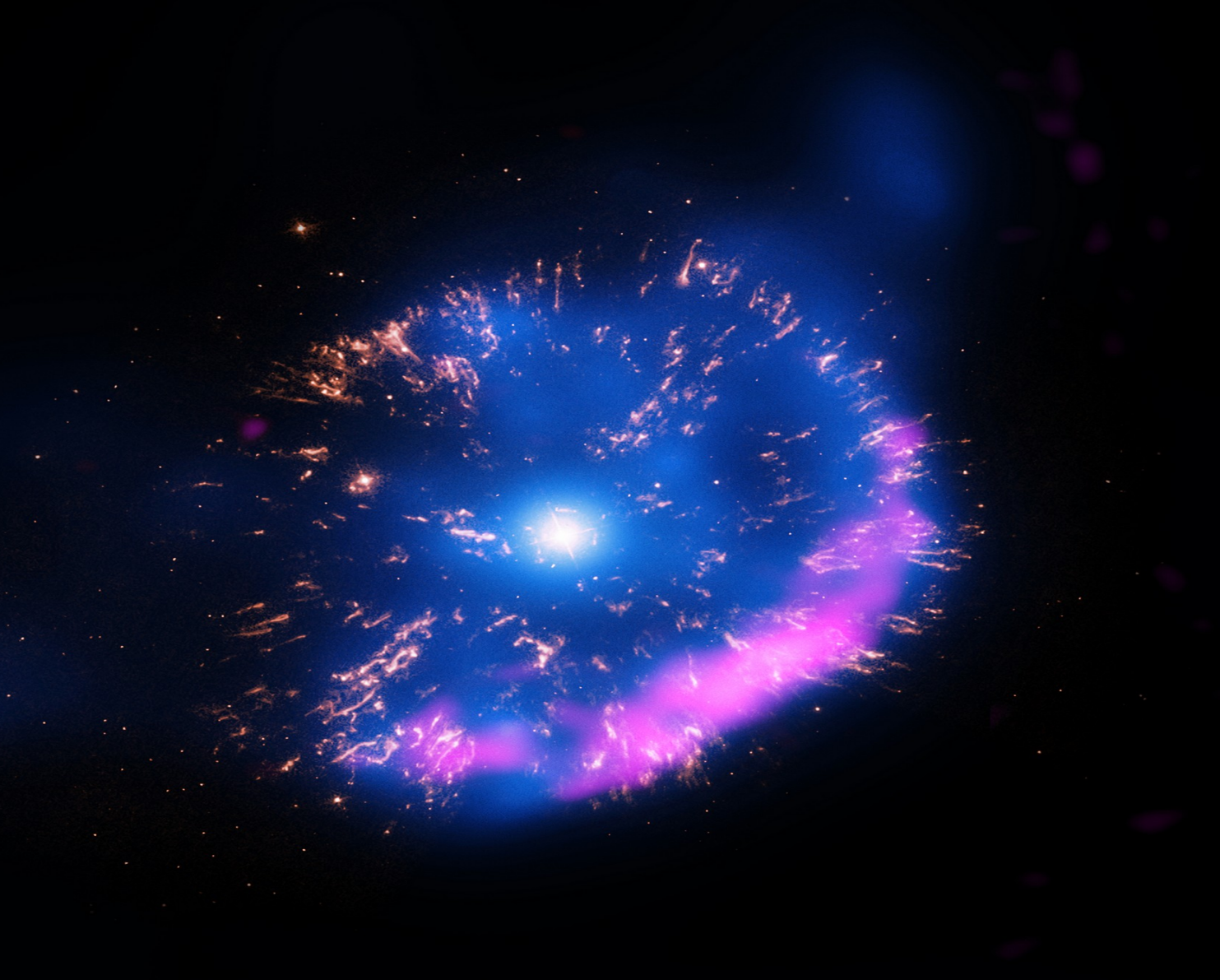
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Conclusions

- Simulations of classical novae indicate the presence of rare radioactive isotopes
- Nuclear reactions that produce these isotopes have poorly constrained uncertainties
- We must reproduce these reactions experimentally to constrain their uncertainties



Future Work



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- Apply the same methodology in other radioactive isotopes



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- Study the relevant nuclear reaction properties in the lab using radioactive ion beams (TRIUMF)



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- COSI gamma-ray telescope expected to launch in 2027



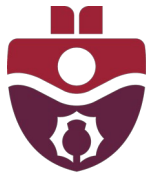
Acknowledgements



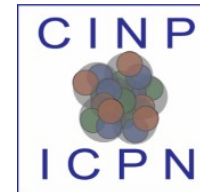
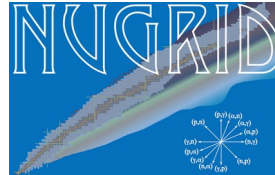
Thanassis Psaltis

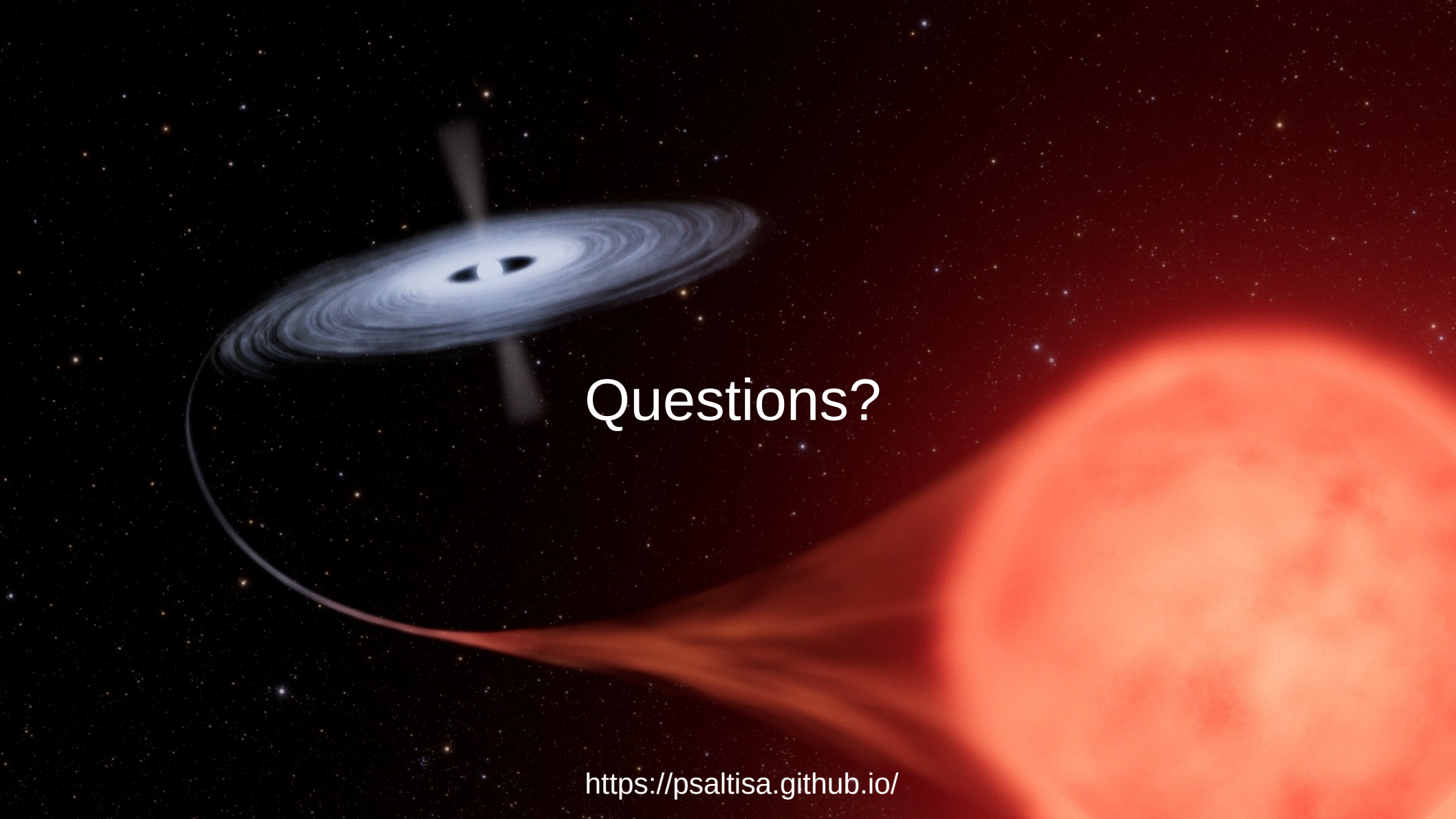


Shane Gervais



**Saint Mary's
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A cosmic scene featuring a spiral galaxy in the upper left, a bright red star in the lower right, and a comet-like streak of light across the center. The background is a dark space filled with numerous small stars.

Questions?

<https://psaltisa.github.io/>